




# On the need for unbiassing azimuthal asymmetry in signals measured by surface detector arrays

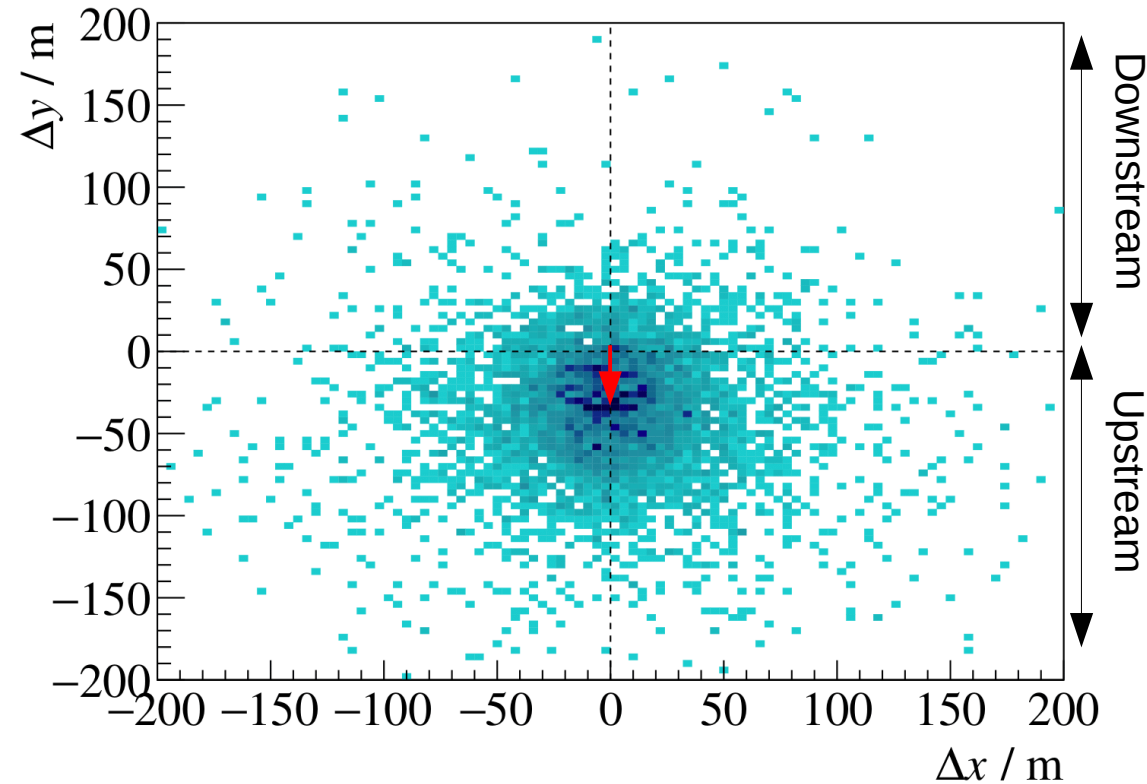
Quentin Luce, Markus Roth, David Schmidt, Darko Veberič  
Institute for Astroparticle Physics, Karlsruhe Institute of Technology , Germany

# Azimuthal asymmetry in signal

Azimuthal asymmetry in signals **observed in the water-Cherenkov detector** of the surface detector is combination of:

- **geometrical effects** from the inclination of the shower
- **attenuation of the charged particles** from their point of emission to the ground

