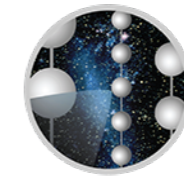




UPPSALA  
UNIVERSITET



**ICECUBE**  
SOUTH POLE NEUTRINO OBSERVATORY

# ONLINE ICRC 2021

THE ASTROPARTICLE PHYSICS CONFERENCE  
Berlin | Germany



37<sup>th</sup> International  
Cosmic Ray Conference  
12–23 July 2021

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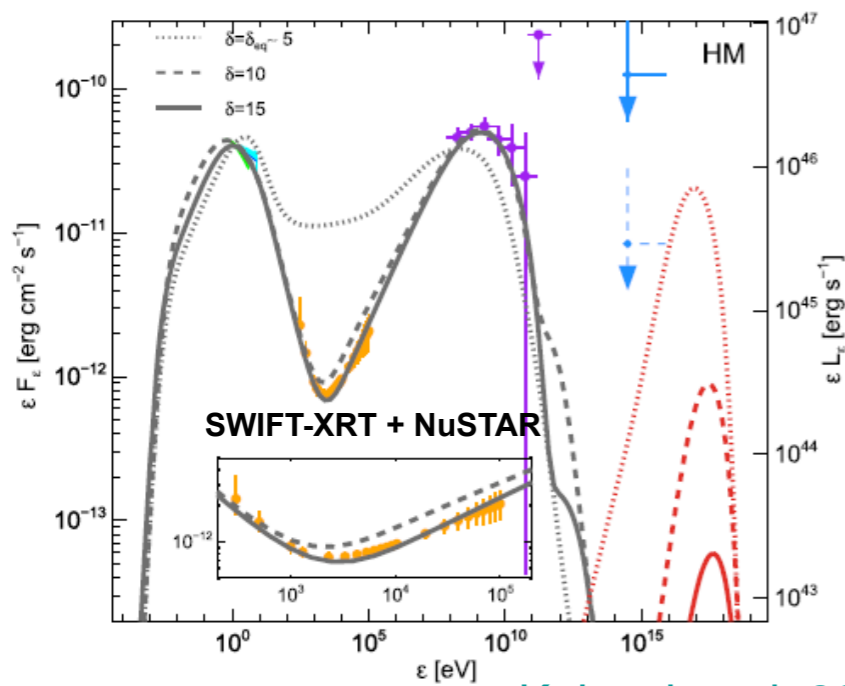
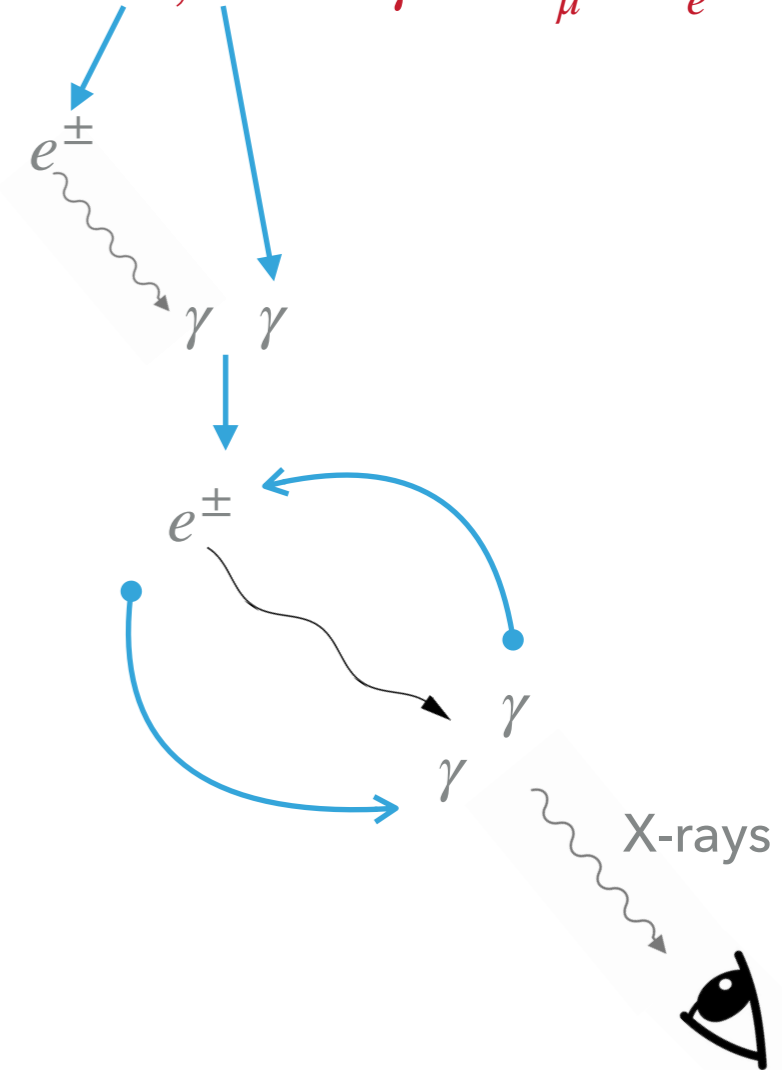
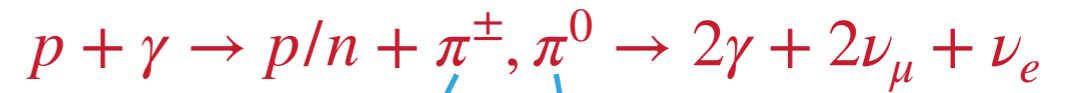
(FOR THE ICECUBE COLLABORATION)

