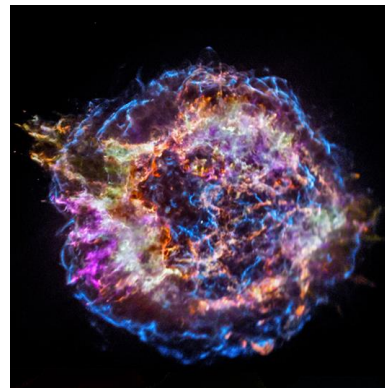


Electron-beam instabilities in the foreshock of oblique high-Mach-number shocks

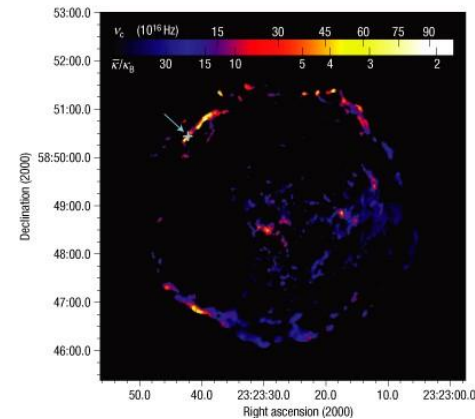
Martin S. Weidl (IPP Garching)

Artem Bohdan, Paul J. Morris, Martin Pohl (DESY Zeuthen)

International Cosmic-Ray Conference 2021



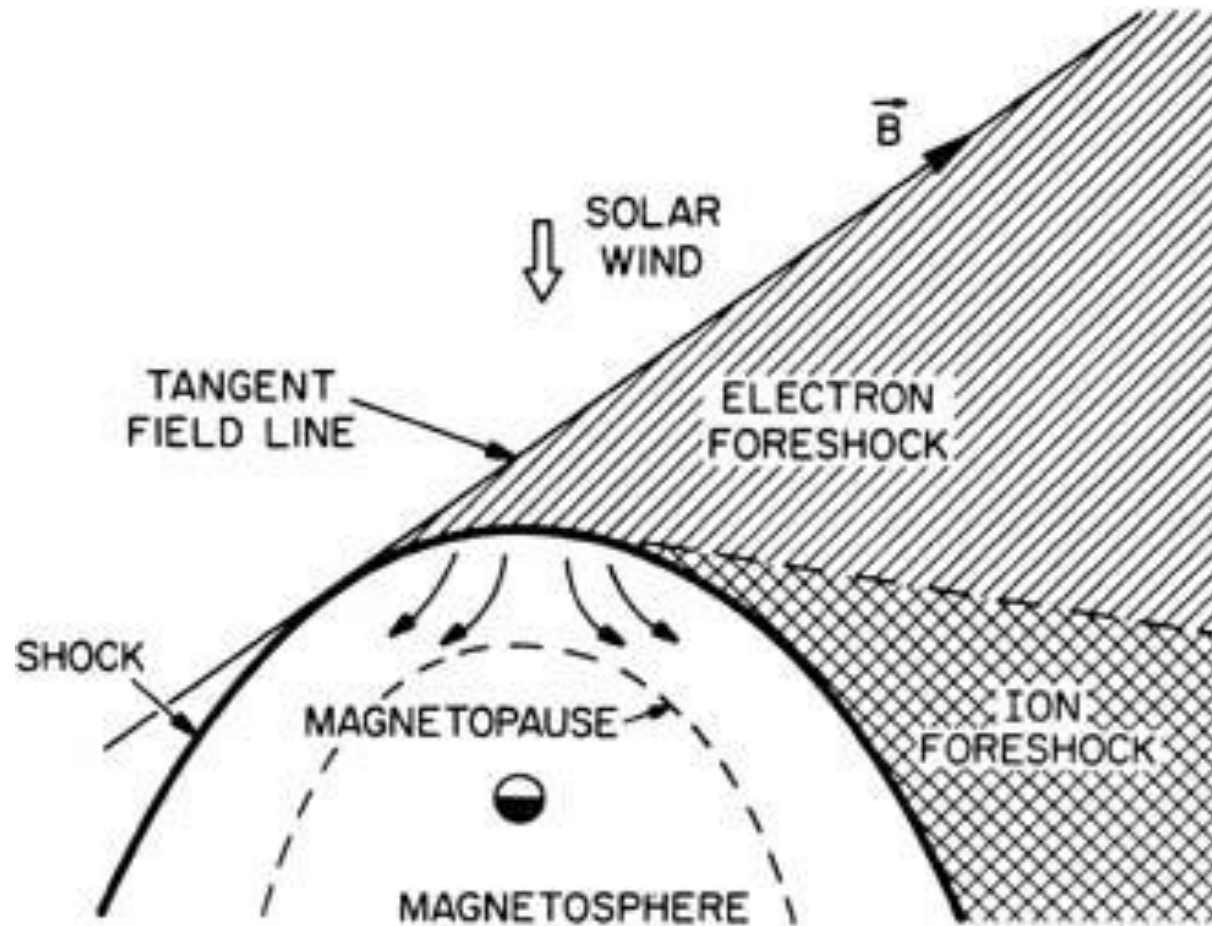
[NASA/CXC/SAO]



[Stage et al 2006]



Foreshock geometry at oblique shocks



[D. A. Gurnett]



Typical electron-foreshock waves

Alfvén waves

Electron-beam firehose instability

[Gary 85]

Langmuir waves

2-stream instability

Quasiparallel whistler waves

Whistler heat-flux instability

[Gary 85]

Oblique whistler waves

[Sentman et al 83]

Left-hand polarised

Electrostatic

Right-hand polarised



Periodic PIC simulations with solar-wind parameters

Borda de Agua et al, JGR 1996

2D3V simulation of electron beam with $T_{\perp} \gg T_{\parallel}$
Parallel Langmuir waves
Oblique whistler waves forward and backward

Fu et al, PoP 2014

2D3V simulation of electron beam with $T_{\perp} > T_{\parallel}$
Electrostatic instability, then **quasiparallel whistler** instability
Only weak oblique whistlers

Micera et al, ApJ 2020

2D3V simulation of electron strahl with $T_{\parallel} > T_{\perp}$
Fast-growing **oblique whistlers** scatter strahl perpendicularly
Then susceptible to **quasiparallel whistlers**

