

ICRC 2021, Berlin, Germany, July 16, 2021

Cosmic Rays from the Termination Shock to the Heliopause: The Role of the Heliospheric Current Sheet

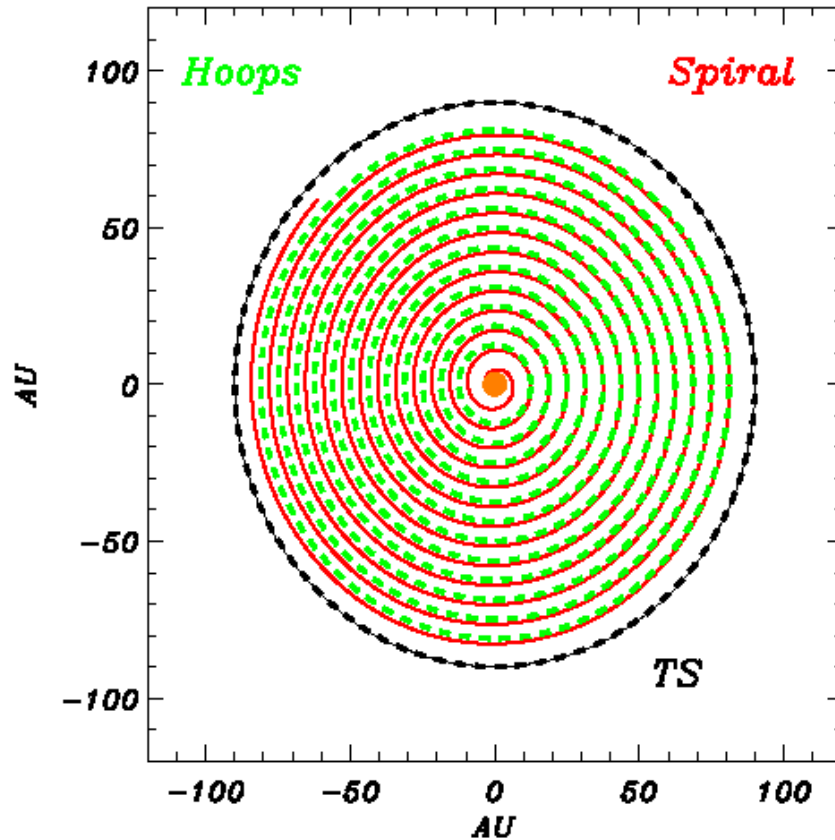
József Kóta

The University of Arizona

Tucson, Arizona, USA

jkota@email.Arizona.edu

Acceleration of ACRs at the TS: Possible Role of the HCS



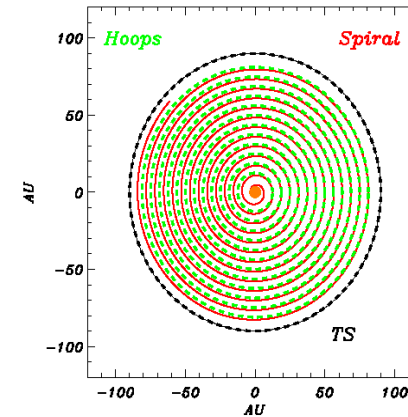
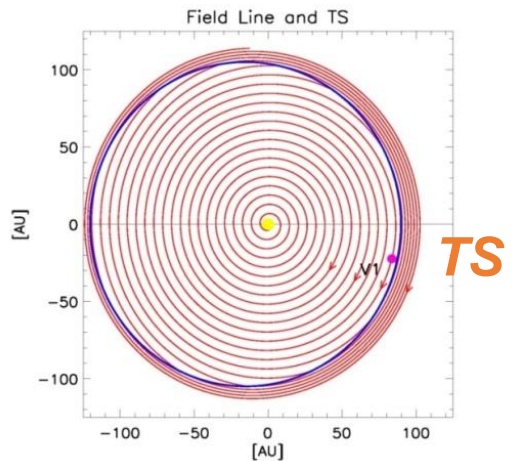
Motivation: GCR- ACR disparity at solar minima: GCRs record high, ACRs are not

Question: What is the effect of the HCS on the accelerated source spectrum of ACRs at the termination shock (TS) ?