## A MODEL-INDEPENDENT ANALYSIS OF NEUTRINO FLARES DETECTED IN ICECUBE FROM X-RAY SELECTED BLAZARS

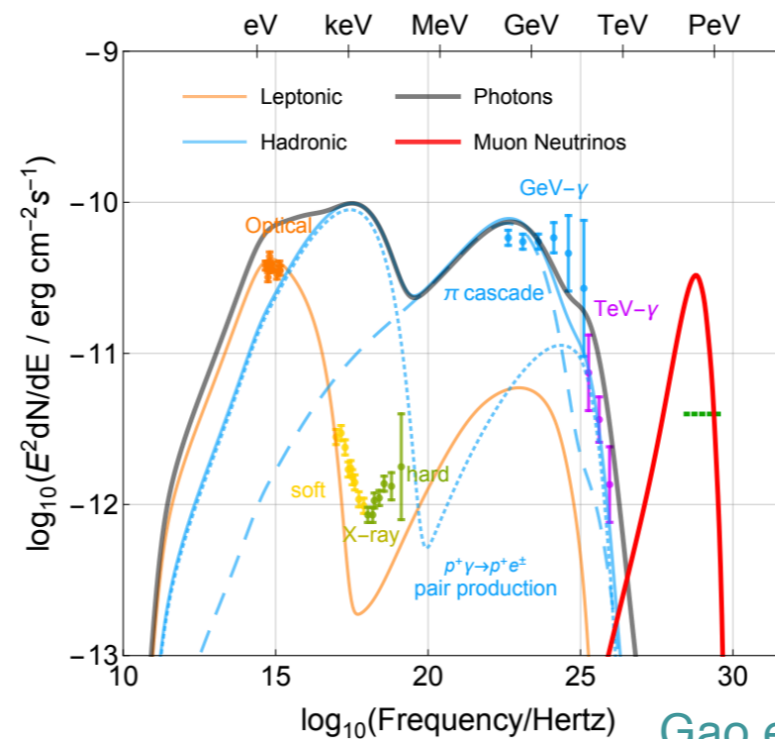
\*PRESENTER

## Proton Induced Cascades (PIC) :

- Secondary  $e^\pm$  pairs generated by pion decay produce high energy photons via Synch./Compton processes
- These photons combine to produce new pairs, which radiate a new generation of photons
- The *synchrotron-pair cascades* shift the extreme proton energies down to the X-ray band