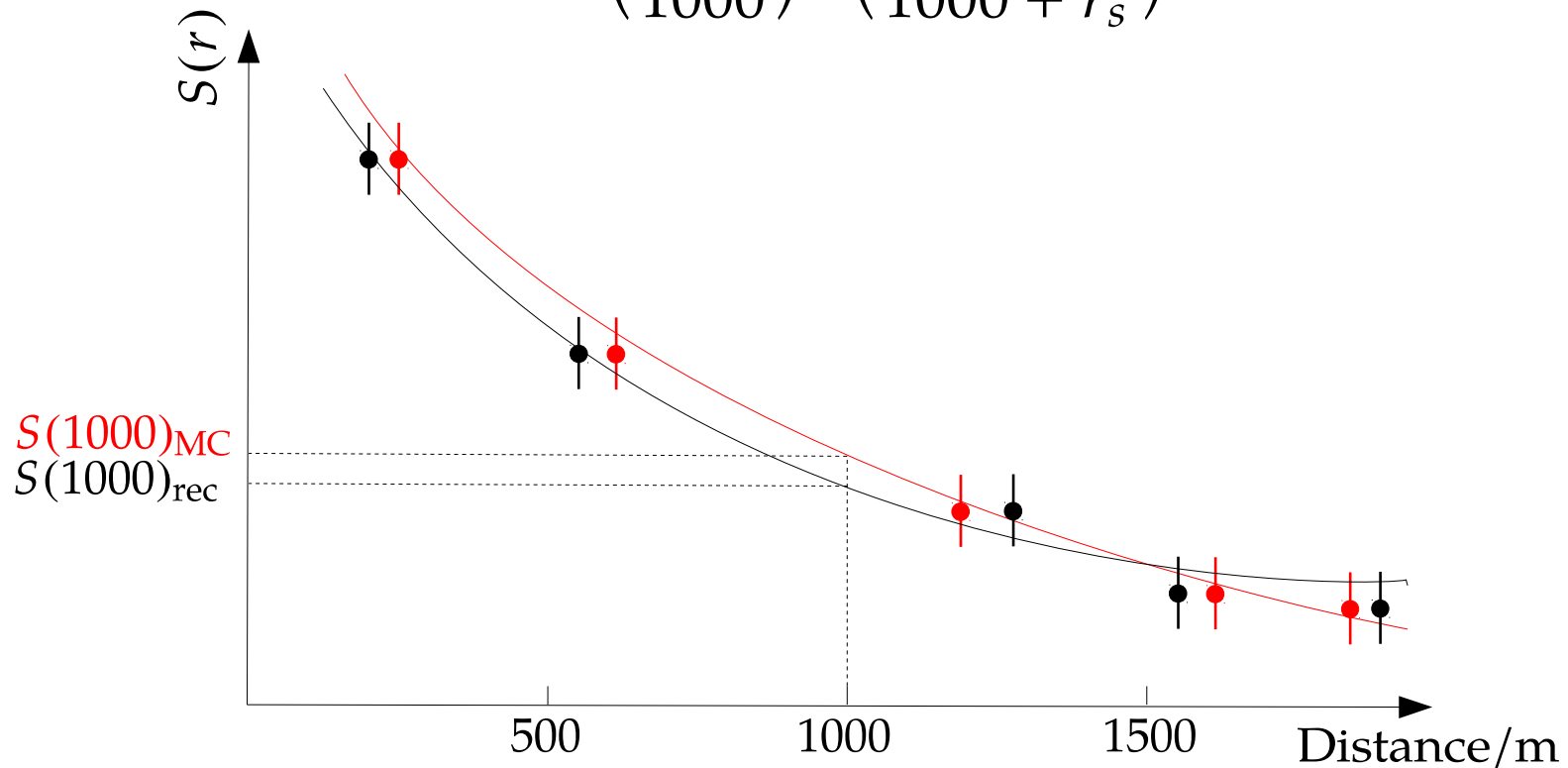
From the **surface detector of the Pierre Auger Observatory**:  
introduction of a **bias (~40 m) in the reconstruction of the core** of the shower when **axisymmetric LDF** is used