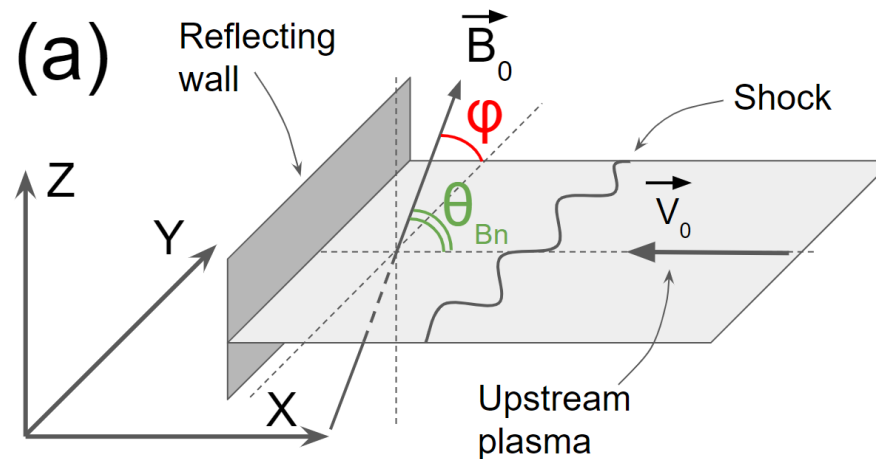
Shock simulation with Mach number 63

Xu et al, ApJ 2020

1D3V simulation of oblique shock with shock-mirrored electrons
Strong **LCP waves** in foreshock
Electrons undergo SDA in shock, then escape and **perform DSA**

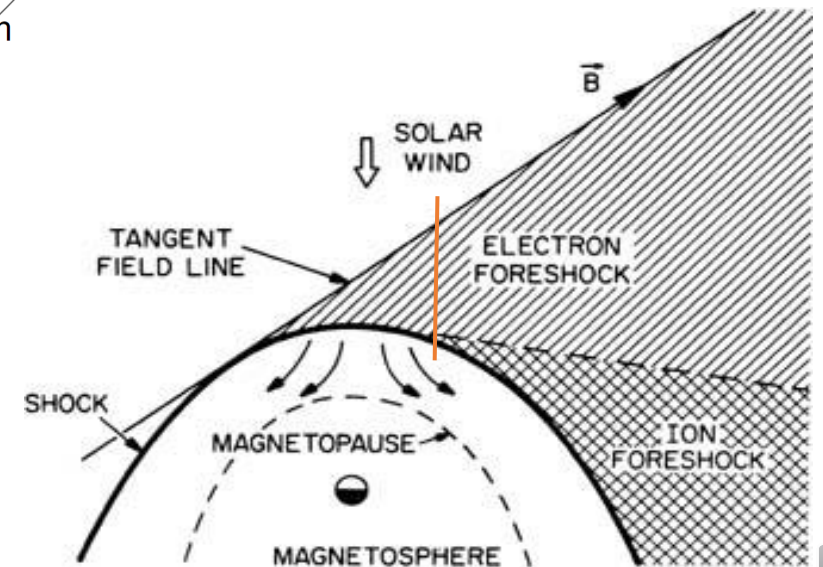


Shock-reflected electrons and the foreshock instabilities they excite

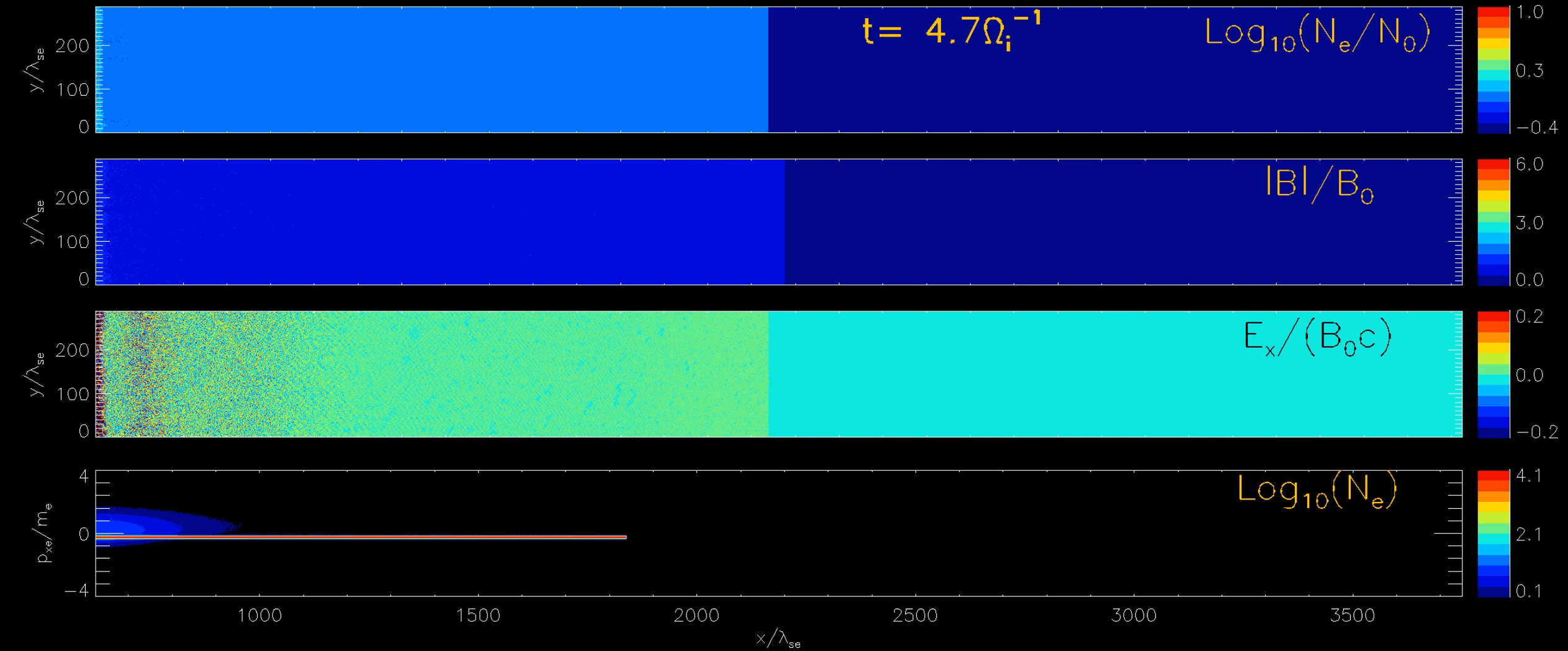


(a) Retrieve properties of shock-reflected electrons from 2D3V shock simulation

(b) Determine unstable wave modes from linear analysis and periodic PIC simulations



