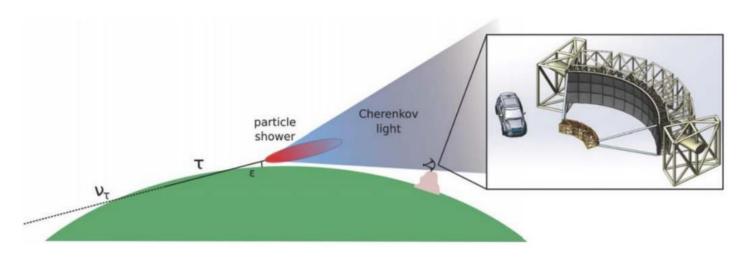




Trinity: An Imaging Air Cherenkov Telescope to Search for Ultra-High-Energy Neutrinos

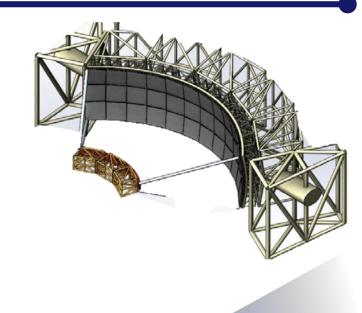


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What is Trinity?



- Trinity is a proposed 18 telescope-strong network of dedicated Imaging Air Cherenkov Telescopes (IACTs) optimized to detect Earth-skimming tau neutrinos.
- Trinity will build on the rich heritage of IACTs, but also take of new technology designed for CTA & EUSO-SPB to reduce costs and improve sensitivity.
- Collaboration of US-UK-IT institutes.
- Currently in the design phase.
- NSF funding has recently been awarded to create a demonstrator system.





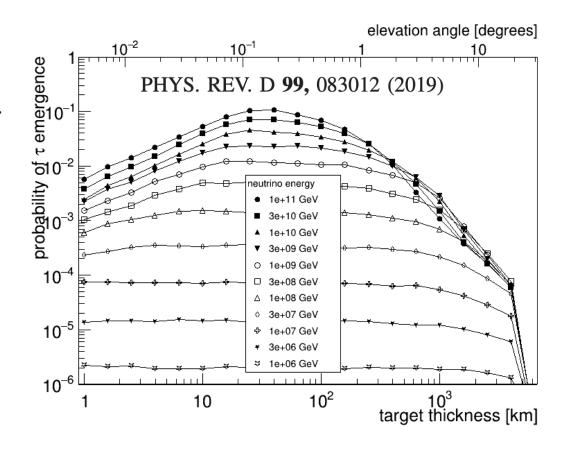
PHYS. REV. D 99, 083012 (2019)

Earth-skimming tau neutrinos



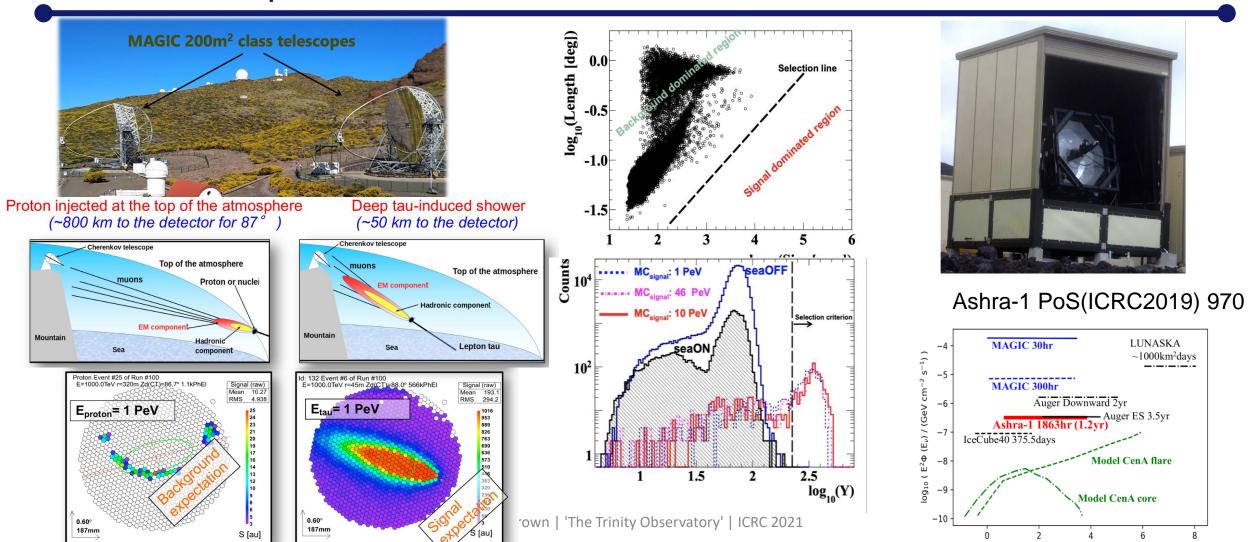
- 1-10⁴ PeV (UHE) tau neutrinos can pass through hundreds of km of bedrock, with a high probability that a tau lepton emerging before it decays
- Tau decays initiates an up-wards going air-shower
- A wide FoV IACT can image the Cherenkov radiation from this air shower to infer the presence and energy of the original tau neutrino