# Azimuthal asymmetry in signal

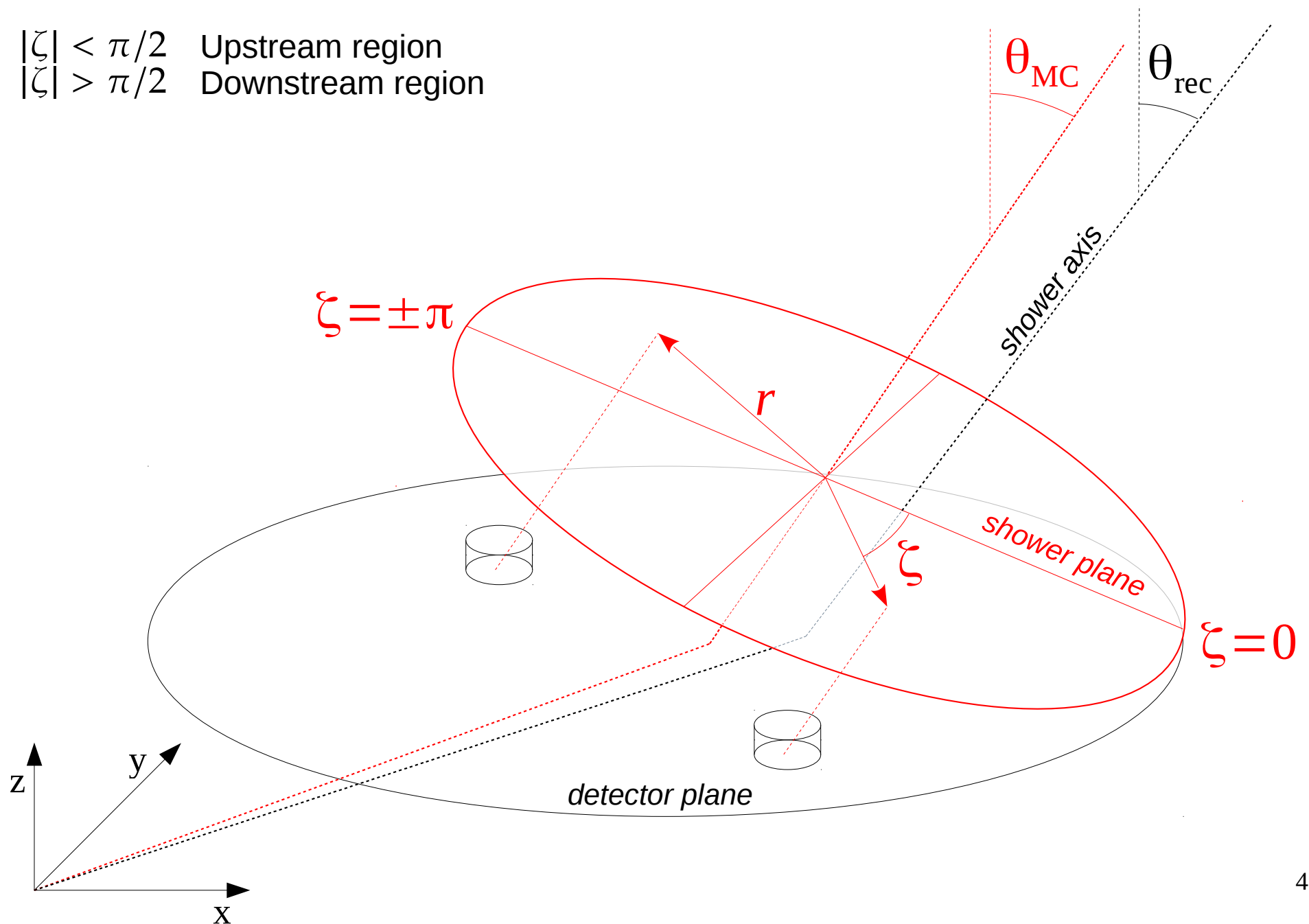
$$S(r) = S(1000) f_{\text{LDF}}(r)$$

$$f_{\text{LDF}}(r) = \left(\frac{r}{1000}\right)^{\beta} \left(\frac{r + r_s}{1000 + r_s}\right)^{\beta + \gamma}$$



# Azimuthal asymmetry in signal

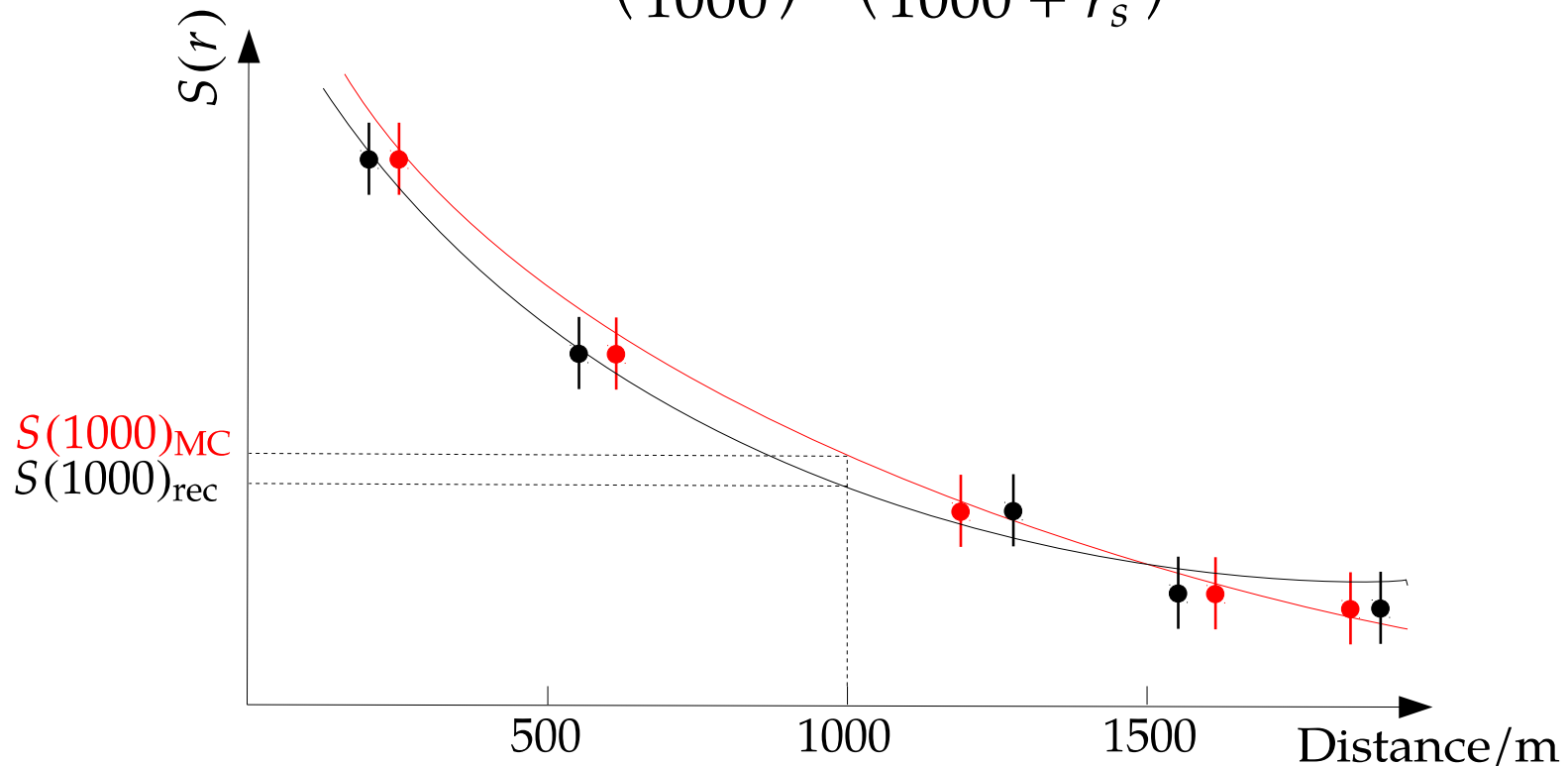
$|\zeta| < \pi/2$  Upstream region  
 $|\zeta| > \pi/2$  Downstream region



# Azimuthal asymmetry in signal

$$S(r, \zeta) = S(1000) f_{LDF}(r) [1 + \alpha(r, \theta, S(1000)) \cos \zeta]$$

$$f_{LDF}(r) = \left(\frac{r}{1000}\right)^\beta \left(\frac{r + r_s}{1000 + r_s}\right)^{\beta+\gamma}$$

