

Observation of large-scale anisotropy in the arrival directions of cosmic rays with LHAASO-KM2A

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高海拔宇宙线观测站

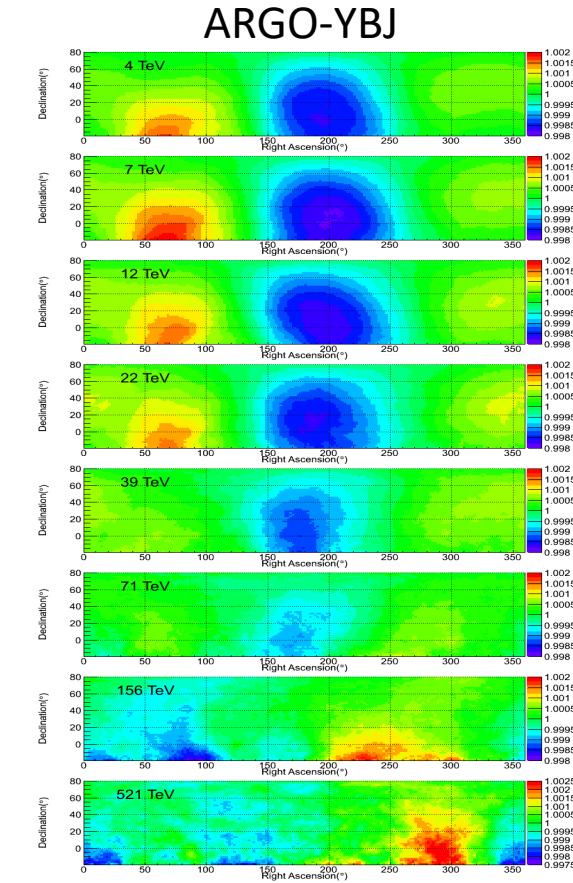
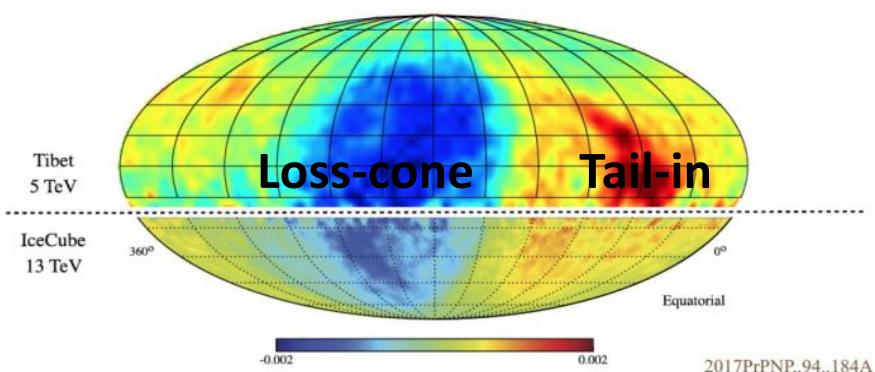
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Introduction

- **Anisotropy :**

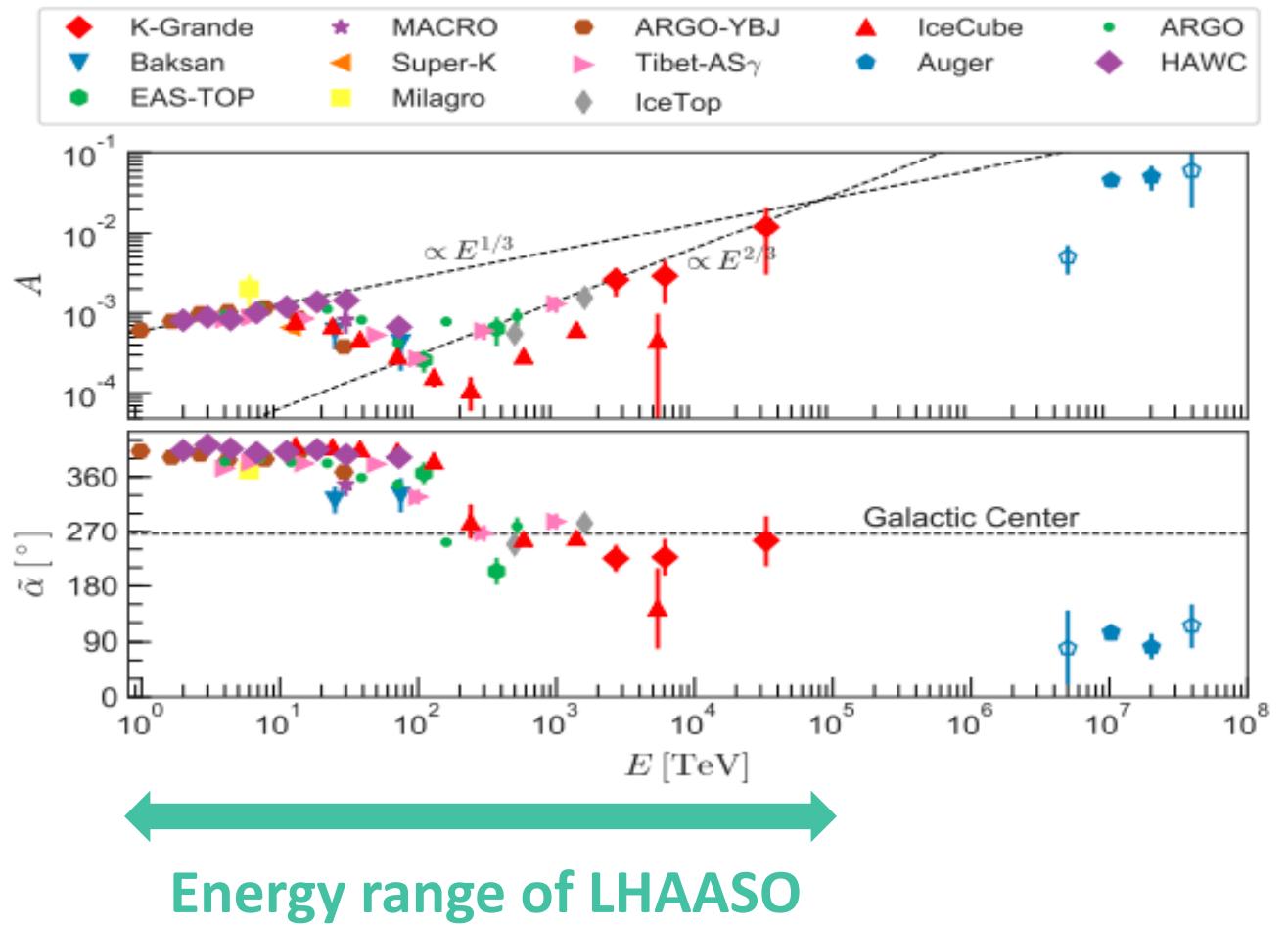
- a probe to study the origin and propagation of cosmic rays.
- the structure and intensity of anisotropy are obviously energy dependent.



