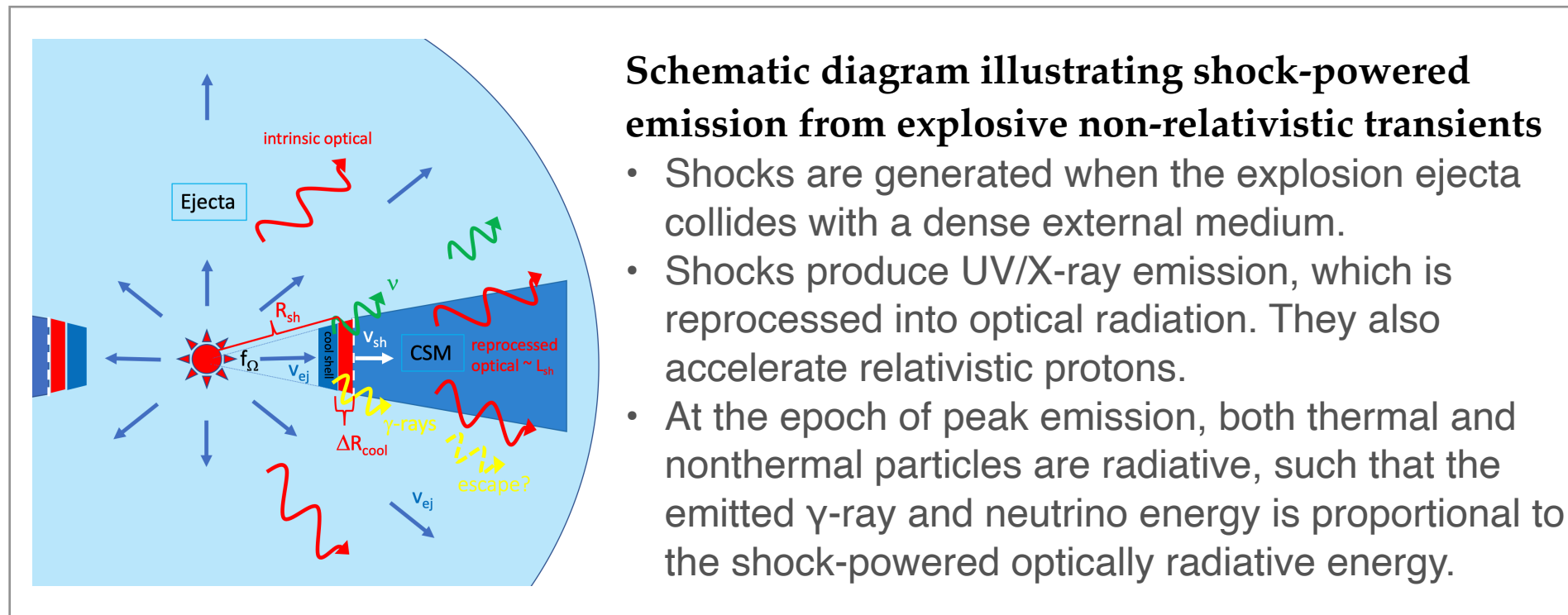


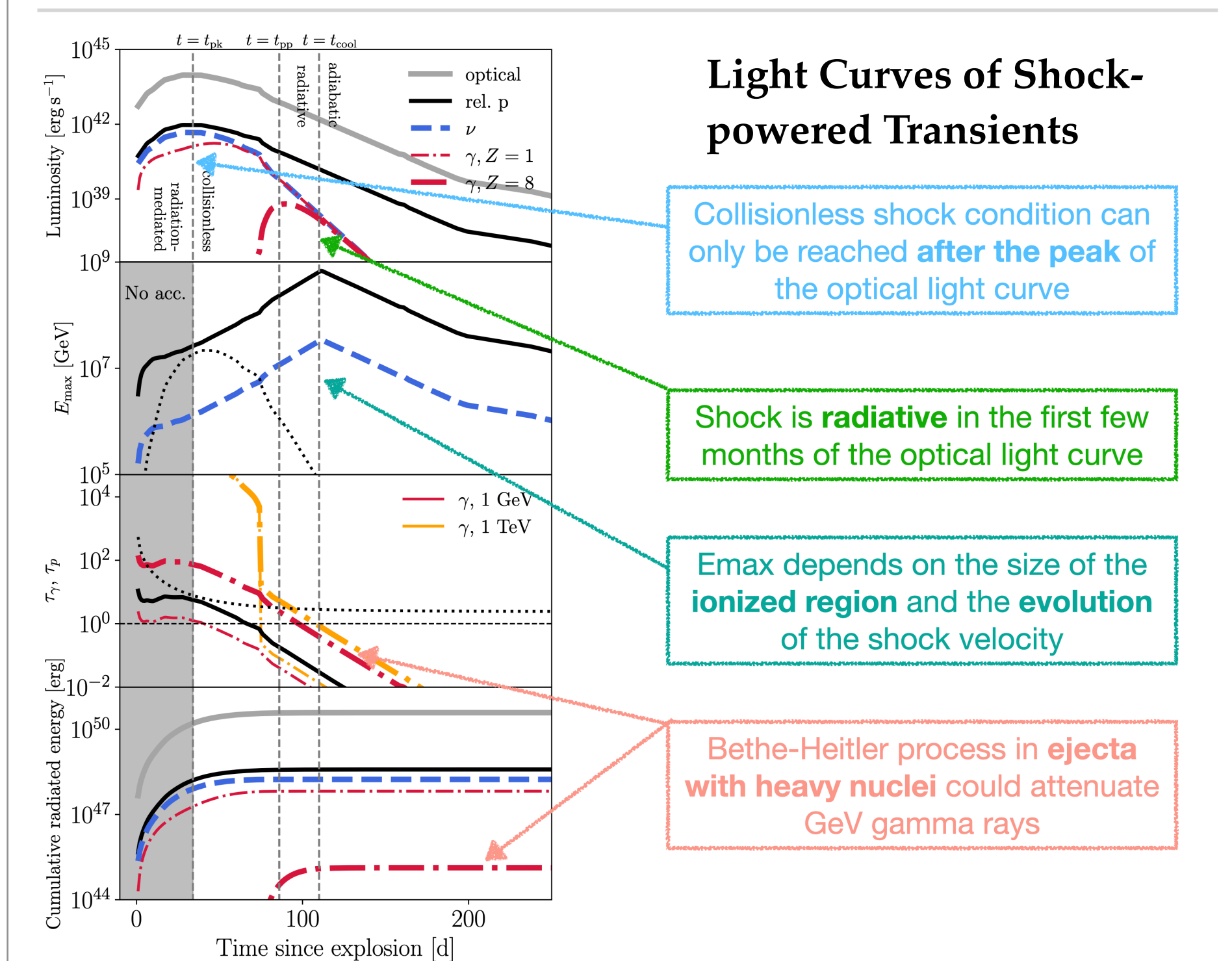
# High-Energy Neutrinos from Non-Relativistic Shock-Powered Transients

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Schematic diagram illustrating shock-powered emission from explosive non-relativistic transients

- Shocks are generated when the explosion ejecta collides with a dense external medium.