We adopt a “**hoop model**” that captures the most essential effects of the wavy HCS in a simple 2D time dependent calculation. The model works well beyond >10-20 AU distances from the Sun. It

Hoop model of Parker Spiral Field

Solve Parker's equation in 2-D + energy using ad-hoc parameters



26-day rotation + radial solar wind = tightly wound spiral field

CR transport at $r > 10$ AU is **dominantly across the spiral field**

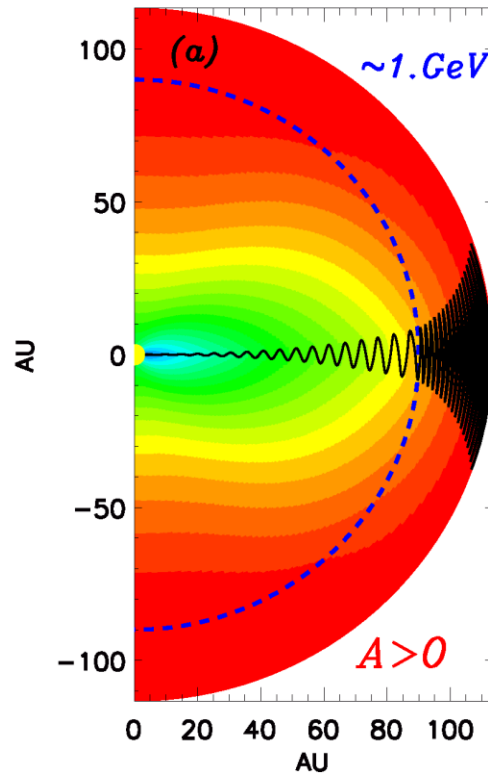
(it is faster to climb the wall than going around – spiral is too long)

Substitute spiral with **hoops**: azimuthal symmetry with wavy HCS
seamlessly changing the tilt angle continuously (**time-variations**)

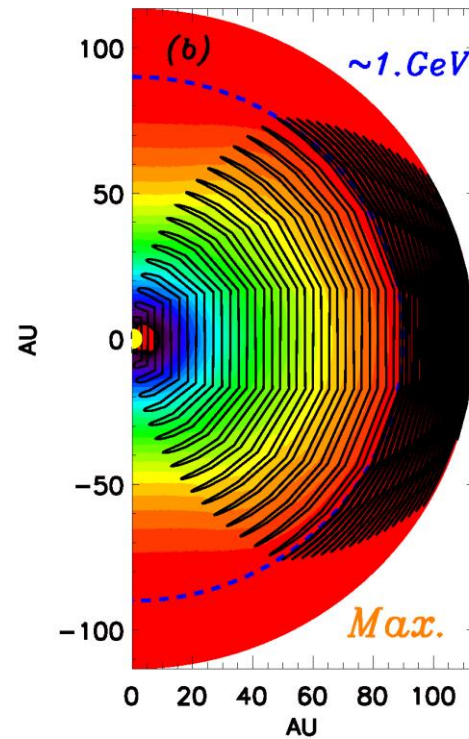
22 Year variation of GCRs & ACRs in the hoop model

GCRs:

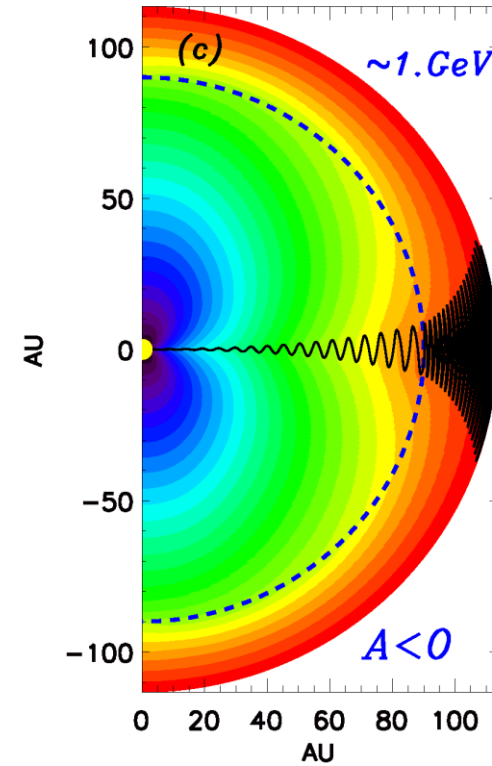
Solar Min $A > 0$



Solar Max.