Bartoli et al. 2018

- Puzzles :**

The irregularity of energy dependence of amplitude and phase –Why ?



LHAASO

- Large High Altitude Air Shower Observatory (LHAASO), 4410 m a.s.l. @ Daocheng, Sichuan, China

Kilometer Array (KM2A):

5195 Electromagnetic particle Detectors (EDs)

1188 Muon Detectors (MDs)

1.3 km² area

Water Cherenkov Detector Array (WCDA):

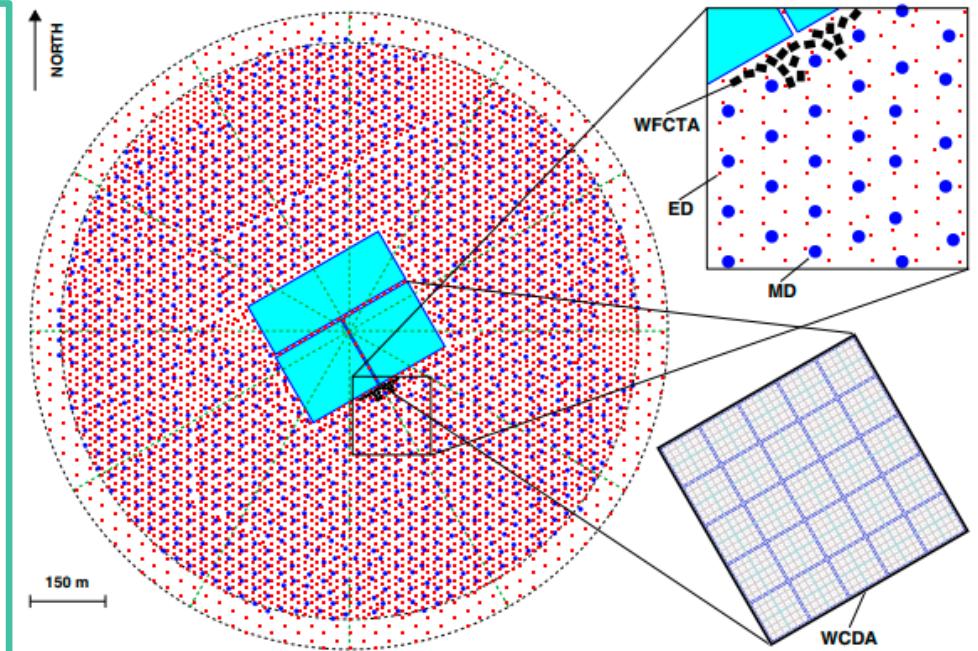