- Shocks produce UV/X-ray emission, which is reprocessed into optical radiation. They also accelerate relativistic protons.
- At the epoch of peak emission, both thermal and nonthermal particles are radiative, such that the emitted  $\gamma$ -ray and neutrino energy is proportional to the shock-powered optically radiative energy.



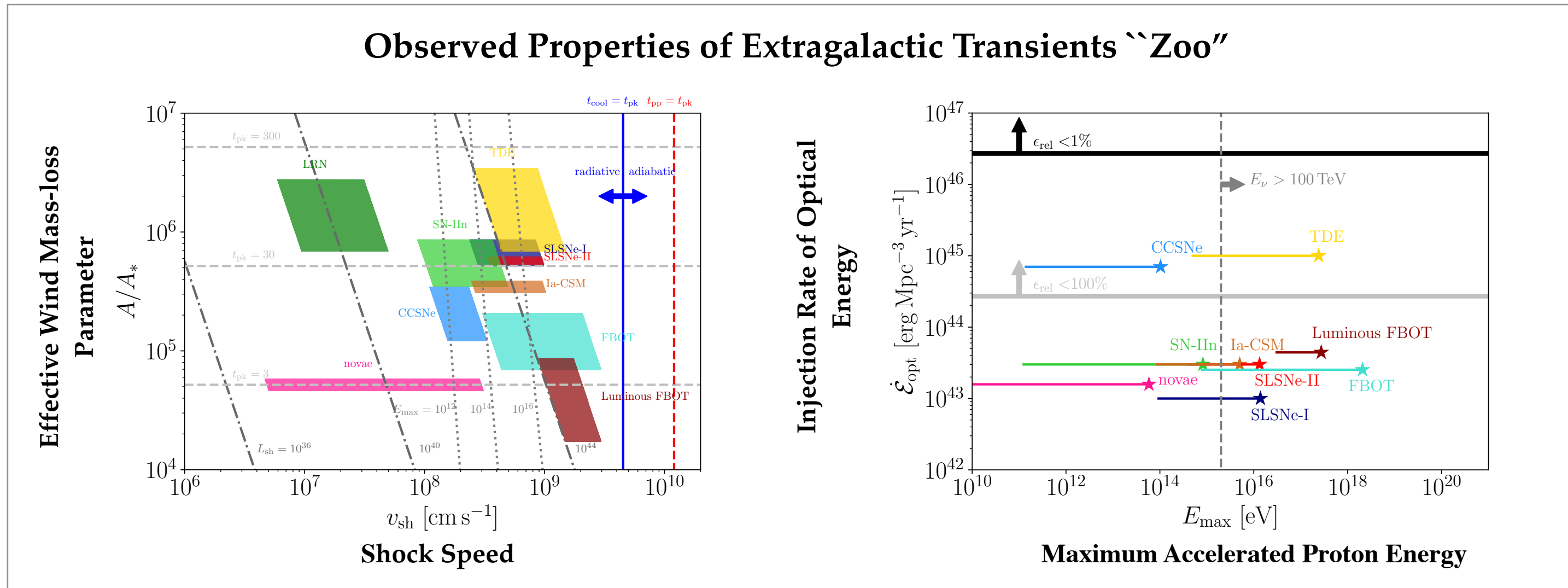
Light Curves of Shock-powered Transients