Keivani et al. 2018



Gao et al. 2018

**X-rays from PIC can be used to constrain the expected neutrino flux from source, if a hadronic contribution is assumed**

**Objective:** Test the hypothesis ‘X-ray bright blazars can be sources of high-energy astrophysical neutrinos’....under the assumption that blazars can flare > twice on avg. over the period of observation

○ A **model independent**, time-dependent, untriggered multi-flare search:

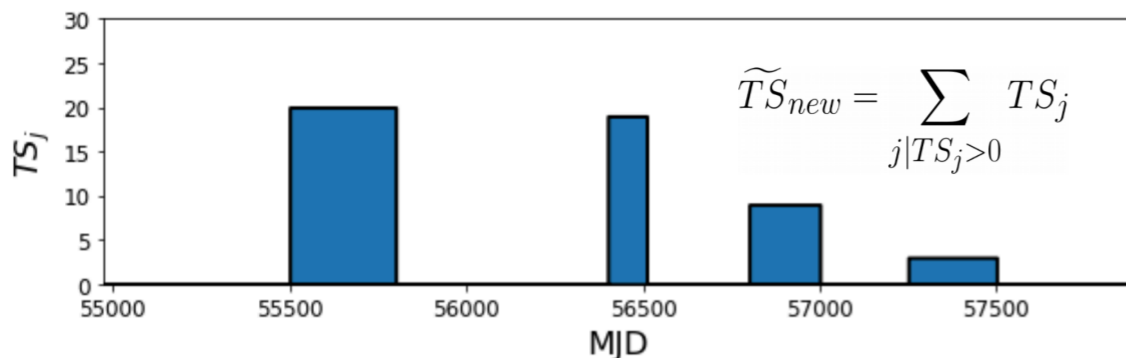
- ➔ search for neutrino flares in 10 years of IceCube data from a catalog of blazars curated by X-ray fluxes, and obtain a local p-value for each source
- ➔ perform a population test using the binomial test statistic to determine the sub-population with statistically significant emission

## Multi-flare method

Reports the combined significance of all flares from a source direction (multi-flare TS) instead of the most significant flare

$$\mathcal{L}(n_s, \gamma, \Delta t_j) = \prod_{i=1}^N \frac{n_s}{N} S_i + (1 - \frac{n_s}{N}) B_i$$

$$TS_j |_{\Delta t_j} = -2 \log \left[ \frac{\Delta T_{\text{data}}}{\Delta t_j} \times \frac{\mathcal{L}(\vec{x}_s, n_s = 0)}{\mathcal{L}(\vec{x}_s, \hat{n}_s, \hat{\gamma}_s)} \right]$$