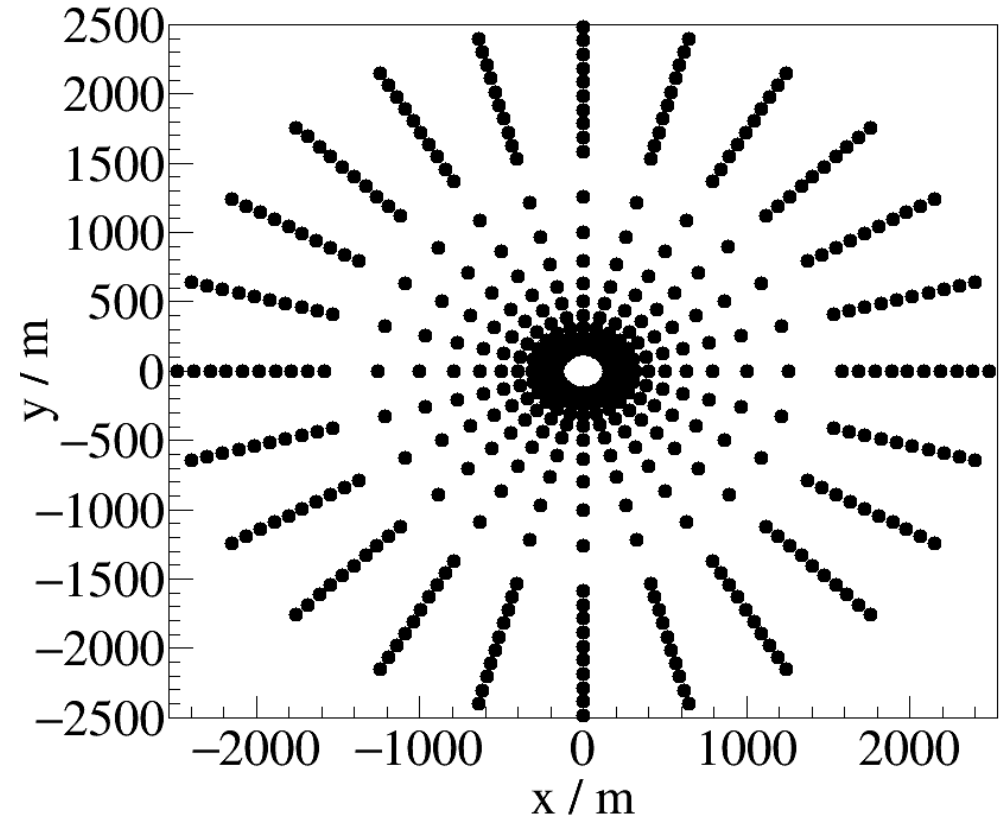


Parametrisation of  $\alpha(r, \theta, S(1000))$  using simulations

# Simulated data-sets

---

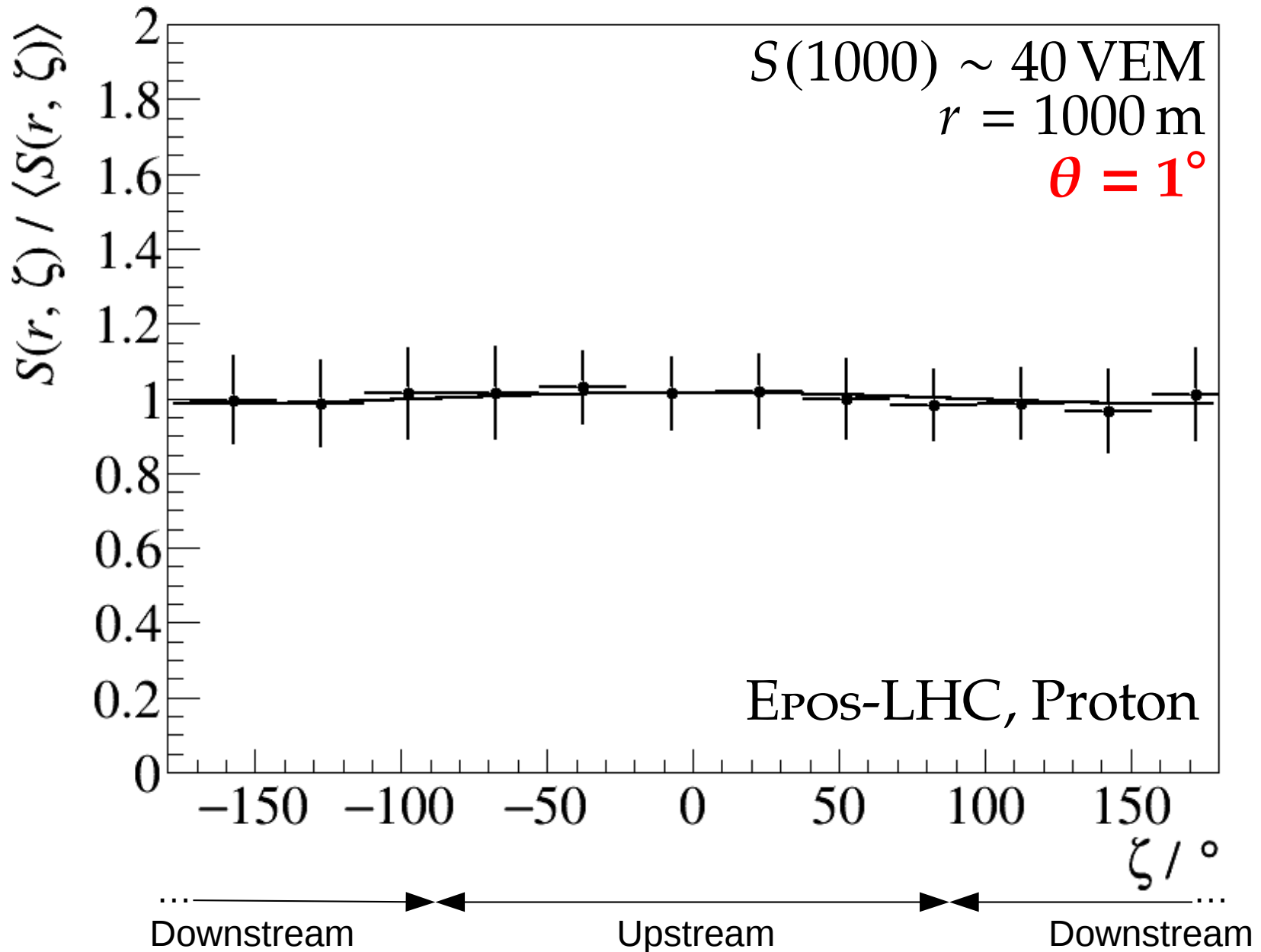
- hadronic model :  
**QGSJet-II.04, EPOS-LHC**
- primary : **Proton**, **Iron**
- Continuous library  
 $\lg(E / \text{eV}) = 18.5 - 20.0$   
 $\theta / ^\circ = 0 - 60$  (flat distribution in  $\sin^2\theta$ )
- surface detector of the **Pierre Auger Observatory**
- **20 dense rings** = 24 detectors at a fixed distance from the shower axis
- Offline simulation and reconstruction frameworks<sup>1,2</sup>