Decay	Secondaries	Probability	Air-shower
$ au ightarrow \mu^- ar{ u}_\mu u_ au$	μ^-	17.4%	weak showers
$ au ightarrow e^-ar{ u}_e u_ au$	e^{-}	17.8%	1 Electromagnetic
$ au ightarrow \pi^- u_ au$	π^-	11.8%	1 Hadronic
$ au ightarrow \pi^-\pi^0 u_ au$	$\pi^-,\pi^0 o 2\gamma$	25.8%	1 Hadronic, 2 Electromagnetic
$ au ightarrow \pi^- 2\pi^0 u_ au$	$\pi^-, 2\pi^0 \rightarrow 4\gamma$	10.79%	1 Hadronic, 4 Electromagnetic
$ au ightarrow \pi^- 3 \pi^0 u_ au$	$\pi^-, 3\pi^0 \rightarrow 6\gamma$	1.23%	1 Hadronic, 6 Electromagnetic
$ au ightarrow \pi^-\pi^-\pi^+ u_ au$	$2\pi^-,\pi^+$	10%	3 Hadronic
$ au ightarrow \pi^- \pi^+ \pi^- \pi^0 u_ au$	$2\pi^-,\pi^+,\pi^0 o 2\gamma$	5.18%	3 Hadronic, 2 Electromagnetic



Past efforts with imaging telescopes

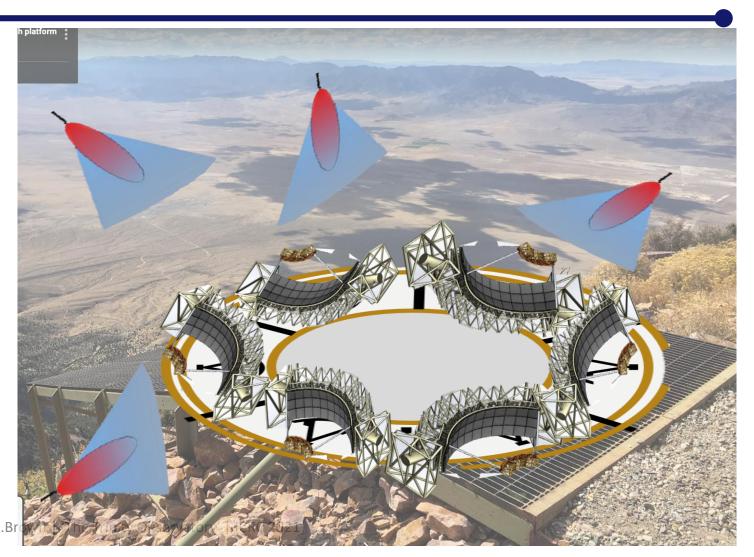




Recipe for Trinity

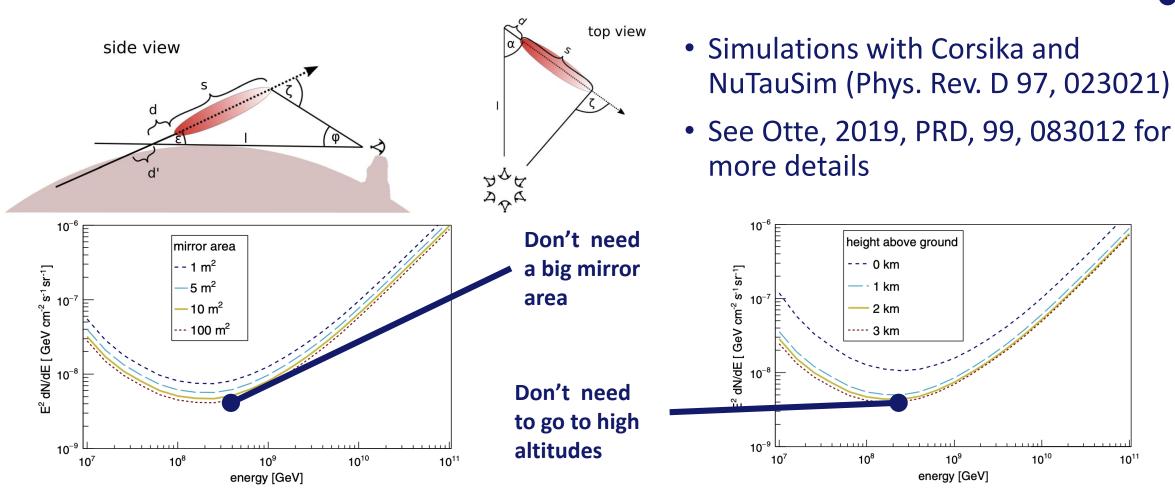


- Array of 6 wide FoV telescopes
- Place on a mountain peak
- Arranged in a circle
- Covering the entire horizon
- Operate at night (1200 h/y)
- Operate remotely
- Goal to have 3 Trinity site (the first of which will be Fresno peak, Utah)