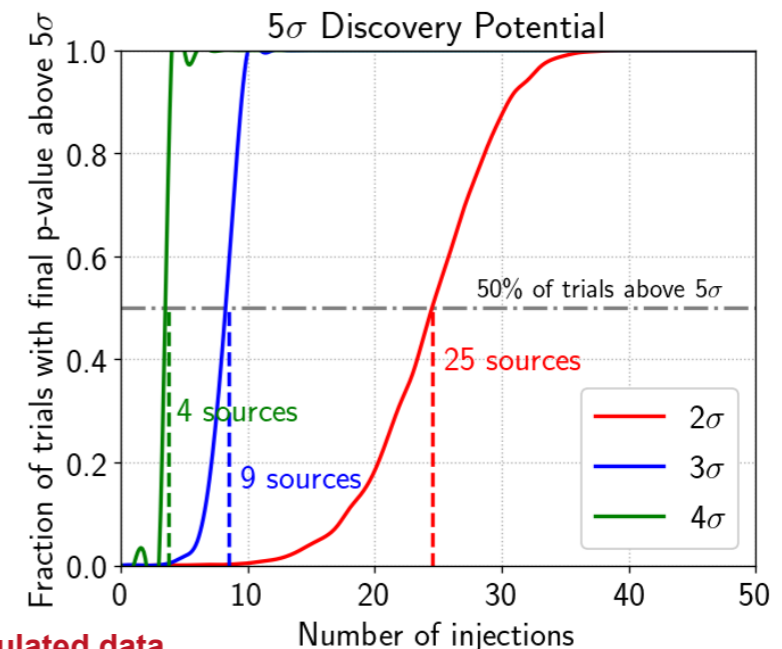


Credit: Will Luszczyk

## Binomial Test

If a sub-population within the catalog has statistically significant emission, the test can reveal how many and which sources are of interest

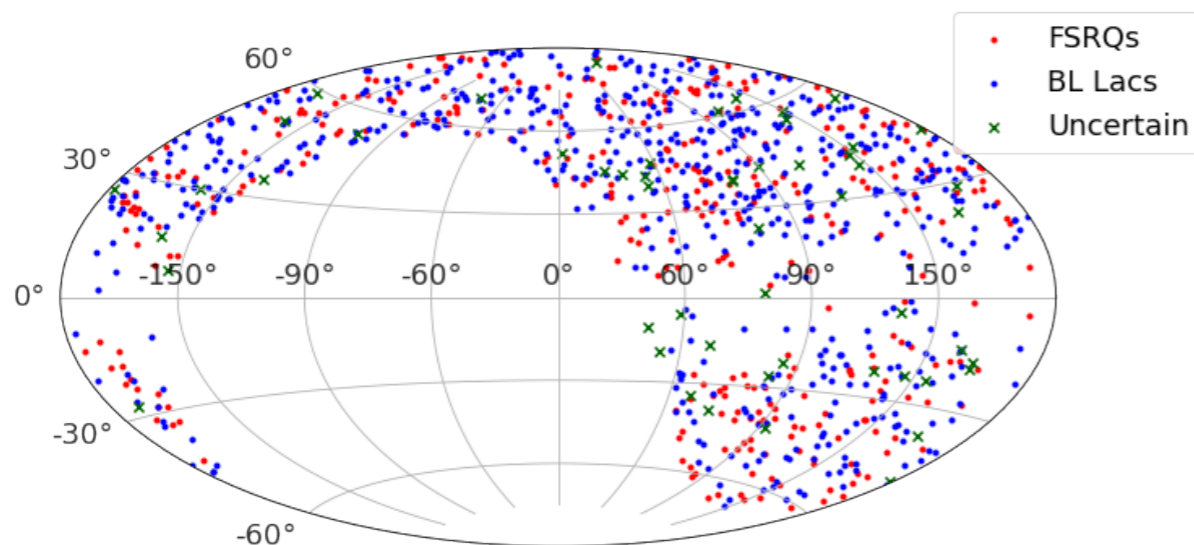
$$P(k) = \sum_{m=k}^N \binom{N}{m} p_k^m (1 - p_k)^{N-m}$$