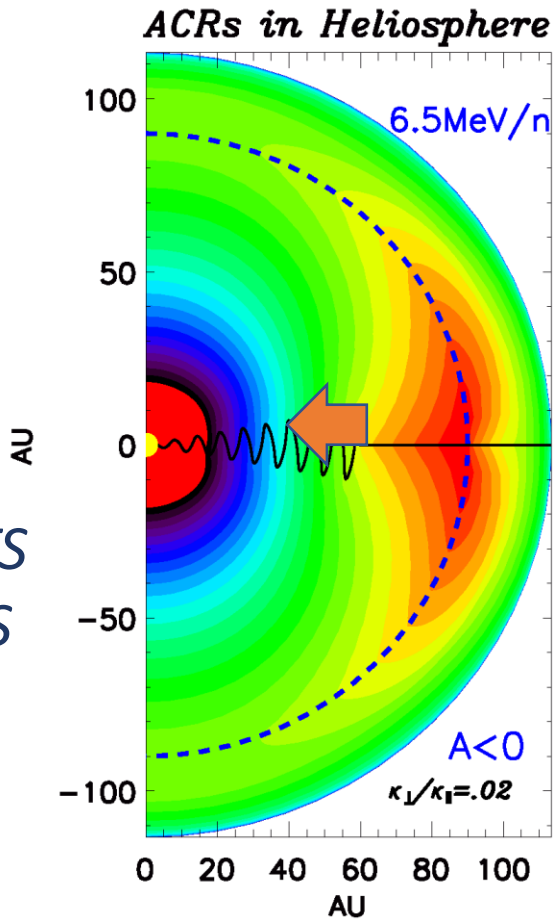


Solar Min. $A < 0$



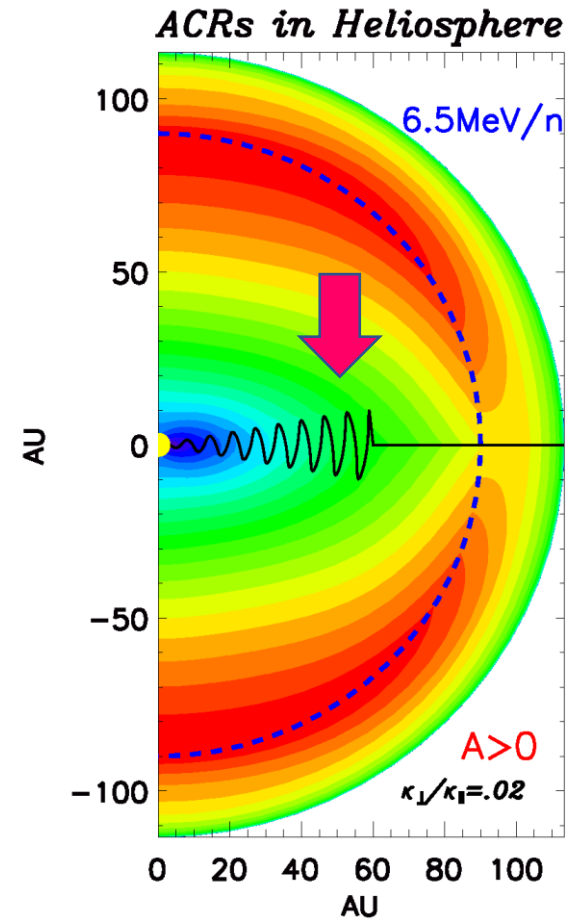
ACRs: Simulations starting from flat HCS (Solar Minima)