Collisionless shock condition can only be reached after the peak of the optical light curve

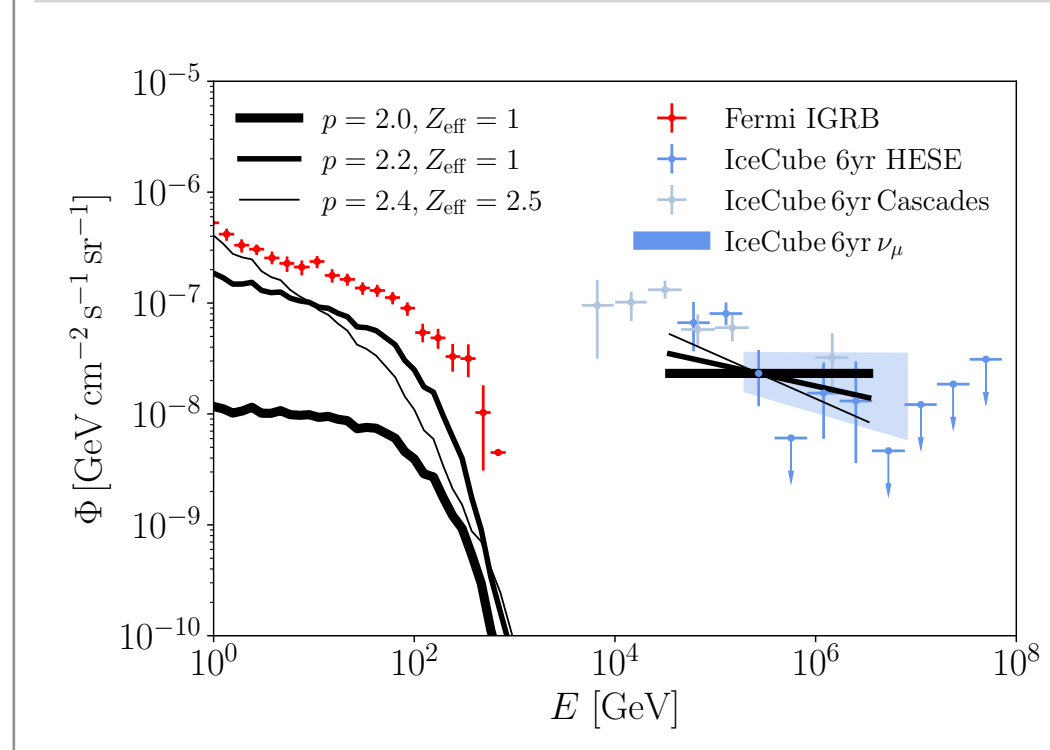
Shock is radiative in the first few months of the optical light curve

$E_{max}$  depends on the size of the ionized region and the evolution of the shock velocity

Bethe-Heitler process in ejecta with heavy nuclei could attenuate GeV gamma rays



Observed Properties of Extragalactic Transients "Zoo"



If the mass ejecta surrounding the shocks are nuclei-abundant, shock-powered transients could be gamma-ray-hidden neutrino sources.

## Conclusions:

- Correlation in optical and GeV gamma-ray light curves of classical novae suggest that they both come from radiation shocks with a particle acceleration efficiency  $< 1\%$ .
- The energy radiated by non-thermal ions in high-energy neutrinos and gamma rays is directly proportional to the transient's shock-powered optical energy.
- Based on observed and inferred energetics of various classes of optical transients, we find that transients powered by non-relativistic shocks can barely explain the IceCube diffuse neutrinos.
- Adiabatic shocks may have a cosmic-ray luminosity higher than the optical luminosity, but for most CSM density profiles, the total shock-dissipated energy is dominated by early times when the shock is radiative.