$$M_A = 30; \theta_{Bn} = 60^\circ$$

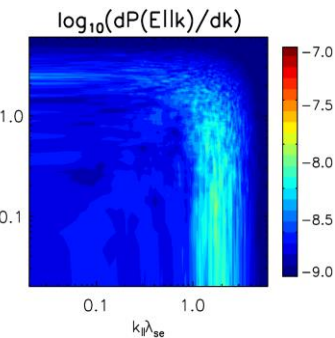
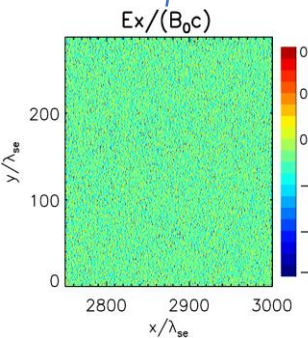
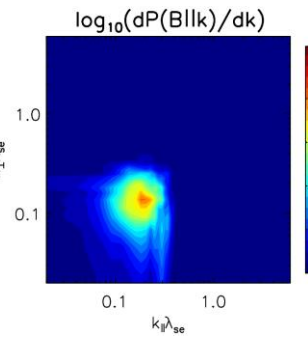
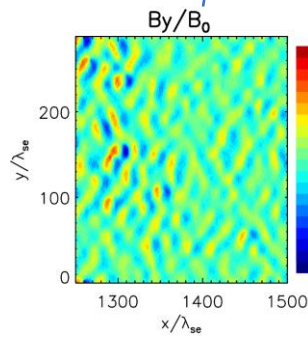
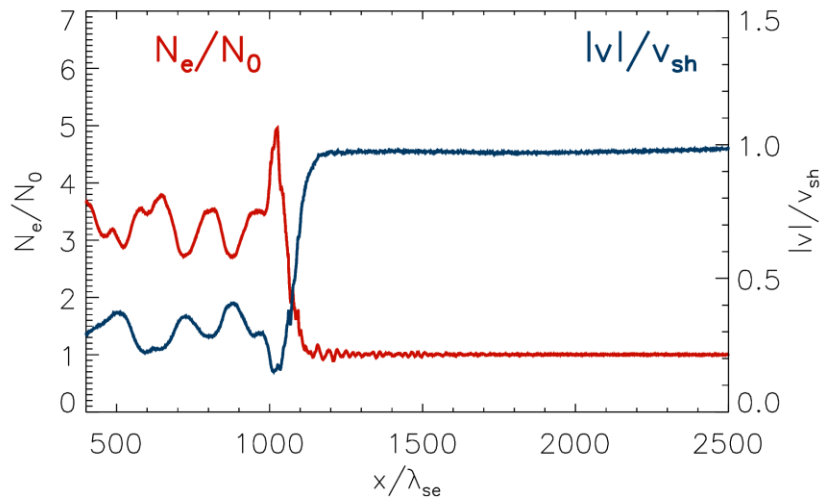
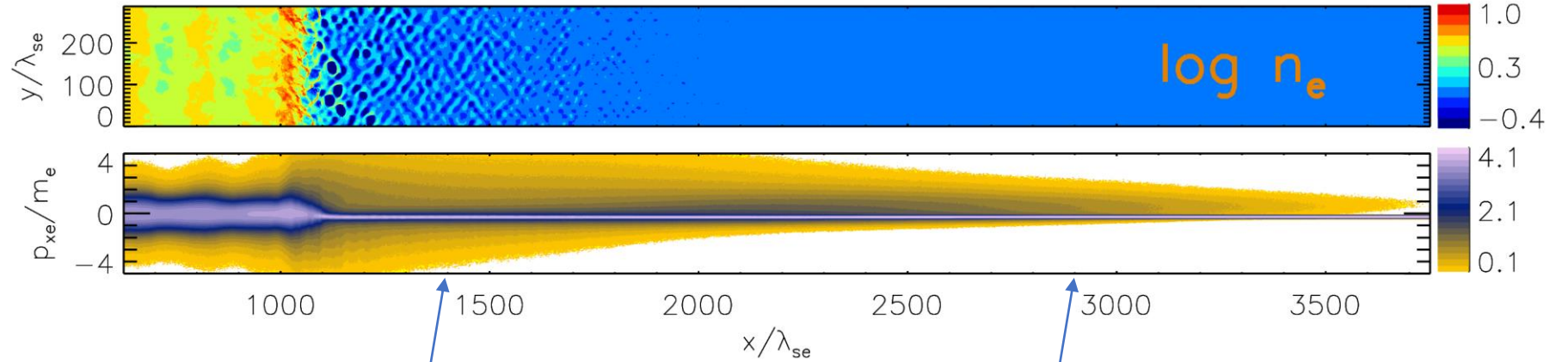