*A < 0 Drift:
Equatorward along TS
Inward along the HCS
Best injected at Poles*



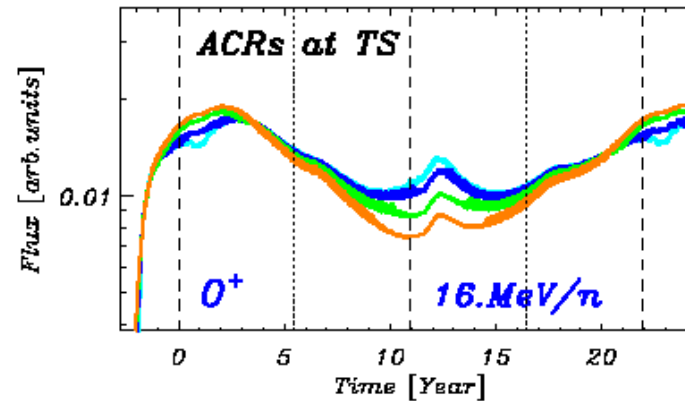
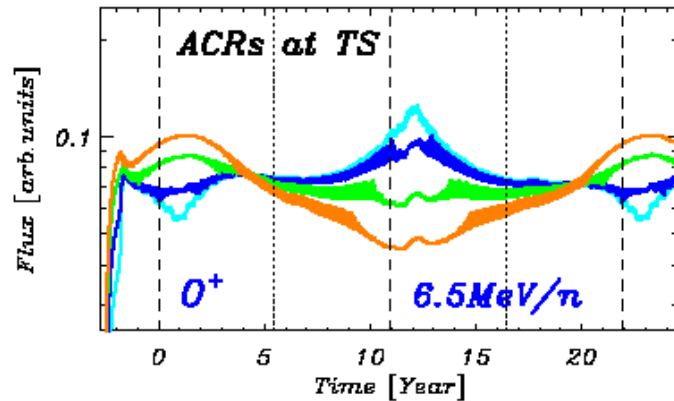
A < 0

*A > 0 Drift:
Poleward along TS
Pole to Equator
Best if injected near
Equator*



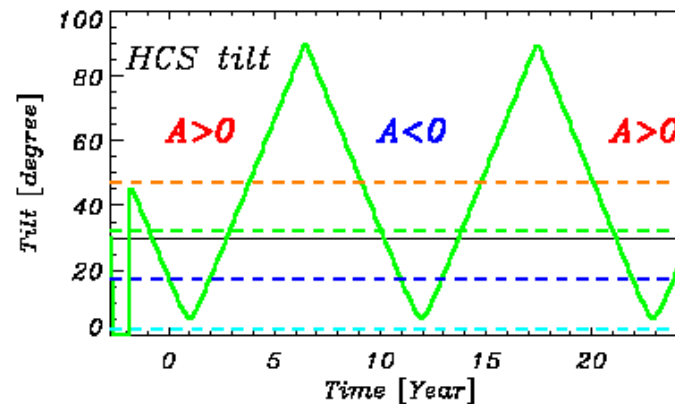
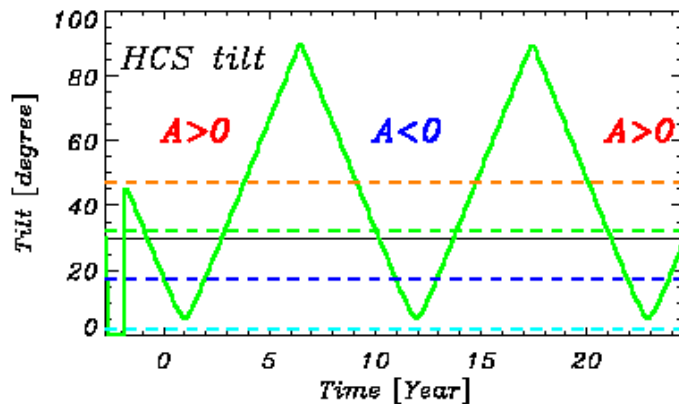
A > 0

Example: 22-year cycle in the accelerated source spectrum of ACRs *at the TS* at different latitudes



22 yr Cyclic variation
Is energy dependent
& can be significant

Line width indicates
26-day variations



The only quantity to
change is the *tilt-angle*
of the HCS.

Polarity is reversed
seamlessly.

Motivation and Summary

- *GCRs and ACRs tend to track each other quite closely. During recent solar minima, however, GCRs reached record high level, ACRs did not (Mewaldt, 2010)*
- *GCRs and ACRs have similar transport properties but distinctly different origin. Their disparity suggests that the source spectrum of ACRs at the TS may have been weaker.*
- *Weaker source spectrum can be caused by a number of possible reasons (Moraal & Stoker, 2010; Leske et al. 2013).*
- *Here we single out the HCS and mimic a 22-year cycle, where the HCS tilt changes, while everything else remain constant.*
- *The changing polarity and tilt leads to time variations of the source spectrum which may be significant*
- *Drift & cross field diffusion are coupled (think of the inverse κ tensor). Drift along the wavy HCS, can, among others, increase the “effective” latitudinal transport and make the distribution of ACRs ‘more spherical’.*