Telescope size & altitude



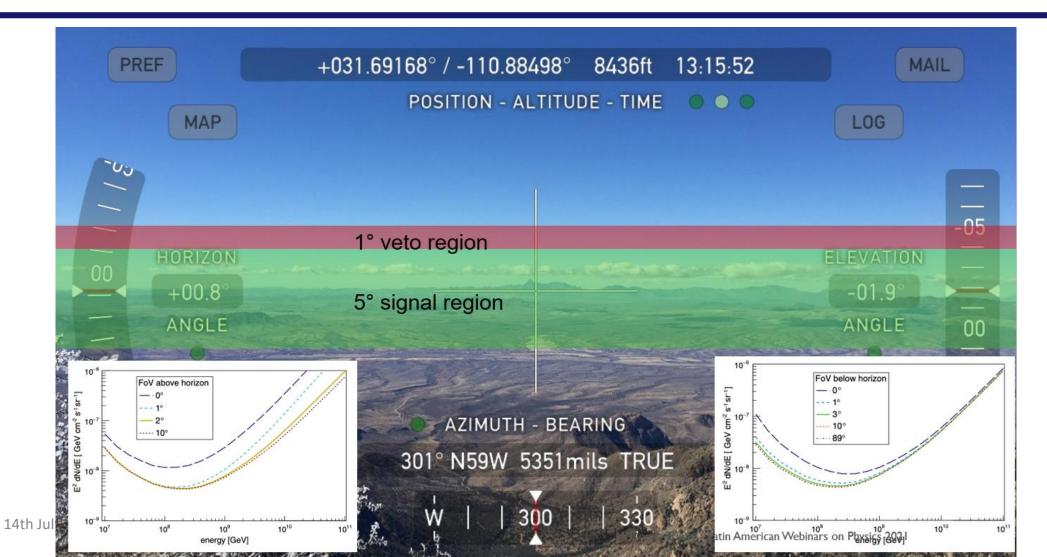


10¹¹

10¹⁰

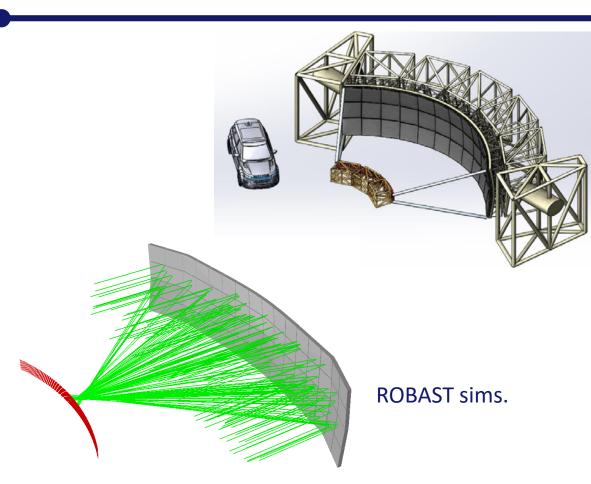
View of the horizon...





Telescope OSS





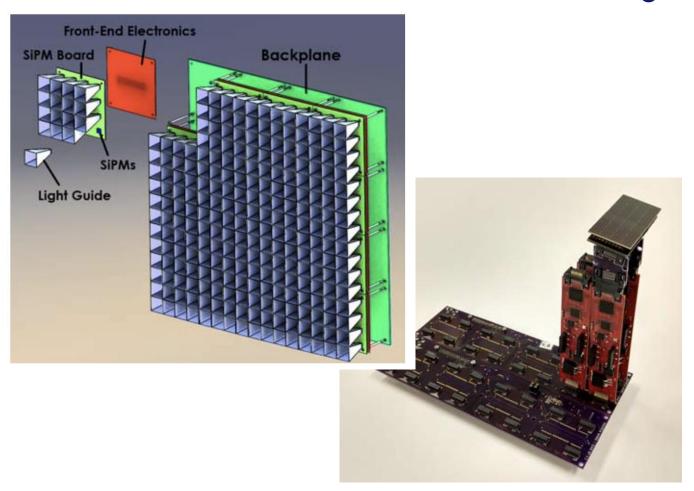
- Spherical mirror profile.
- Focal length ~4.2m.
- 0.3 deg. angular resolution.
- 60 deg x 5 deg FoV.
- 12 x 3m total mirror area.
- Fixed in azimuth.
- Movement in elevation only.