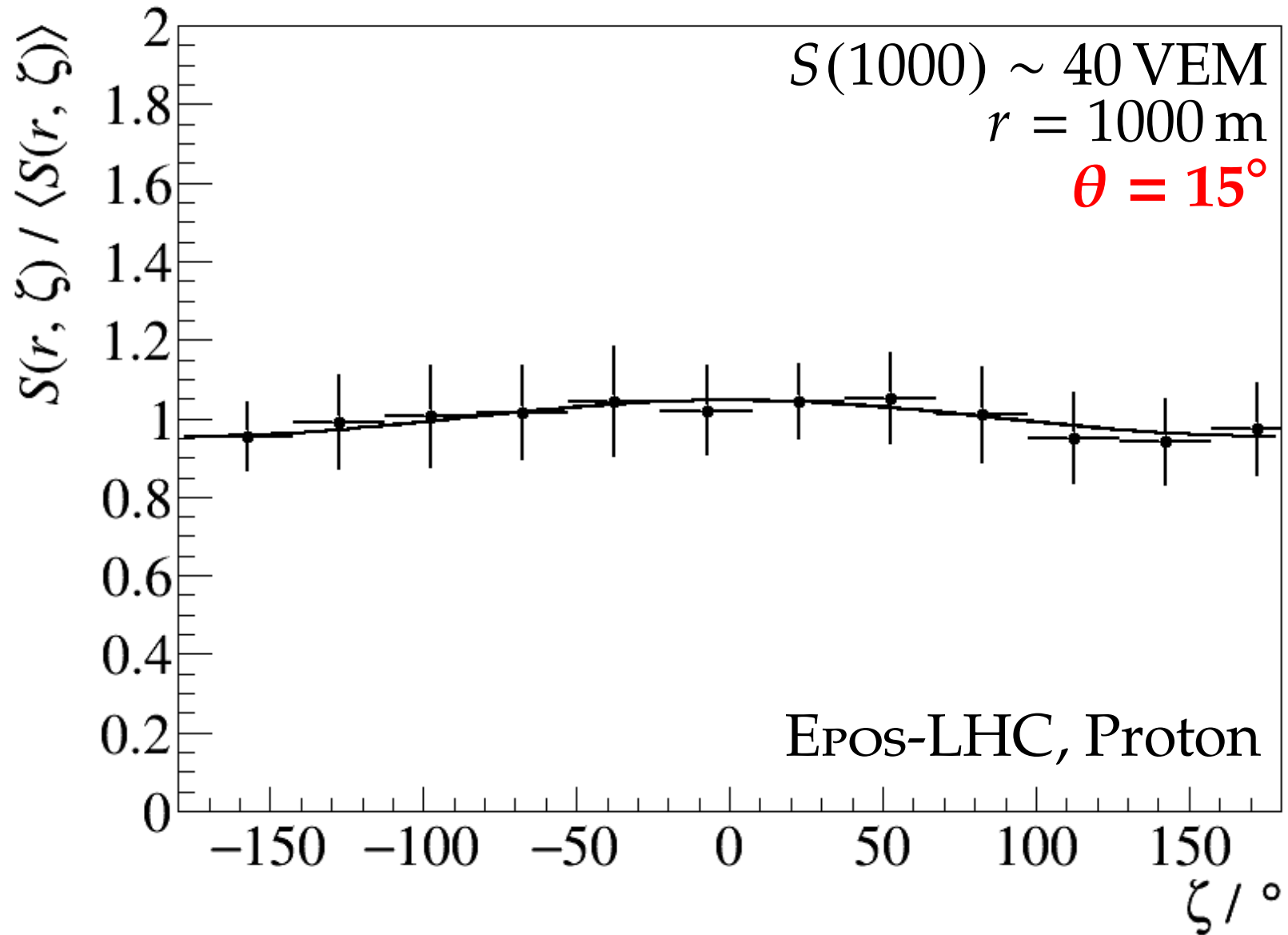
<sup>1</sup> S. Argirò et al., *Nucl. Instrum. Meth. A* **580** (2007) 1485-1496, [0707.1652]