3 water ponds

3120 detectors

78,000 m² area

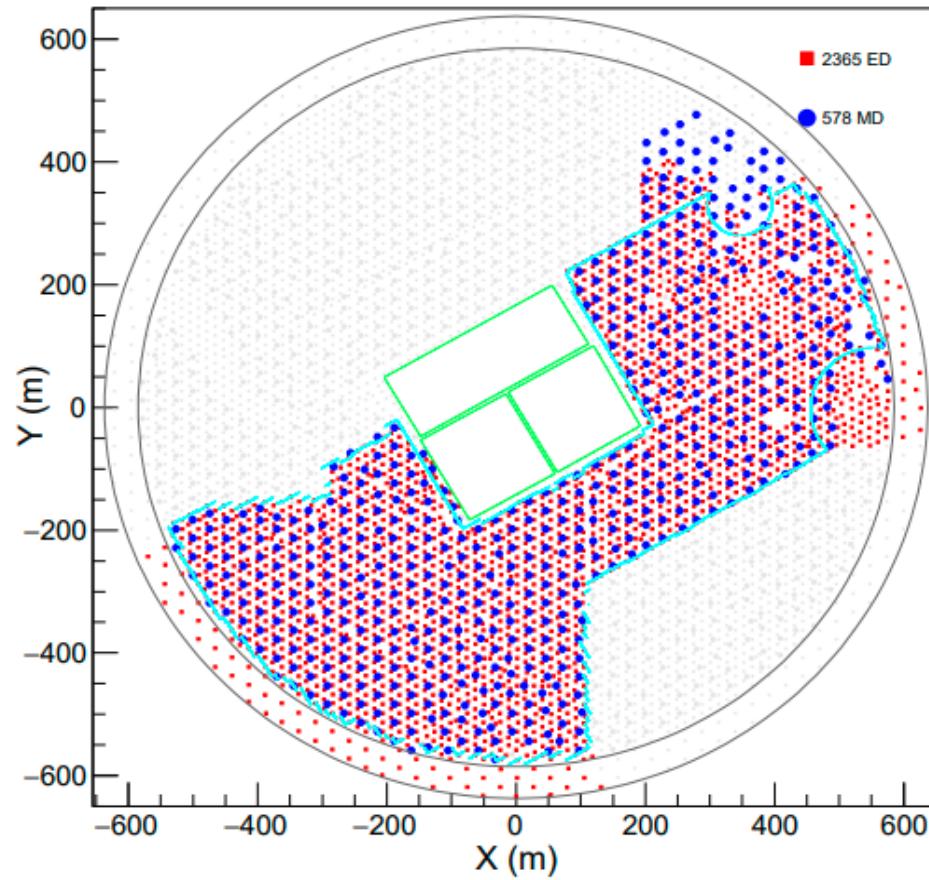
Wide Field Cherenkov Telescope Array (WFCTA):

18 wide-field-of-view Cherenkov/fluorescence telescopes



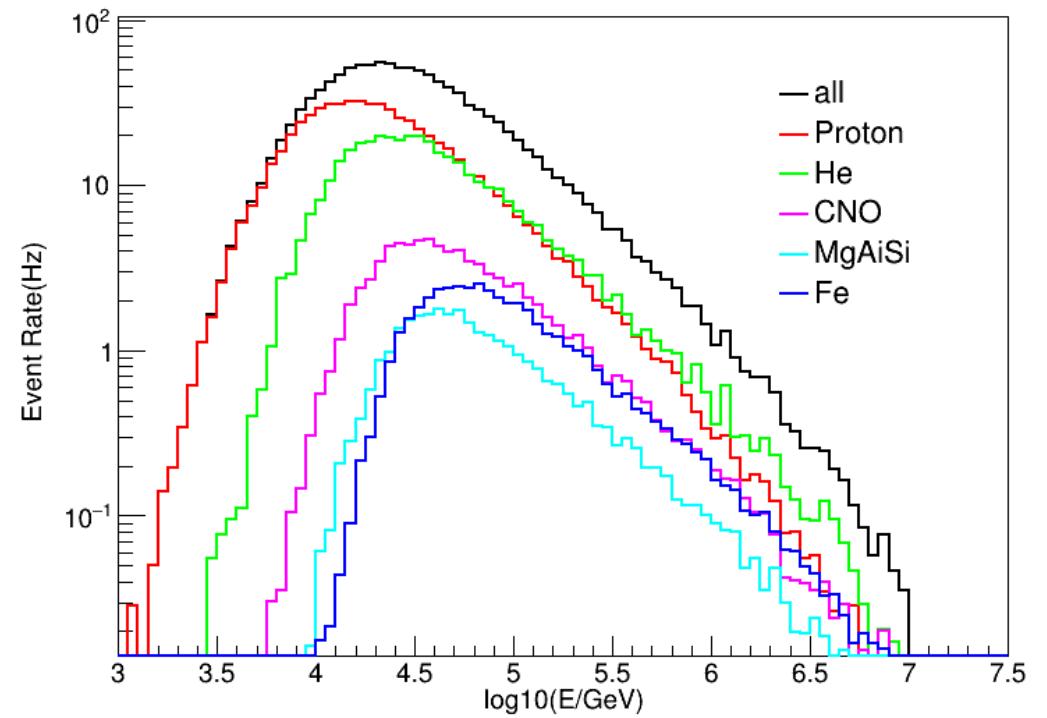
Data selection

- Half array of KM2A
- more than 20 EDs triggered
- Zenith < 50 degree
- Core within the array
- 2020001-2020366