Preliminary drawing at INFN-DFA Padova

Camera



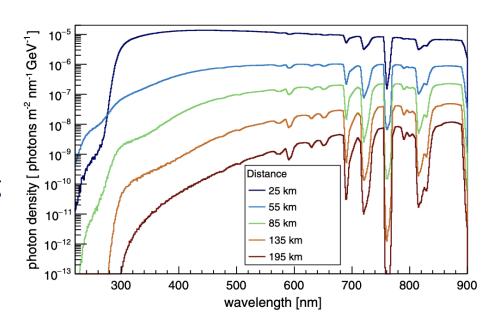
- Camera based upon the EUSO-SPB2 Cherenkov telescope
 - PMMA light guide
 - 3300 SiPMs (from CTA development)
 - MUSIC ASIC
 - AGET digitizer (from EUSO-SPB2)



Atmospheric Calibration

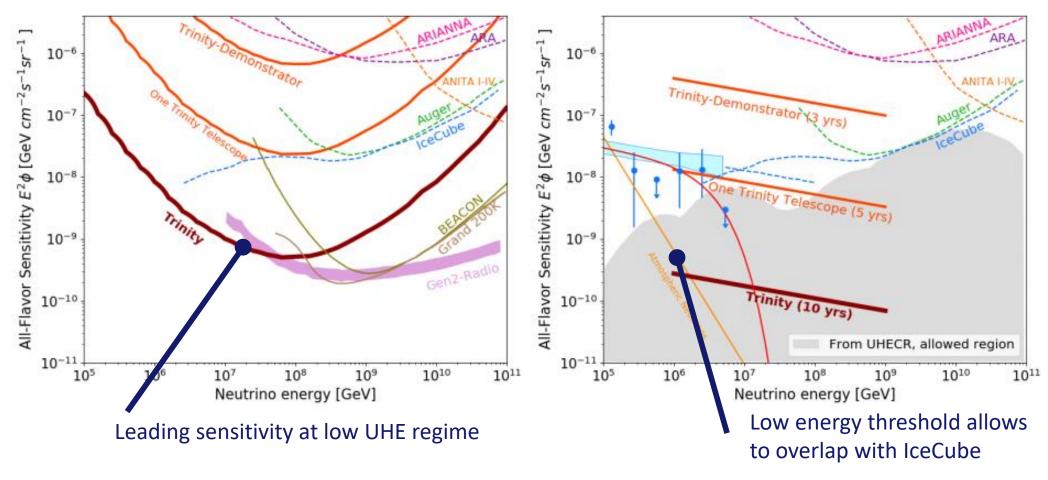


- Will observe air shower in ground-layer of atmosphere where there is a high and variable amount of atmospheric dust: implies we need to monitor atmospheric transmission
- Will have 3 separate, independent approaches
 - Monitor stars close to the horizon (Ebr et al. 2021)
 - Place calibration MWL light beacons on surrounding mountains (Weincke et al. 1999)
 - Periodically fly a UAV to cross-check the two approaches (Brown, 2018)



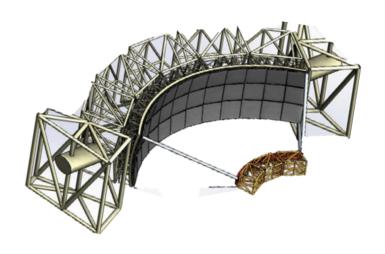
Expected Performance



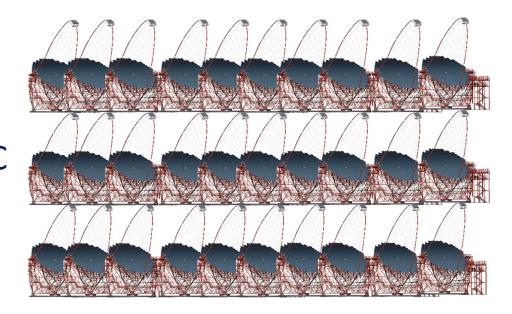


Sensitivity in a nut-shell





= many x MAGIC



See Andrew Wang, et al. poster for a full description of Trinity's sensitivity calculations

Outlook and take-home points



- Trinity will use dedicated wide FoV IACTs to view the entire horizon from a mountain top, to search for Earthskimming tau neutrinos
- Trinity will have the unique capability of a low-energy threshold allow it to observe the tail of the IceCube diffuse neutrino spectrum.
- Trinity will play a crucial role bridging the gap between IceCube's known signal, and expected sensitivity of UHE radio detectors.