Movie – see video



Near vs far upstream in the electron foreshock

$M_A = 30$
 $\theta_{Bn} = 60^\circ$
 $m_p = 50$



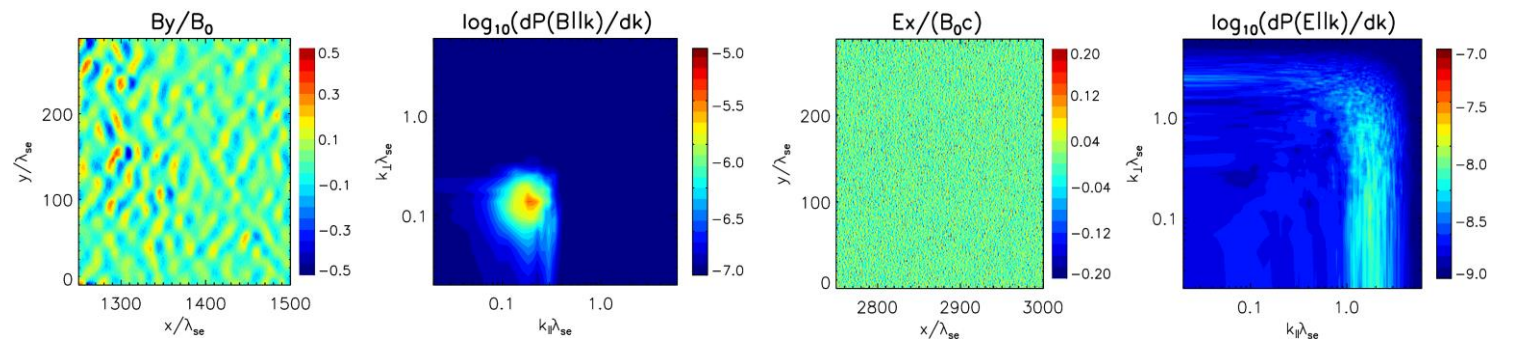
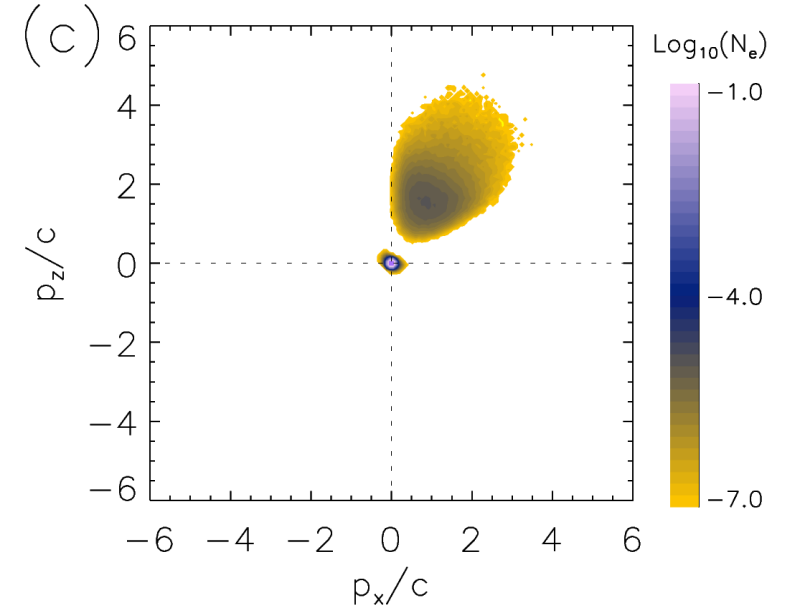
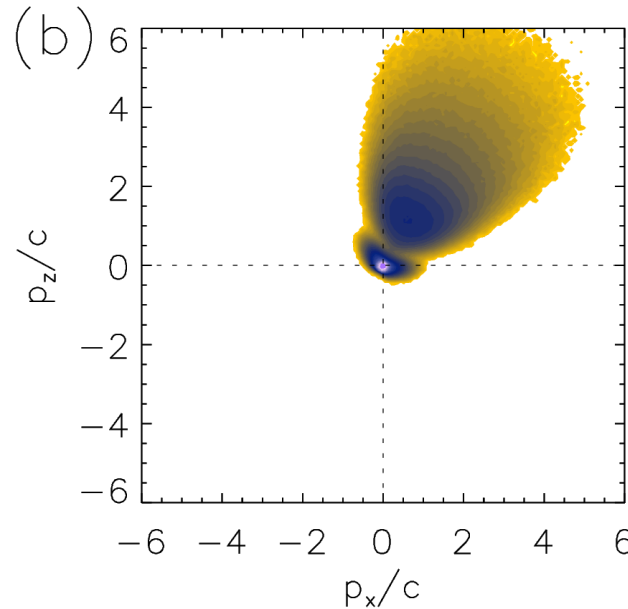
Near upstream:
Oblique whistlers

Far upstream:
Parallel electrostatic waves



Electron distributions

	Near upstrm	Far upstrm
n_{beam}	$0.025 n_0$	$0.0027 n_0$
β_{beam}	0.89	0.91
γ_{beam}	2.2	2.4
v_{th}	$0.46c$	$0.40c$
$(v_{th,\perp}/v_{th,\parallel})$	0.8	0.8



Near upstream:
Oblique whistlers

Far upstream:
Parallel electrostatic waves

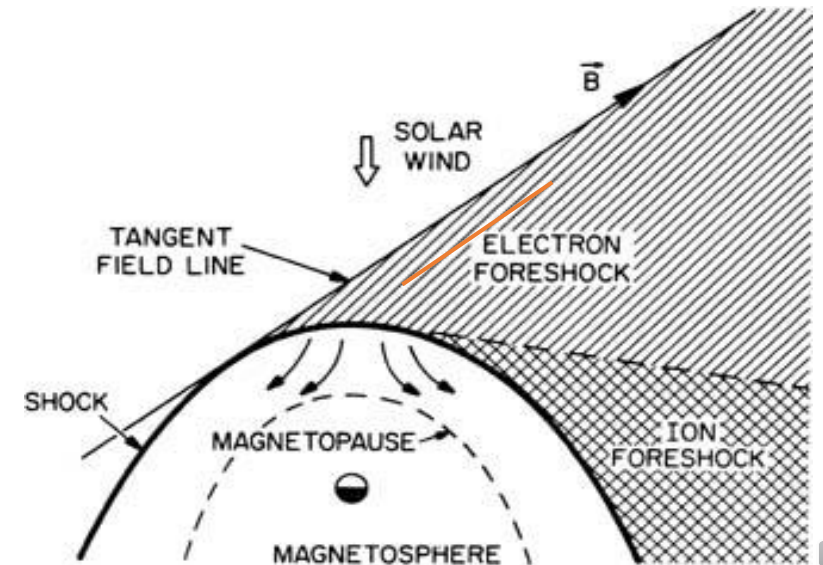
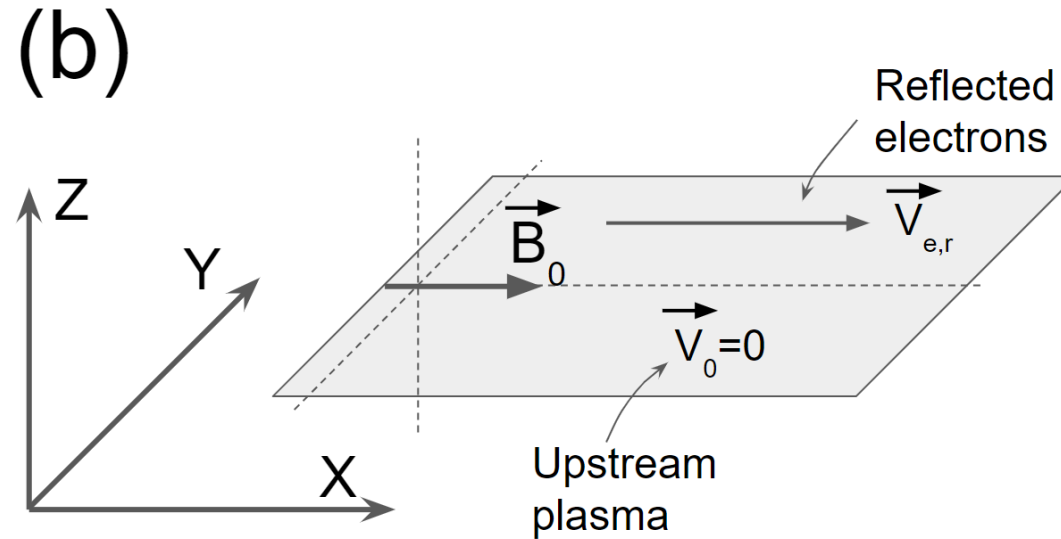
See Paul Morris' talk...