Simulation

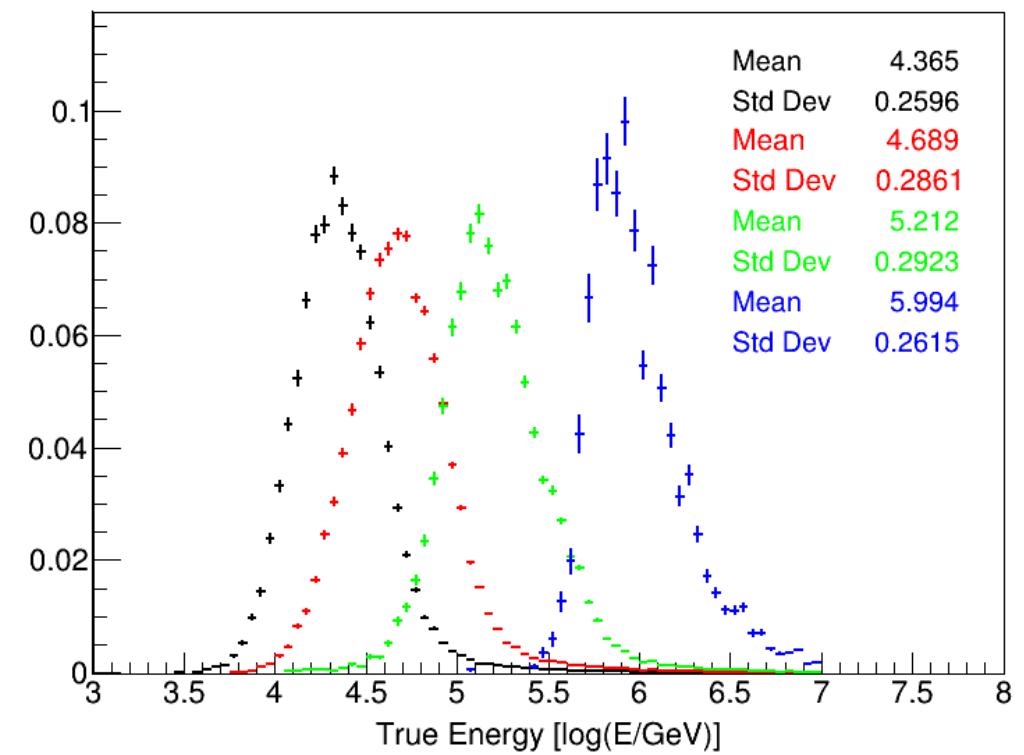
- QGSJETII-04_Gheisha
- 1TeV-10PeV
- Zenith angle 0-70°
- $\sim 7.8 \times 10^7$ events
- Gaisser model
- G4KM2A for half array



Energy estimation