<sup>2</sup> A. Aab et al. [Pierre Auger Coll.], *JINST* **15** (2020) P10021, [2007.09035]

# Amplitude of the asymmetry

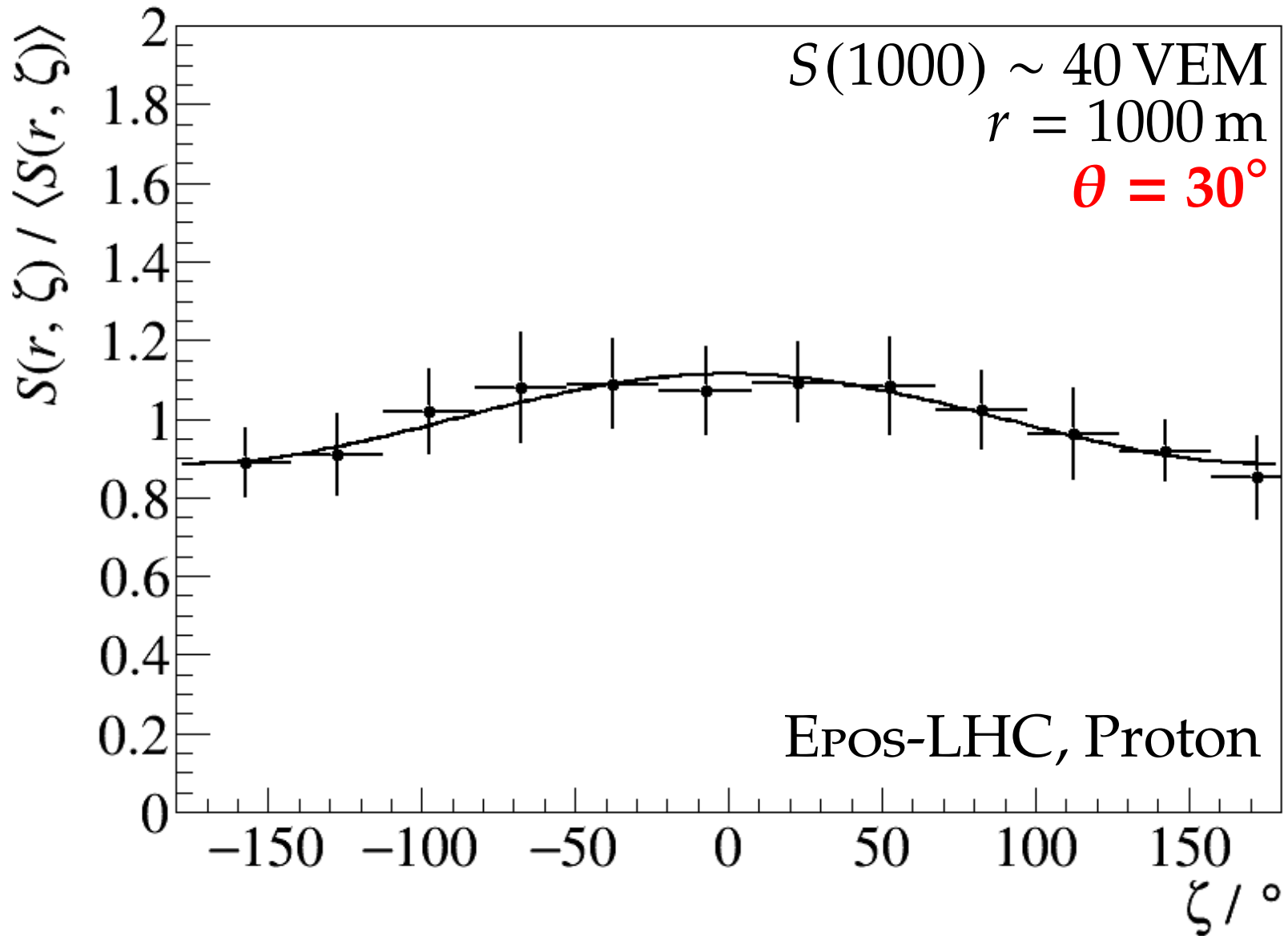


# Amplitude of the asymmetry

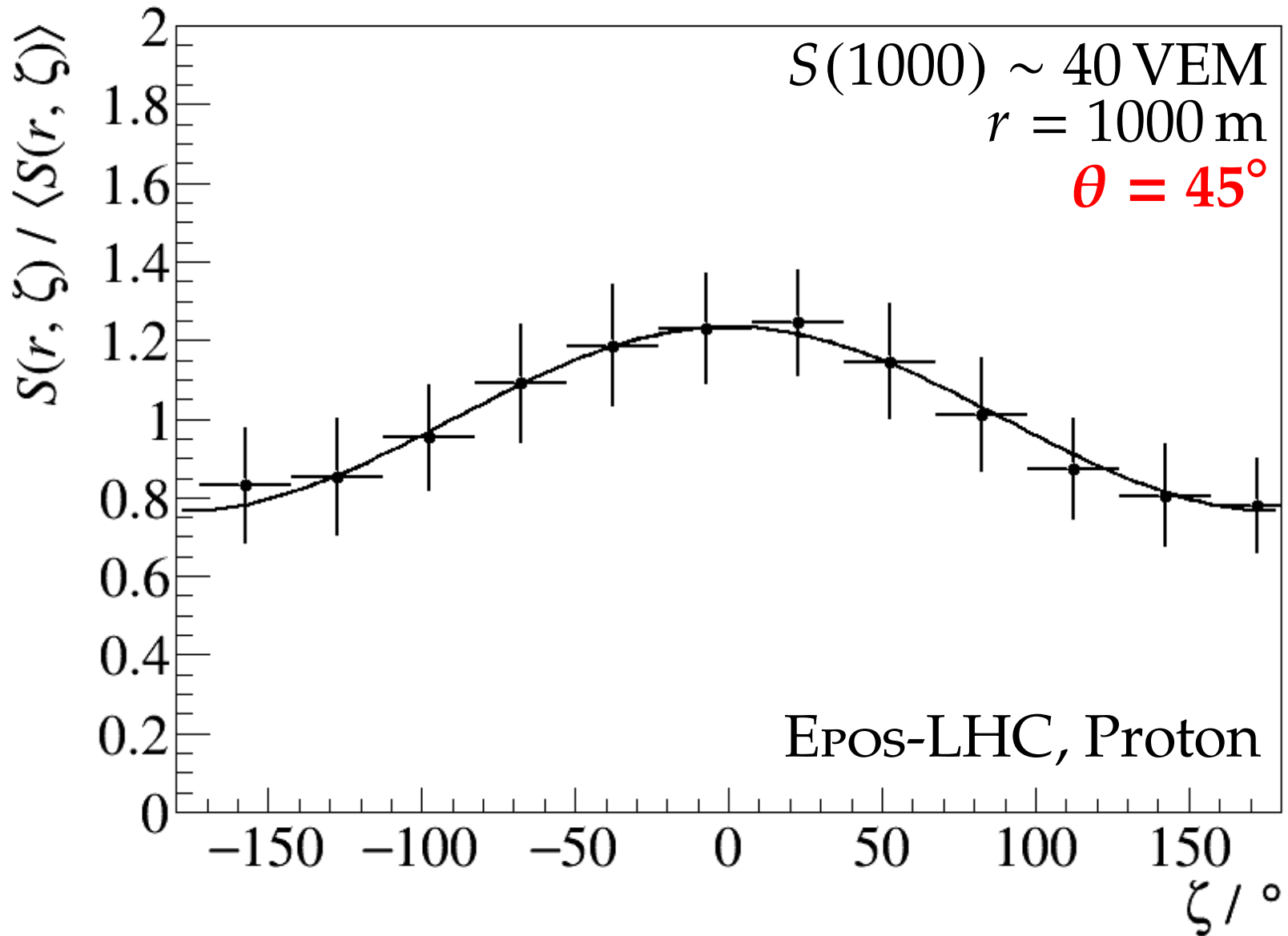