Reflection by Buneman-driven waves
at the shock creates relativistic
foreshock electrons

Periodic simulations: far upstream

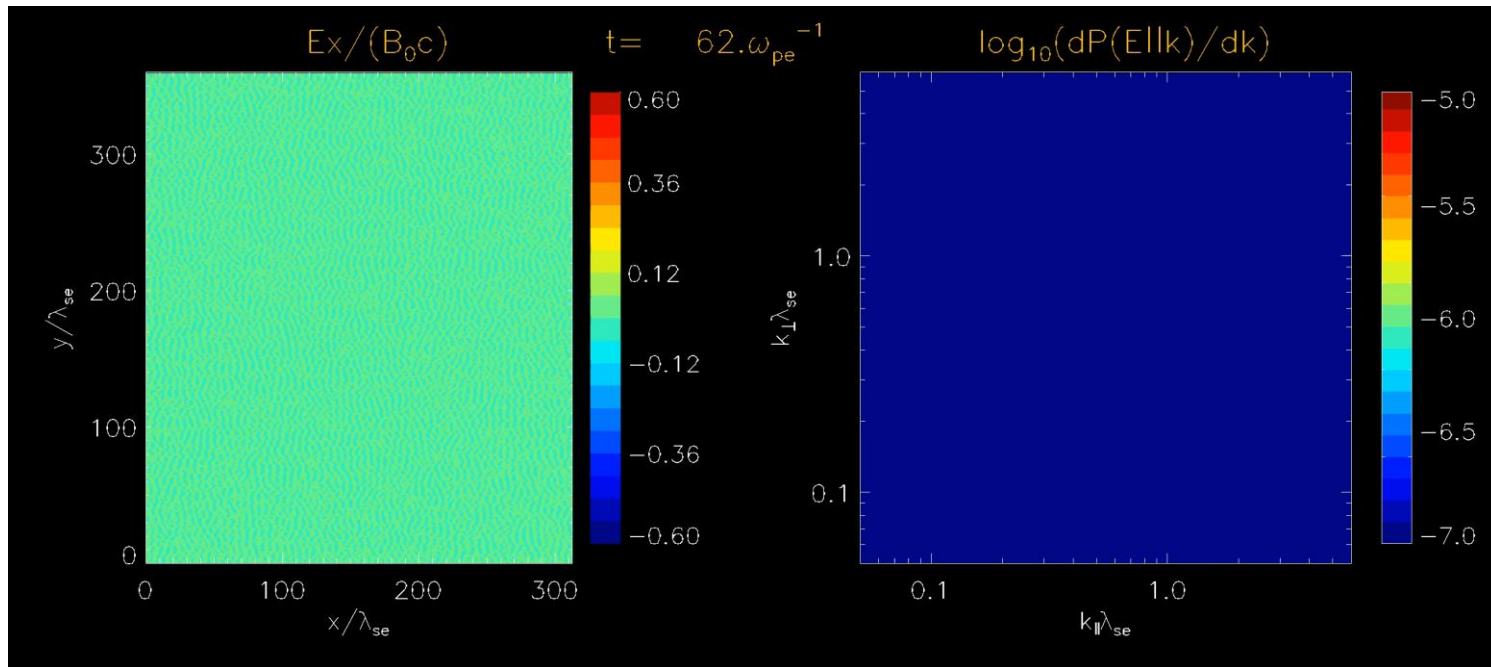
	Far upstrm
n_{beam}	$0.0027 n_0$
β_{beam}	0.91
γ_{beam}	2.4
v_{th}	$0.40c$
$(v_{th,\perp}/v_{th,\parallel})$	0.8



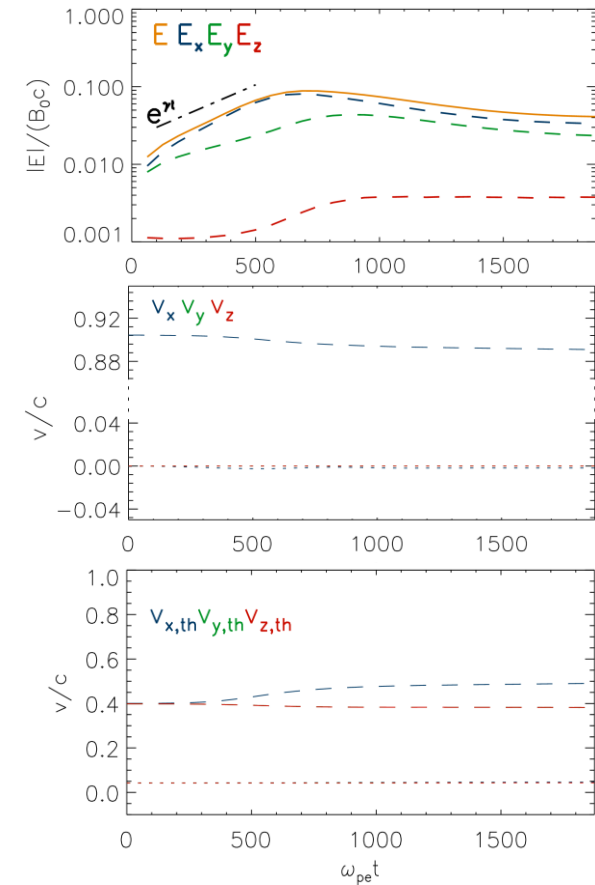
(b) Determine unstable wave modes from linear analysis and periodic PIC simulations