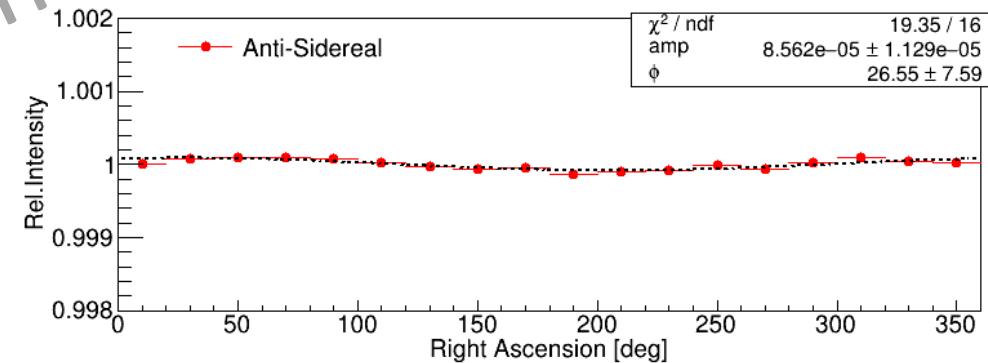
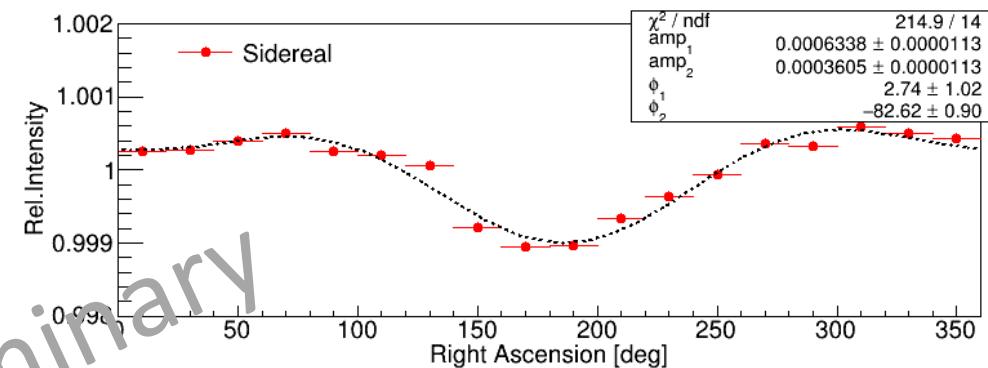
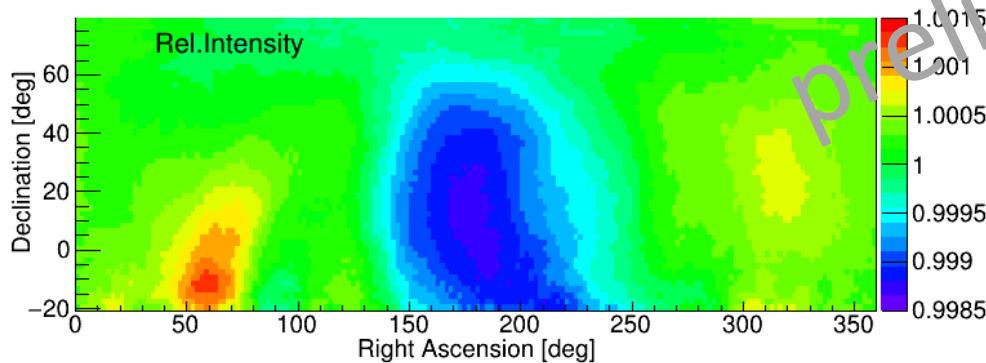
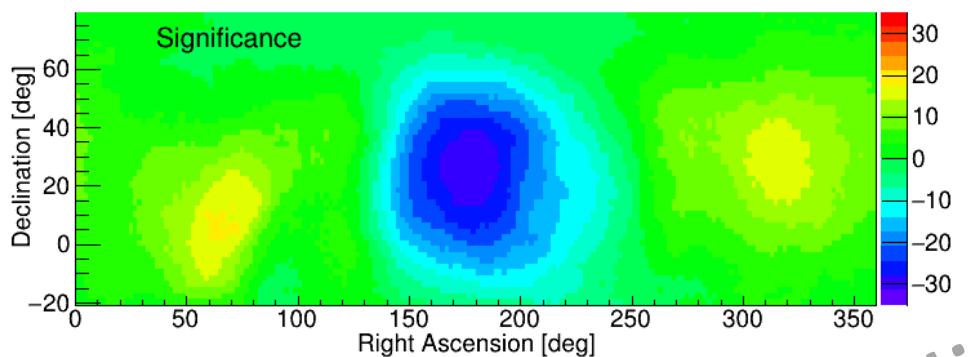
- $\log E = f(\log \sqrt{NpE * NuM})$