\*Plot based on simulated data

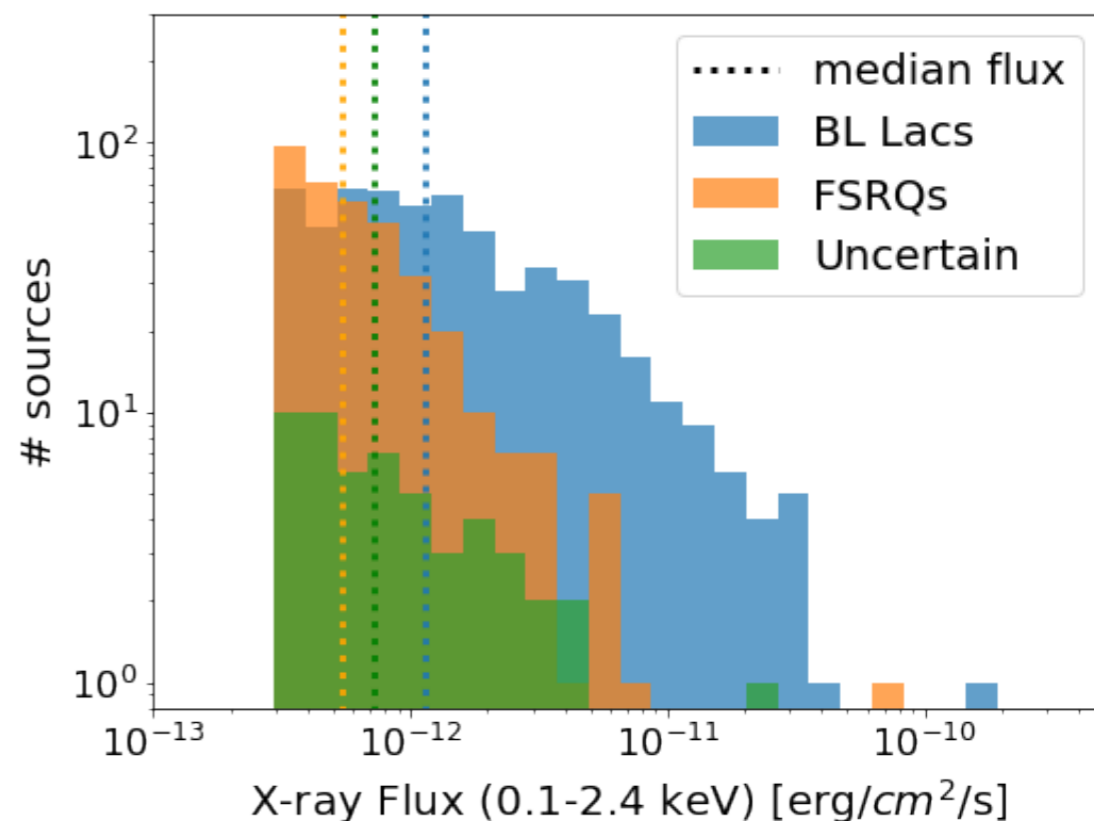
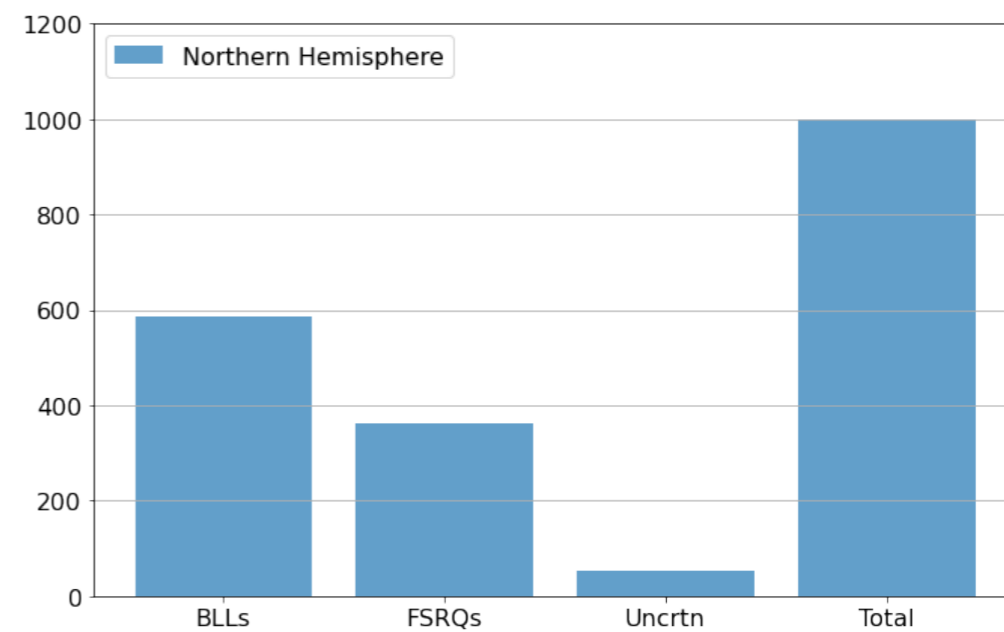
## ROMABZCAT 5TH EDITION (2015)