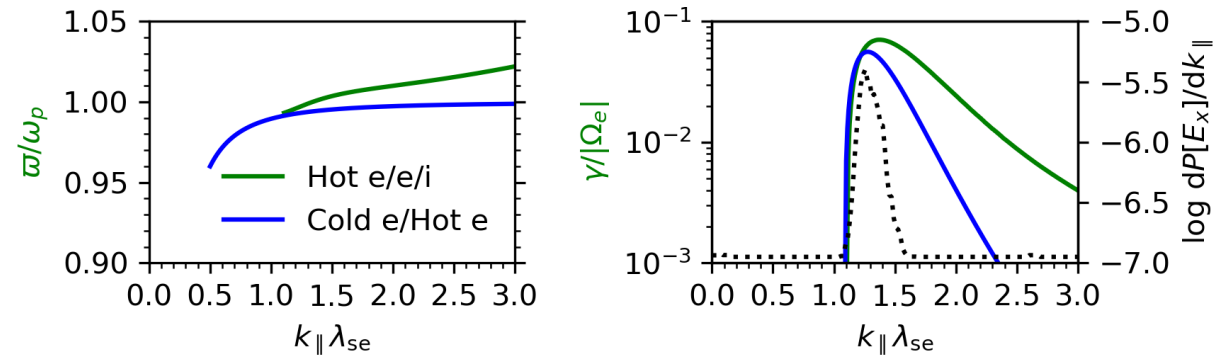
Far upstream: electron-acoustic instability



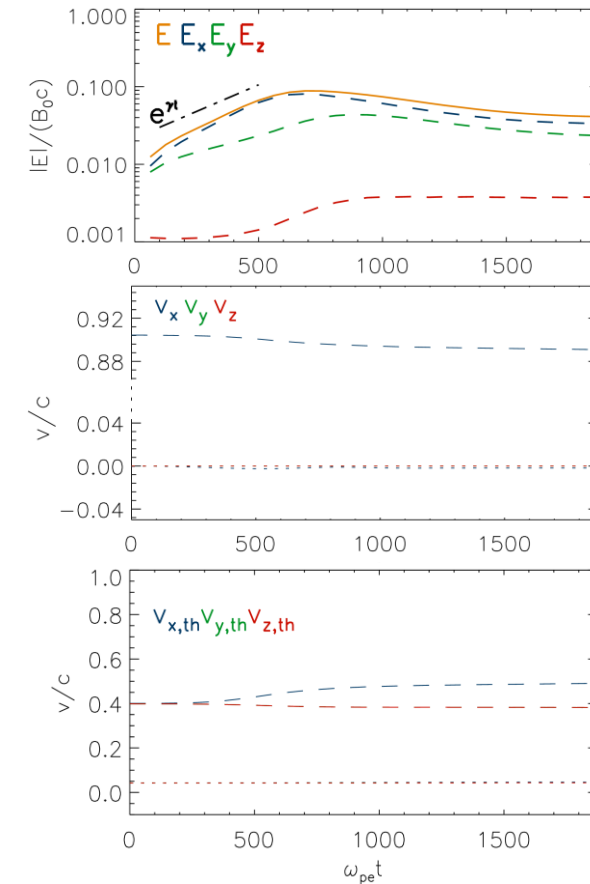
Movie – see video



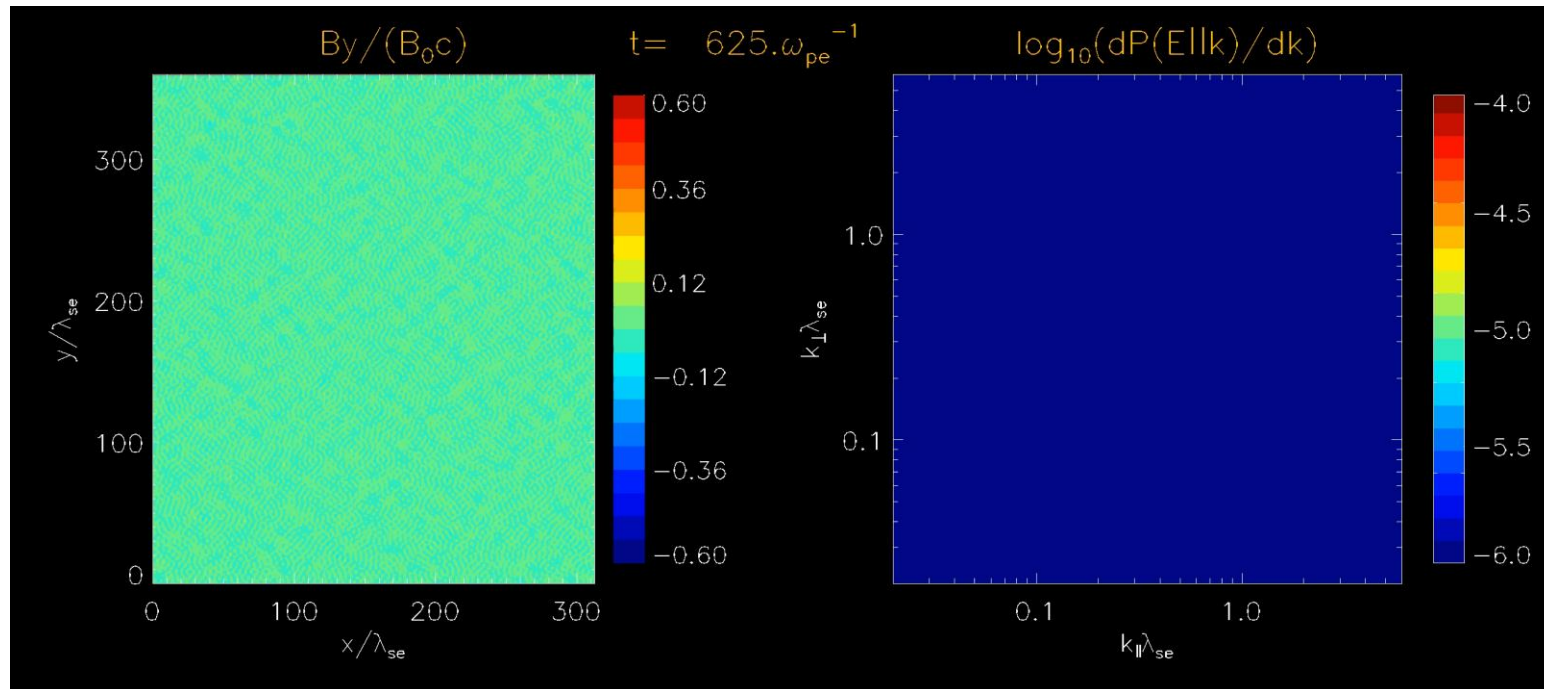
Far upstream: electron-acoustic instability