Bin	log(Erec)	Energy (TeV)	Number (10^8)
1	3.50~4.5	23	64.67
2	4.5~5.00	49	73.99
3	5.00~5.75	163	16.65
4	>=5.75	985	0.86

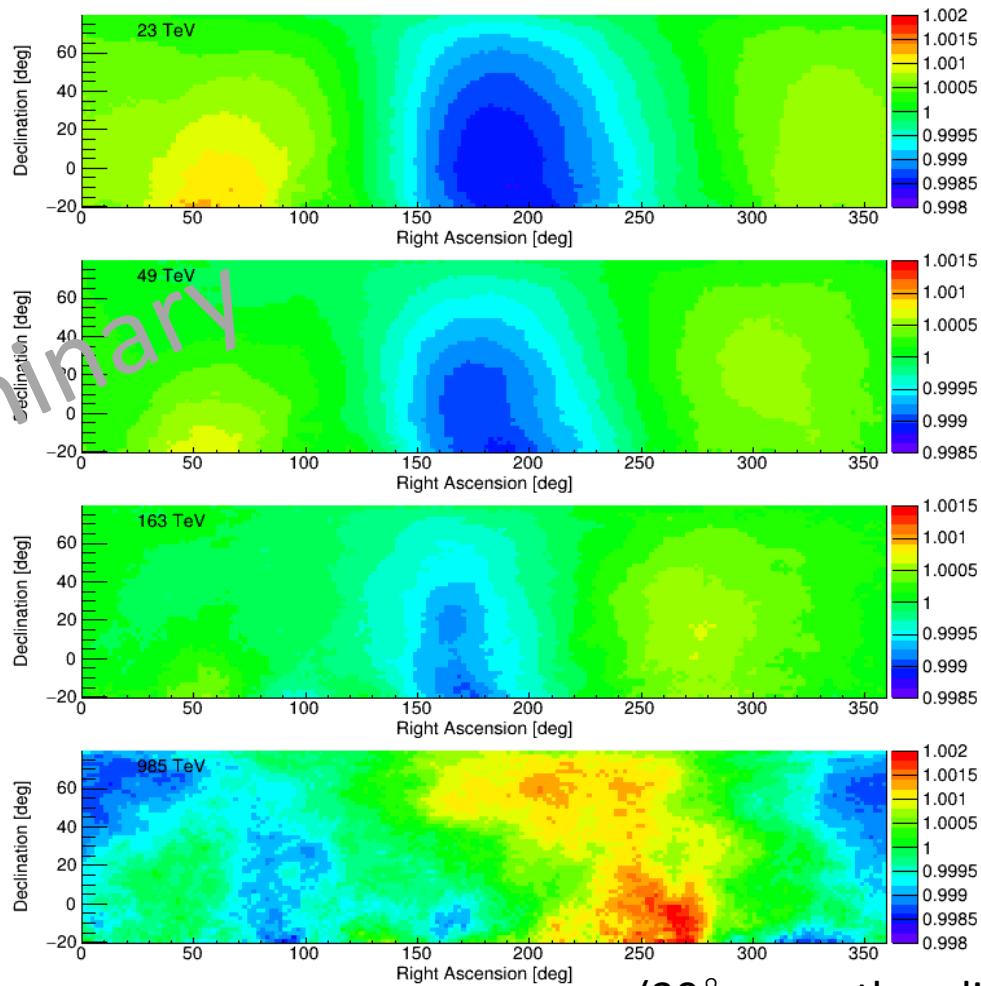
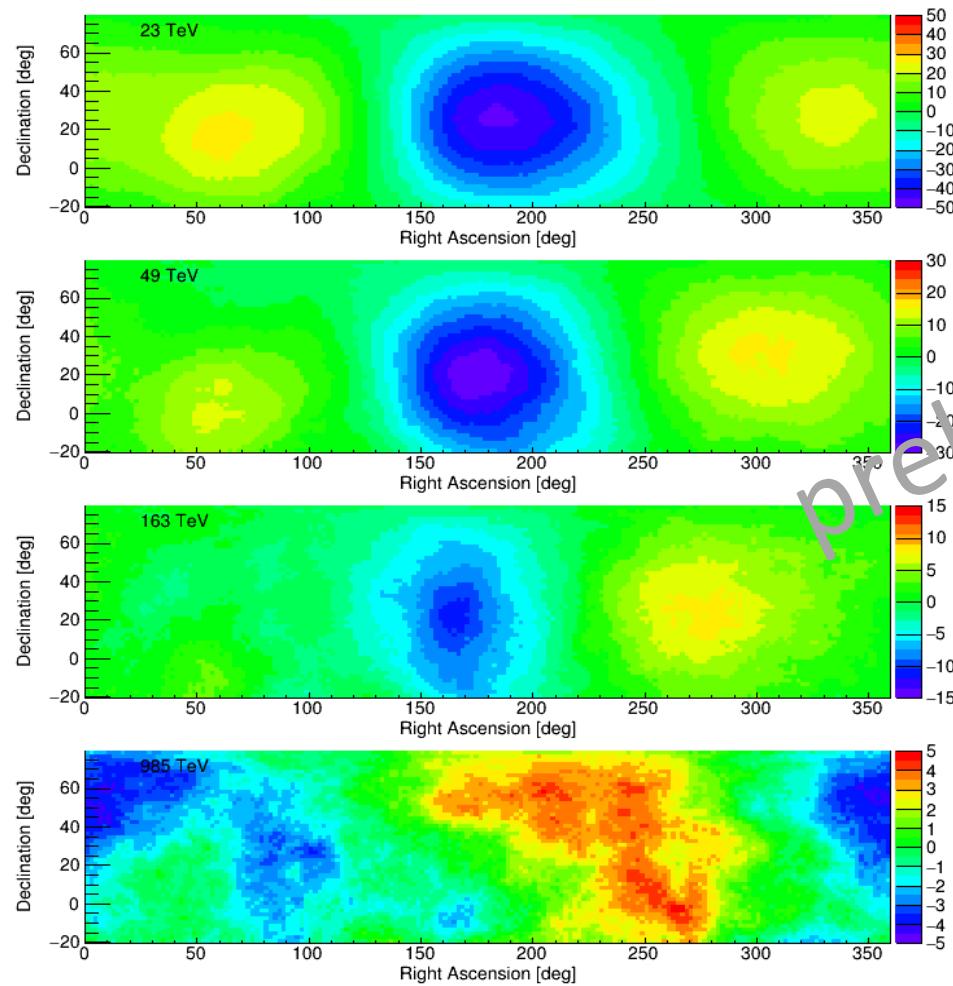


Anisotropy with LHAASO-KM2A

- $\sim 42\text{TeV}, 1.56 \times 10^{10}$ events
- A spurious frame (anti-sidereal) is used to estimate the systematic error