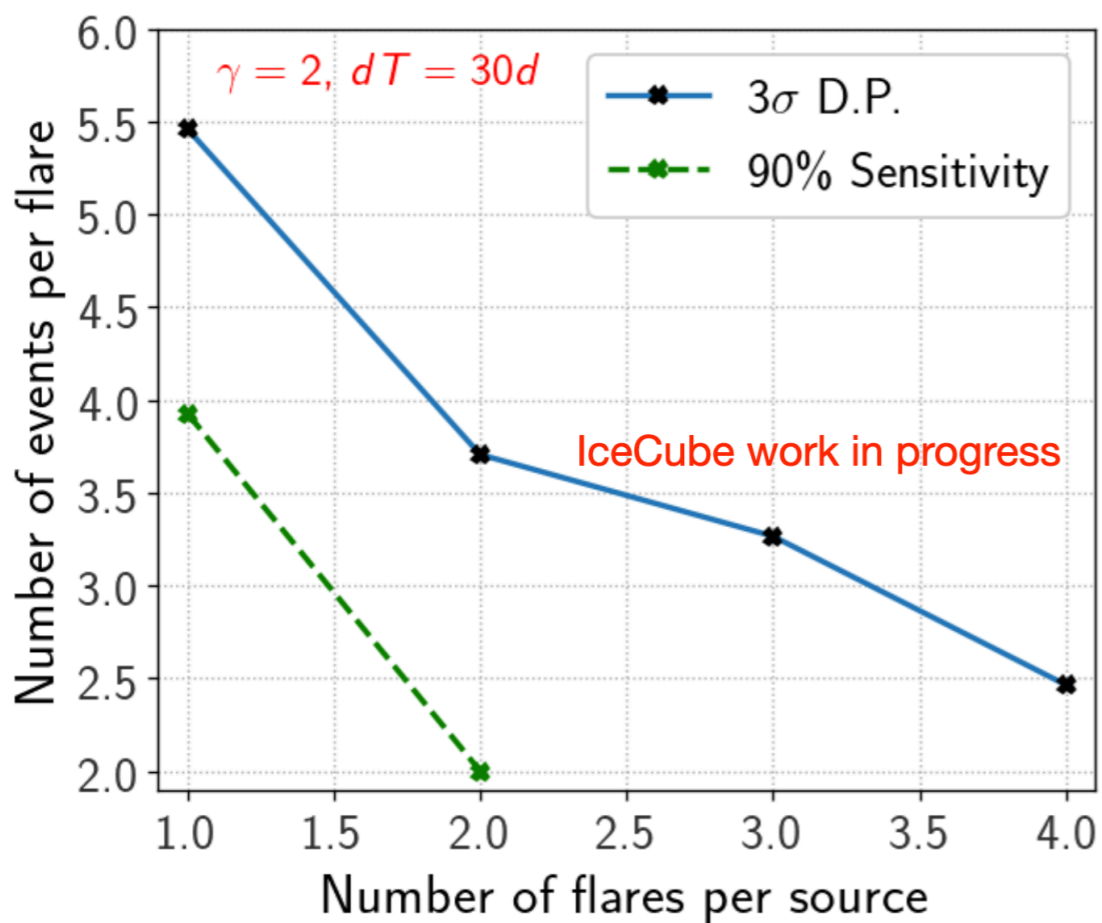
- Multi-frequency blazar catalog; X-ray fluxes (**0.1 - 2.4 keV**) taken from **ROSAT All Sky Survey**
- ✓ We select 1000 blazars from the Northern hemisphere (-5, +85 deg.) with the highest X-ray fluxes in the catalog