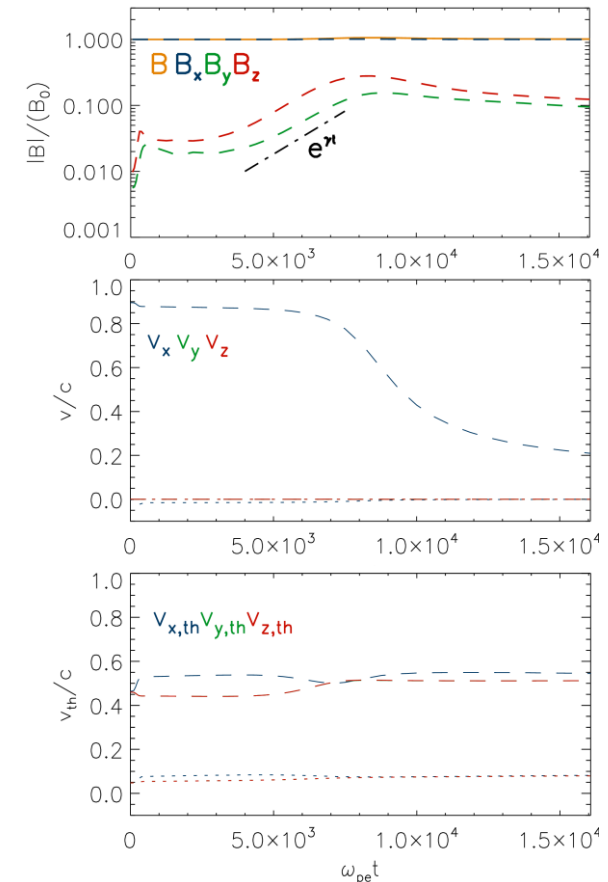
Blue lines: Electron-acoustic instability (e/e 2-stream w/ hot beam)
 Green lines: WHAMP (linear dispersion solver)
 Both calculations assume **isotropic Gaussians** with **relativistic beam-electron mass** $m_{beam} = \gamma_{beam} m_e$



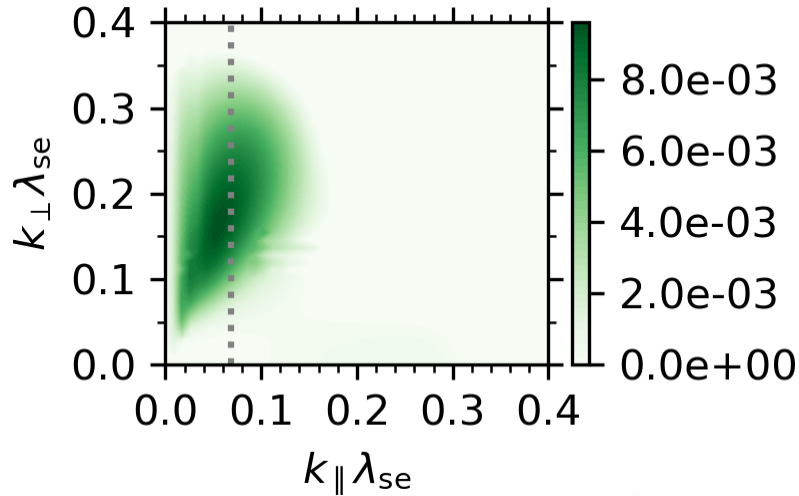
Near upstream: oblique-whistler instability



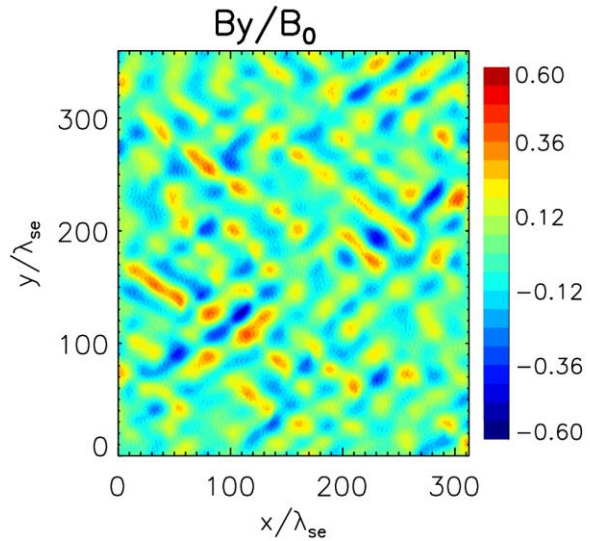
Movie – see video



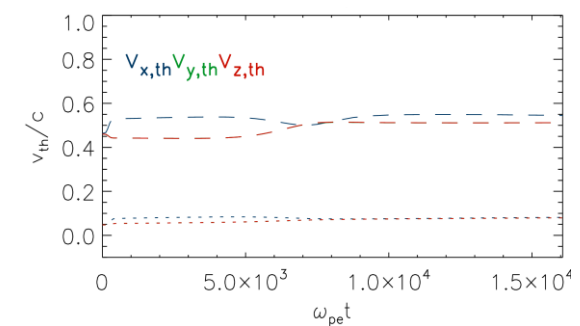
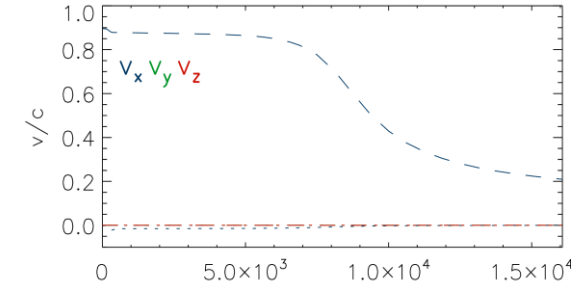
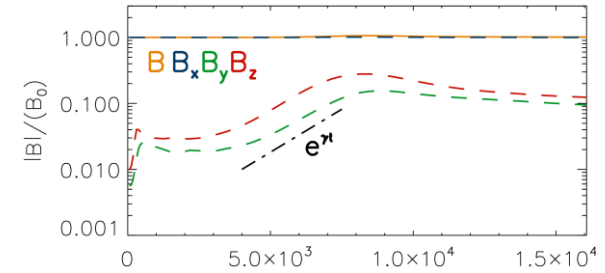
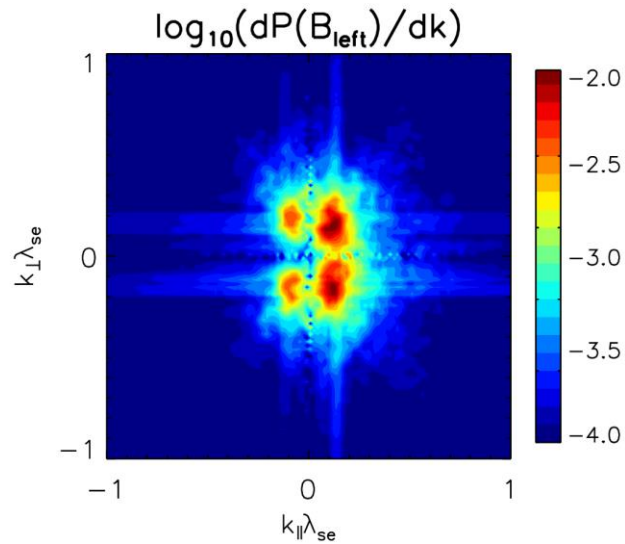
Near upstream: oblique-whistler instability