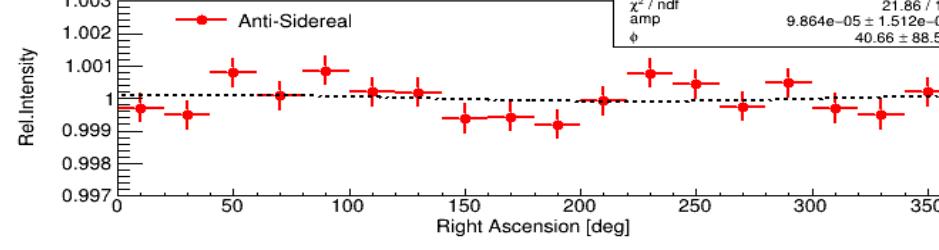
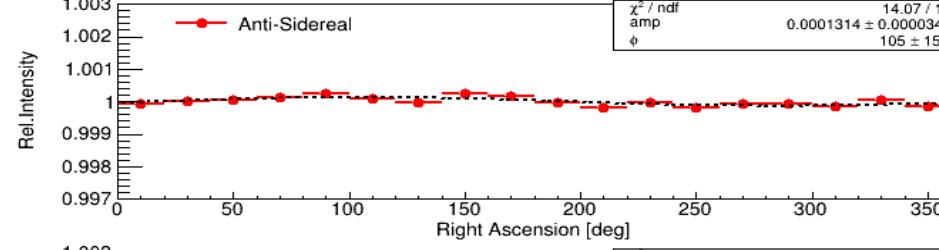
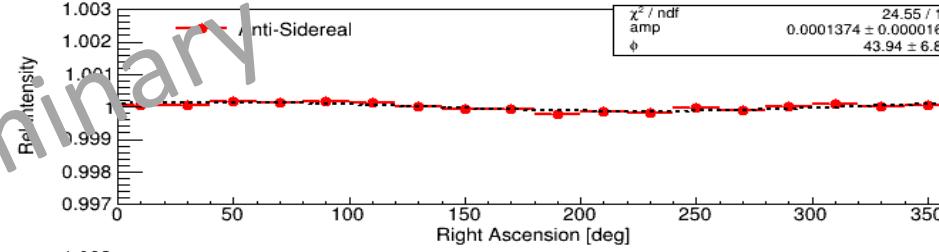
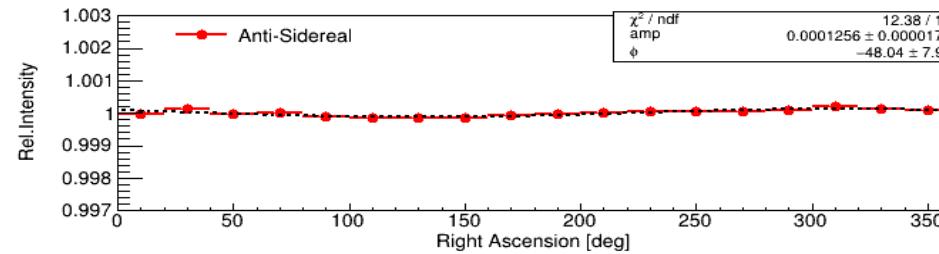
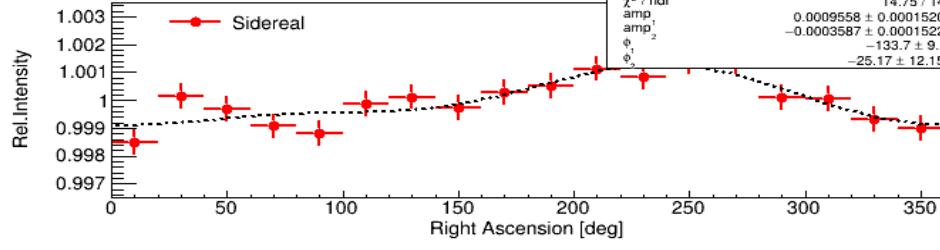
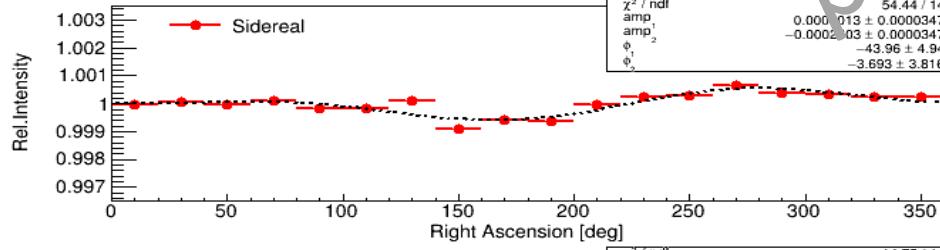
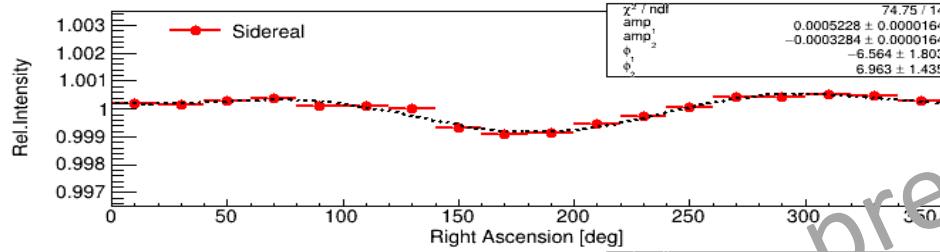
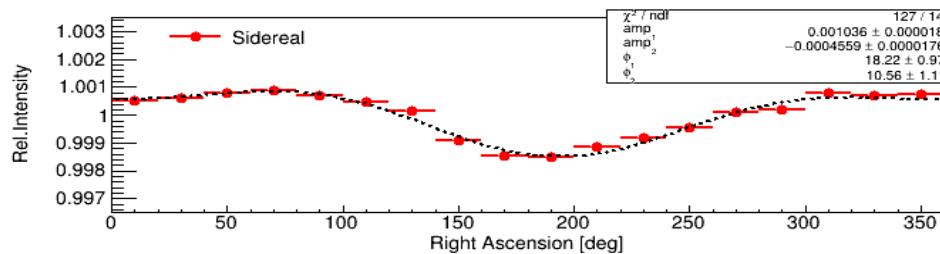