After selection cuts:

- BL Lacs: **586**, FSRQs: **361**, Uncertain type: **53**
- X-ray fluxes b/w  **$3.1e-13$**  to  **$1.8e-10$  erg/cm<sup>2</sup>/s**

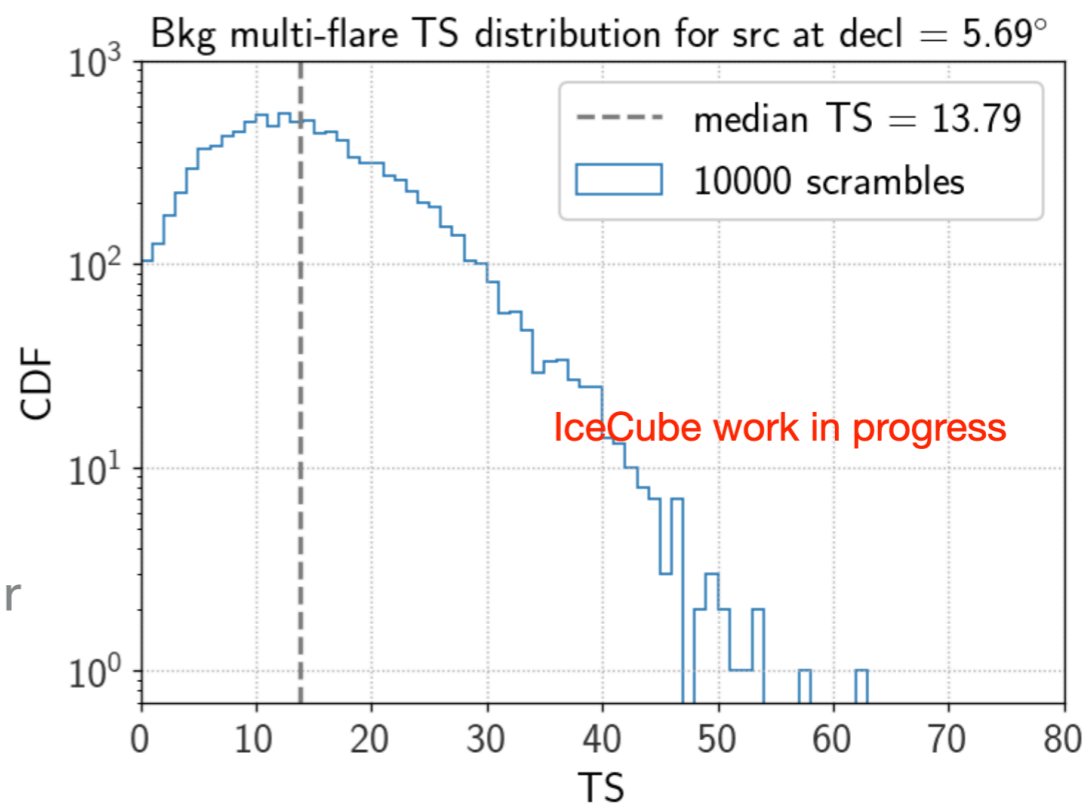




- 10000 trials for a single source location (at the declination of TXS 0506+056) with and w/o signal injection
- Fixed spectral index and flare duration for injected signal ( $\gamma = 2, \delta T = 30 \text{ days}$ )
- Box profile flares; S/B = 2000

Strength of individual flares from a single source required to obtain a final discovery potential (D.P.) of 3-sigma and a 90% C.L. sensitivity

Background multi-flare TS distribution for the declination of TXS across 10k trials



**Full catalog sensitivity and final p-values for each source will be published soon!!**