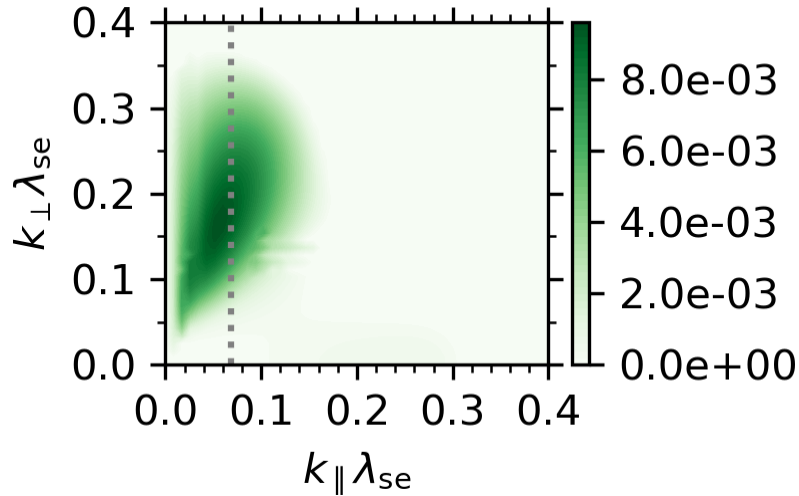
Left: Lin. dispersion solution
Bottom: Magnetic field in simulation and its helicity decomposition



Right-hand polarised waves



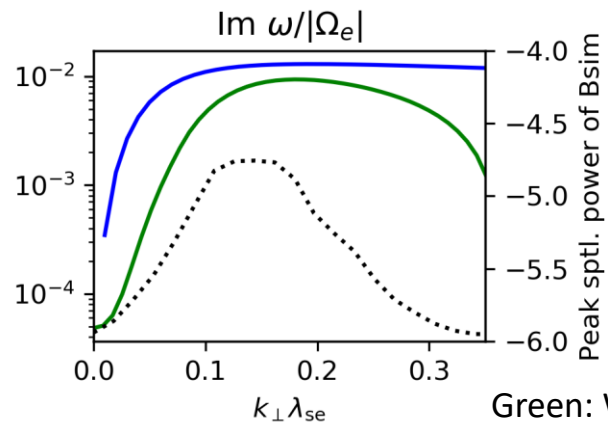
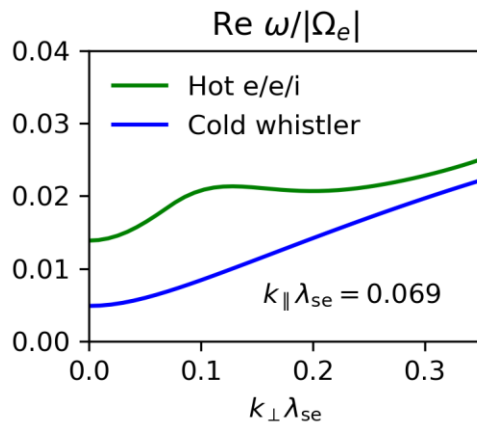
Near upstream: oblique-whistler instability



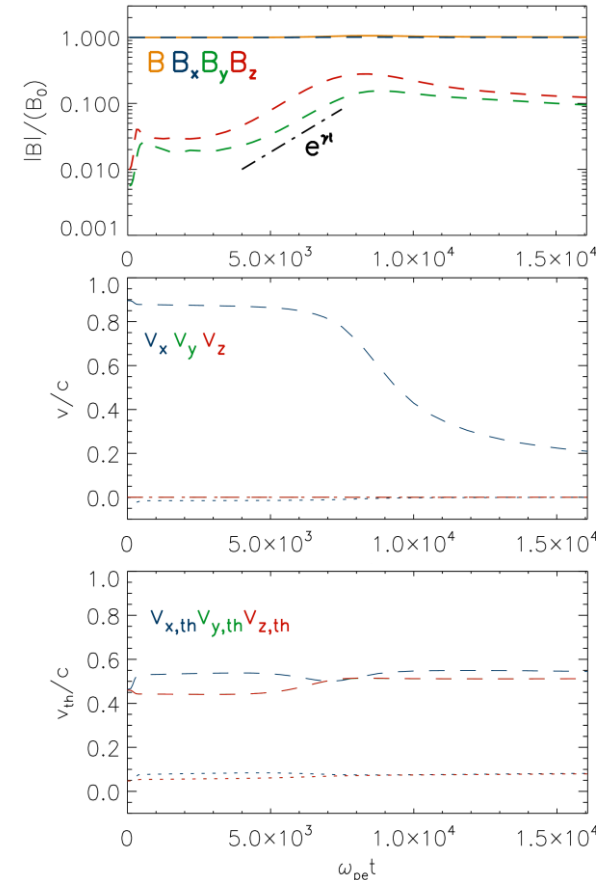
$$\omega - k_{\parallel} v_{\parallel} = 0.9944 |\Omega_e|$$

(electron gyroresonance)

Right-hand polarised waves



Green: WHAMP
 Blue lines: cold-whistler dispersion (left),
 Stepanov/Kitsenko 1961 (right)



Summary

- Shock-reflected electron beams tend to have a fairly isotropic distribution ($T_{\perp} < T_{\parallel} < 2 T_{\perp}$)
- In the **far upstream**, the electrostatic **electron-acoustic** instability excites parallel Langmuir waves
- In the **near upstream**, the denser reflected-electron beam is able to excite gyroresonant **oblique whistlers** before the shock arrives
- We observe no whistler heat-flux or electron-beam firehose instabilities

