

On the nature of primary particles producing air showers with energies greater than 5 EeV



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Abstract. To study the nature of particles with energies greater than 5 EeV, the database of the Yakutsk array was analyzed. Showers coming one after the other are highlighted within a time interval of 1-20 hours. Some periodicity was found in the registration of such showers during the daily observation cycle with an average time of T = 8 hours. The characteristics of the selected showers: energy, zenith and azimuthal angles were found to be close in magnitude. Consequently, we can assume the same origin nature of the primary particles that initiate such showers. Existing discrepancy in the arrival time of showers at the Earth's level can be attributed to the participation in various processes in outer space: the interaction of particles with different charges with galactic magnetic field, acceleration of particles due to the frictional mechanisms followed by re-emission with higher energy. And time delay at the shock front. If this hypothesis is correct, then the analysis of such air shower events will make it possible to obtain information on the processes of interaction of shock waves with the matter of the universe.



Table 1. Pairs of showers characteristics registered at the Yakutsk array

| | | | | | | | | | Time difference between two showers is $\Delta T = 14.4$ h, zenith and | | | |
|----------|----------|---------------------|-------|-------|------|-------|------|-------|--|------|--|--|
| Data | Time | lg(E ₀) | cosθ | φ | δ | α | b | 1 | ρ_{μ}/ρ_{s} | Xmax | azimuth angles are close. Other characteristics are different - | |
| 22.01.14 | 07:08:49 | 18.71 | 0.576 | 195.9 | 60.5 | 352.2 | -0.5 | 113.3 | 0.047 | 1048 | energy of the first shower is smaller, which can be explained by | |
| 22.01.14 | 21:49:08 | 19.03 | 0.687 | 188.5 | 65.4 | 193.5 | 52 | 122.2 | 0.410 | 749 | very low X_{max} and low muon fraction. It can be underdeveloped | |
| | | | | | | | | | | | air shower produced by proton, in this case energy would be | |

underestimated.

On the other hand, it can be produced by γ -ray, which is possible considering low depth of maximum and low muon fraction $\rho_{\mu'}\rho_s$.

Table 2. Pairs of showers with the most similar properties.

| n/n | Date | ∆t [h] | ΔlgE | Δθ [°] | Δφ [''] | õ1 [°] | α1 [°] | ô₂ [°] | α ₂ [°] |
|-----|----------|--------|------|--------|---------|--------|--------|--------|--------------------|
| 1. | 18.04.03 | 12.13 | 0.03 | 11.2 | 148.1 | 47.7 | 293.0 | 54.4 | 337.4 |
| 2. | 02.05.03 | 07.15 | 0.08 | 9.9 | 137.0 | 55.3 | 332.0 | 64.0 | 337.0 |
| 3. | 31.03.04 | 01.51 | 0.01 | 14.4 | 16.1 | 67.7 | 146.8 | 71.8 | 123.0 |
| 4. | 22.01.09 | 11.11 | 0.01 | 7.2 | 86.7 | 65.9 | 141.2 | 62.3 | 318.2 |

This table shows air showers with closest characteristics. Δt is 12 and less hours, energy difference is smaller than factor of 1.08 and very close galactic coordinates – less than 5°.

Their locations on the sky map are within a circle with radius of 5 °. In this case we can assume that these air showers are originated from the same source of cosmic rays.

Distribution of the arrival of EAS events with $E_0 \ge 5 \text{ EeV}$



The figure shows the distribution of air showers with $E_0 \ge 5$ EeV on the sky map, coming one after another with an interval of less than 24 hours. Diamonds - first showers and stars second showers, showers from table 2. Crosses are showers with energies with $E_0 \ge 10$ EeV, registered by the Yakutsk radio array, during the observation periods of 1986-1989 and 2009-2018. Triangles – air showers with $E_0 \sim 100$ EeV; squares show air showers with very low muon content - candidates for gamma-ray produced air showers. In addition, the boundaries of the most active regions with X-ray, radio and optical sources are plotted: the constellation Ursa, Virgo, M82, and Markarian 421. In addition, a hot spot found by TA data is plotted. It can be seen that some of the showers coincide in their coordinates or are close to the boundaries of these regions, which indicates that active regions along with other sources may be sources of cosmic rays with ultra-high energies.

Summary

Judging by the analysis of paired showers, the nature of the primary particles producing air showers is diverse. Not all paired events have close declination and right ascension. Perhaps some part of the events diverged more due accuracy of the zenith angle determination. On the other hand, the discrepancy can be influenced by the fact that paired particles can have different charges and, hence, the magnetic field of the shock wave will affect the trajectory of these particles in different ways. At the same time, there are paired events in which both declination and right ascension are quite close. There are much fewer such events among the selected shower pairs. On the table 2 showers with closes characteristics are shown. However, even for those showers, it is clear that the discrepancy between the galactic coordinates of showers is significant. Unfortunately, using experimental data and known active sources we can't explicitly tell from which region of the celestial sphere those showers are concentrated near the galactic plane, and some near metagalactic plane. The absence of a these sources yet.

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