

# Effelsberg Monitoring of AGN Jets with VHE Astroparticle Emissions

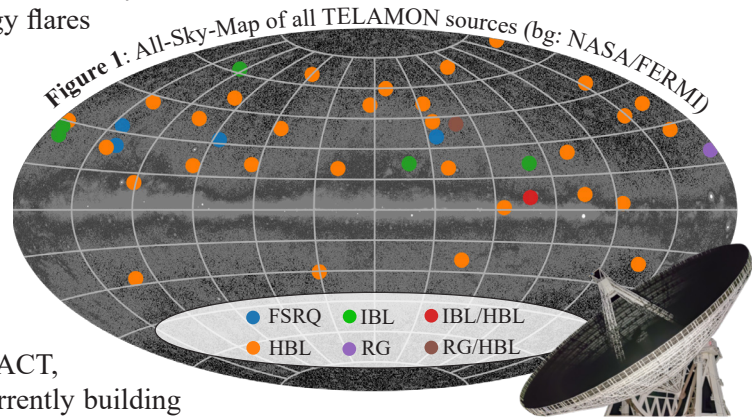
M. Kadler<sup>1</sup>, U. Bach<sup>2</sup>, D. Berge<sup>3</sup>, S. Buson<sup>1</sup>, D. Dorner<sup>1</sup>, P.G. Edwards<sup>4</sup>, F. Eppel<sup>1</sup>, M. Giroletti<sup>5</sup>, A. Gokus<sup>1,6</sup>, O. Hervet<sup>7</sup>, J. Heßdörfer<sup>1</sup>, S. Koyama<sup>8</sup>, A. Kraus<sup>2</sup>, T.P. Krichbaum<sup>2</sup>, E. Lindfors<sup>9</sup>, K. Mannheim<sup>1</sup>, R. de Menezes<sup>10</sup>, R. Ojha<sup>11</sup>, G.F. Paraschos<sup>2</sup>, E. Pueschel<sup>3</sup>, F. Rösch<sup>1</sup>, E. Ros<sup>2</sup>, B. Schleicher<sup>1</sup>, J. Sinapius<sup>3</sup>, J. Sitarek<sup>12</sup>, J. Wilms<sup>6</sup>, M. Zacharias<sup>13</sup>

## What is the TELAMON program?

The TELAMON program is using the Effelsberg 100-m telescope to monitor the radio spectra of active galactic nuclei (AGN) under scrutiny in astroparticle physics, namely **TeV blazars and candidate neutrino-associated AGN**. We perform high-frequency observations every 2-4 weeks at multiple frequencies up to 44 GHz. We aim to characterize the radio variability of very-high-energy emitting AGN jets and trace dynamical processes in the pc-scale jets of blazars related to high-energy flares or neutrino detections. Our sample covers about 40 sources and is dominated by high synchrotron peaked objects.

## What is unique about TELAMON?

The large Effelsberg dish can yield superior radio data over other programs for very-high-energy (VHE) emitting blazars, which are often faint radio sources. Moreover, we are coordinating our observations with FERMI-LAT, FACT, H.E.S.S. and VERITAS AGN monitoring groups. We are currently building a unique sample of TeV- and neutrino-associated AGN, excluding bright low-peaked blazars, which are well covered in other monitoring programs. As a selection criterion, we include all sources whose low-state flux density falls below 500 mJy. Sources south of 30° are also observed by ATCA. This leads to a sample that is complete (down to 10-20 mJy) for HBLs. The TELAMON target list is dynamically updated with new neutrino detections.



## Example Source: S2 0109+22

For each monitored source, we derive continuous dynamic spectra and light curves. Figure 2 shows as an example the source S2 0109+22 which exhibits flaring activity with a continuous increase in flux density over about 100 days both at 14mm and 7mm. Our sampling rate is well suited to sufficiently resolve such time scales.

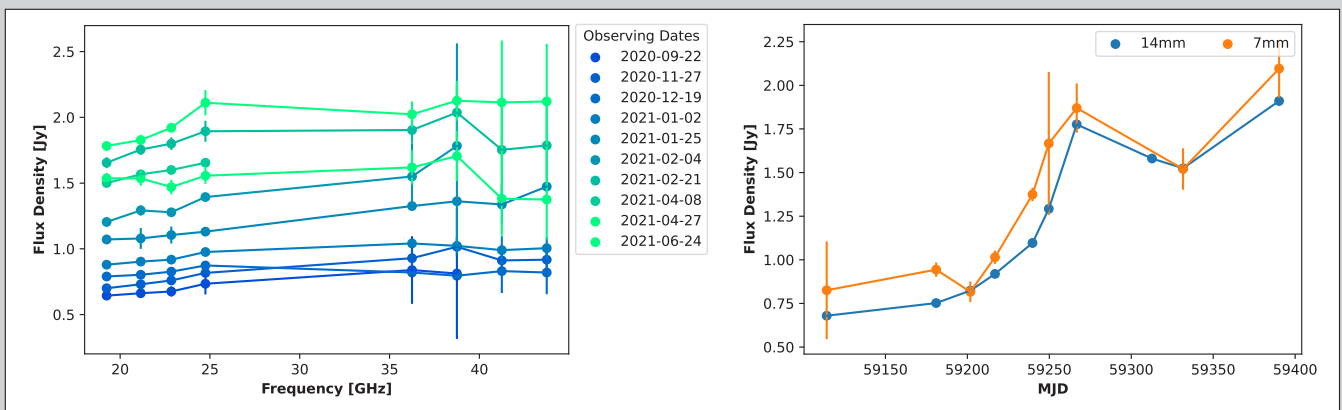


Figure 2: Example results of spectra (left) and light curves (right, averaged over all subbands) for the source S2 0109+22.

<sup>1</sup>Institut für Theoretische Physik und Astrophysik, Universität Würzburg - <sup>2</sup>Max Planck-Institut für Radioastronomie Bonn -

<sup>3</sup>DESY Zeuthen - <sup>4</sup>CSIRO Astronomy and Space Science Australia - <sup>5</sup>INAF-Istituto di Radioastronomia Bologna

<sup>6</sup>Dr. Karl Remeis Observatory and ECAP - <sup>7</sup>Santa Cruz Institute UCSC - <sup>8</sup>Institute of Astronomy and Astrophysics Taipei

<sup>9</sup>Finnish Centre of for Astronomy with ESO - <sup>10</sup>Universidade de São Paulo

<sup>12</sup>NASA HQ - <sup>13</sup>University of Łódź