Energy dependence

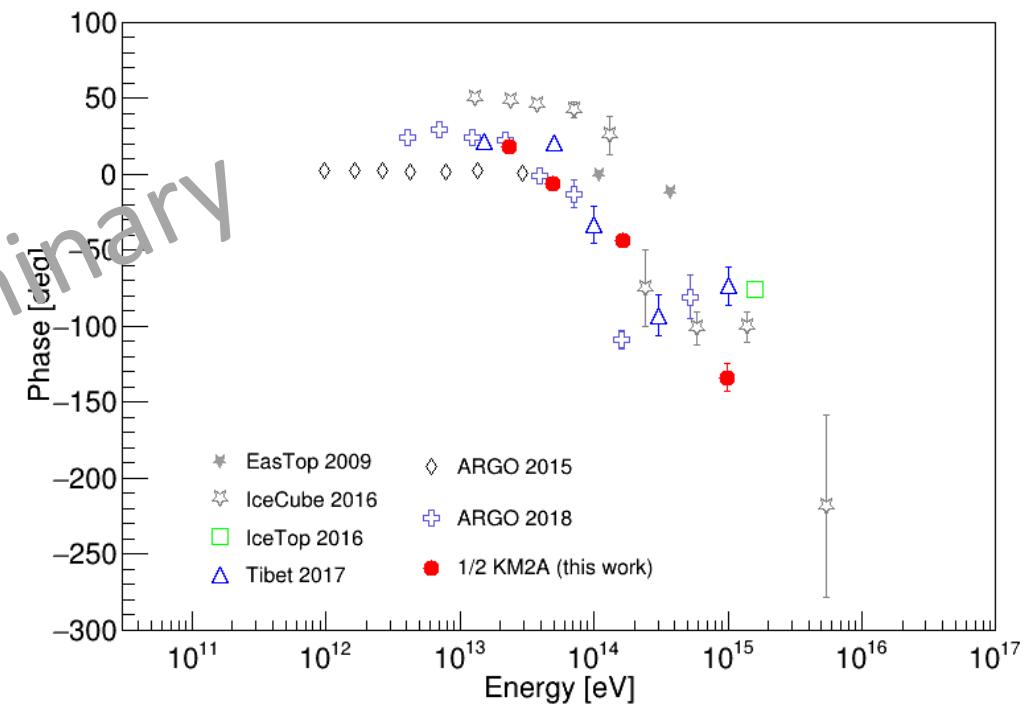
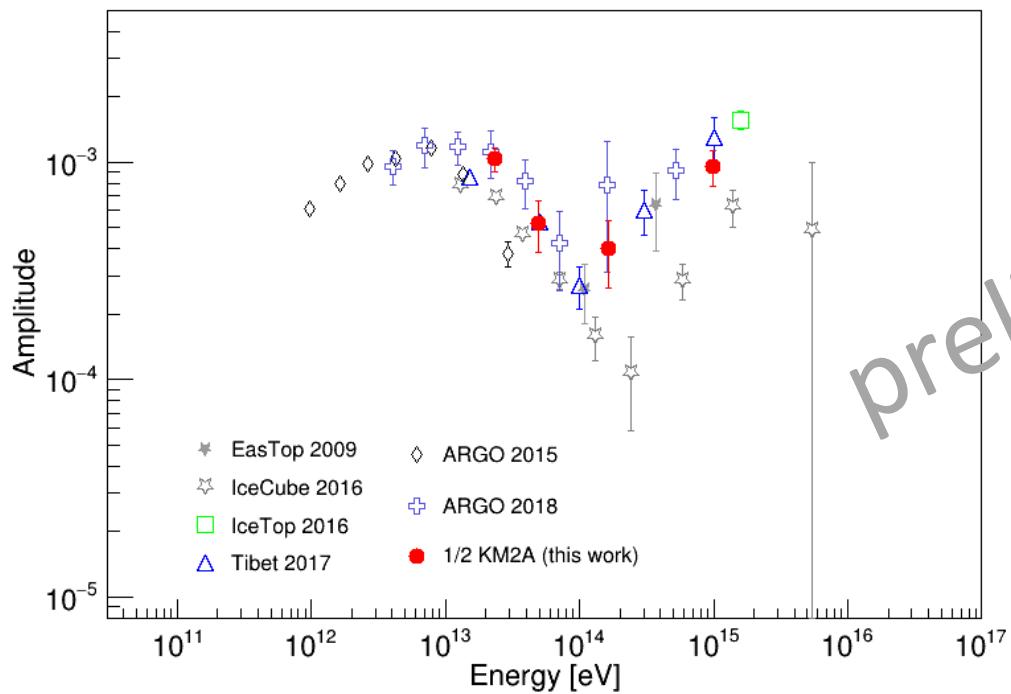


(30° smooth radius)

Energy dependence



Compared with others



Conclusion

- The half array of LHAASO-KM2A observed the large-scale anisotropy from 23 TeV up to 985 TeV.
- An inverse anisotropy is observed with a significance of 5σ at 985 TeV.
- The evolution of large-scale anisotropy observed by KM2A is consistent with other experiments.