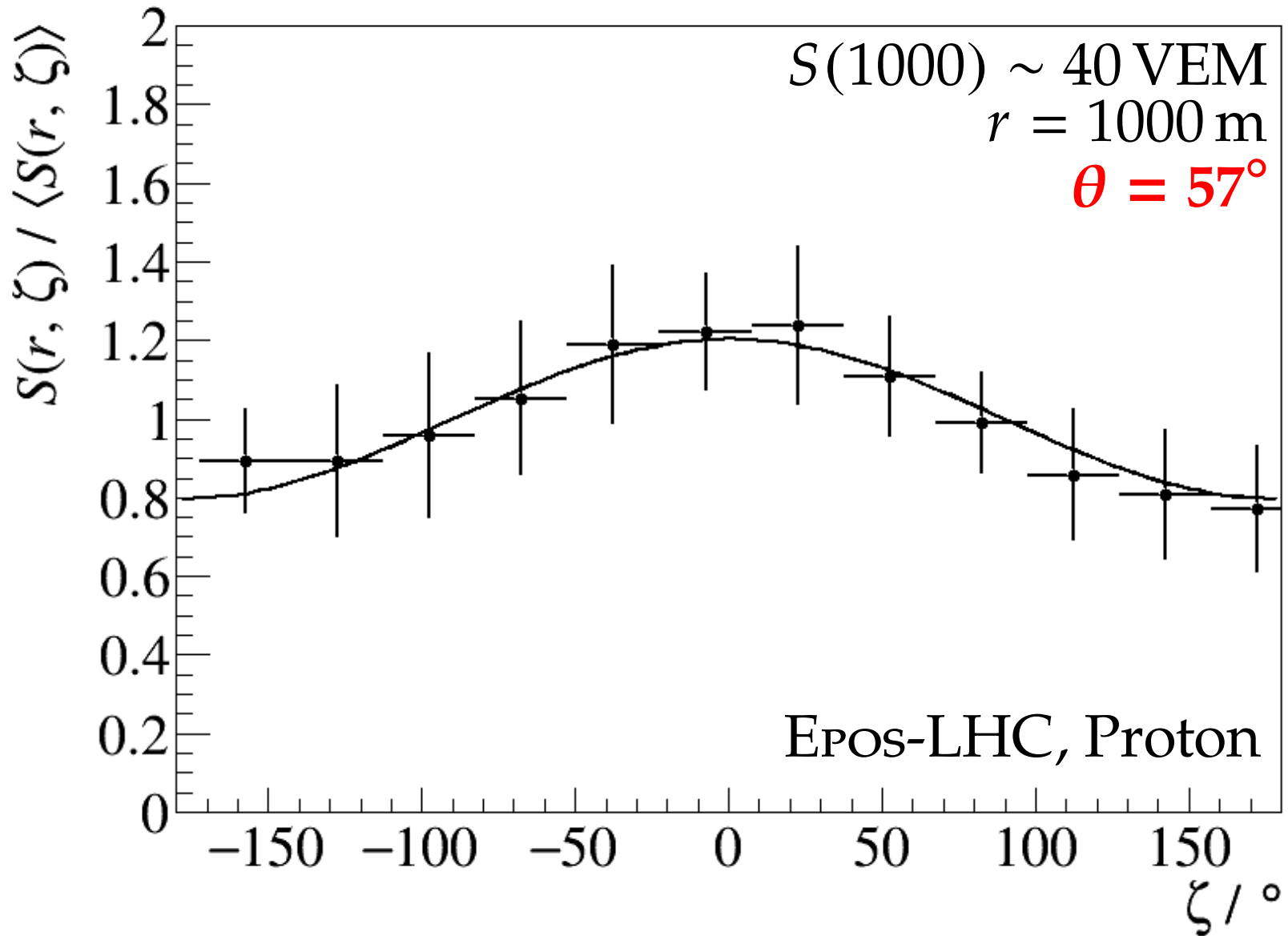
# Amplitude of the asymmetry



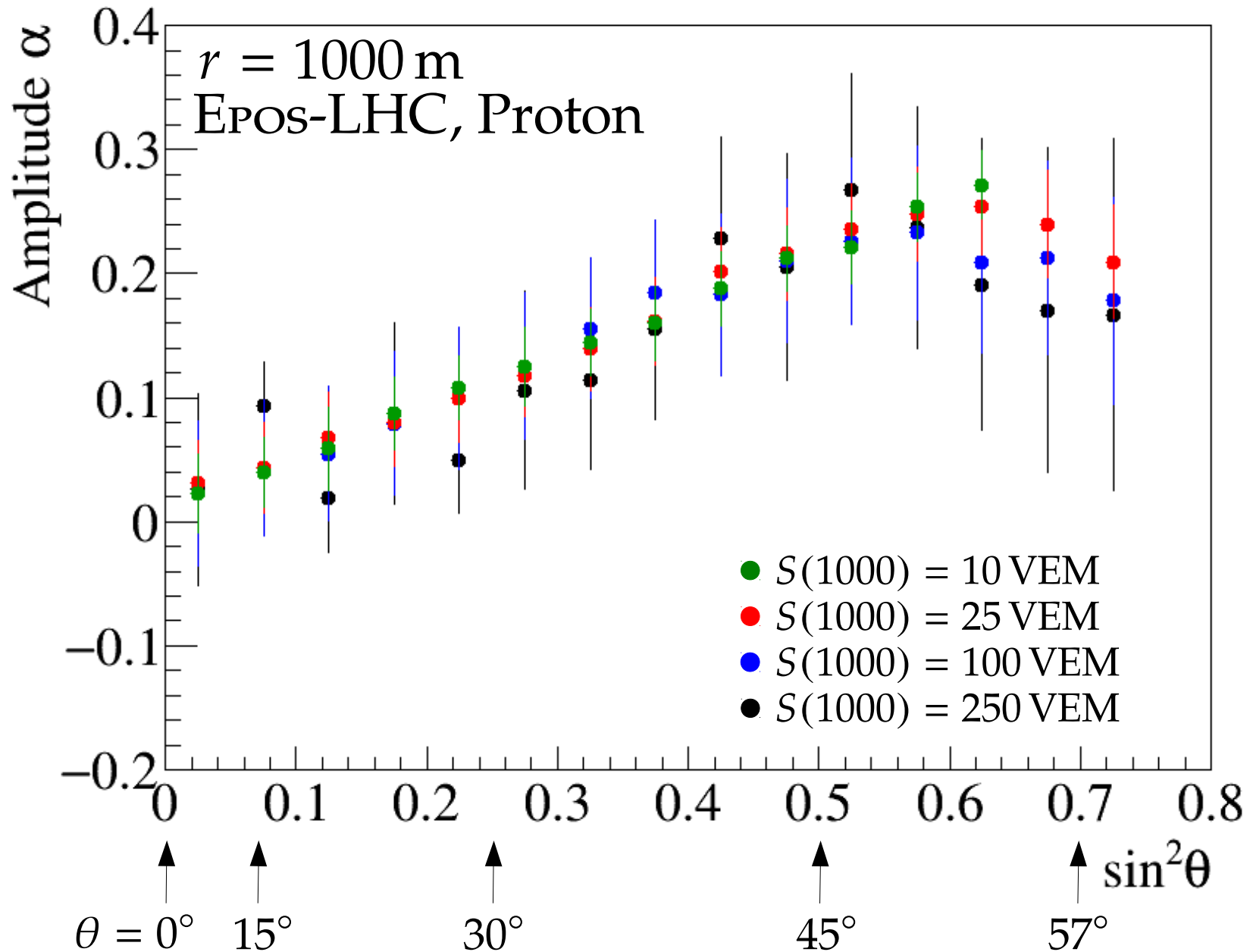
# Amplitude of the asymmetry



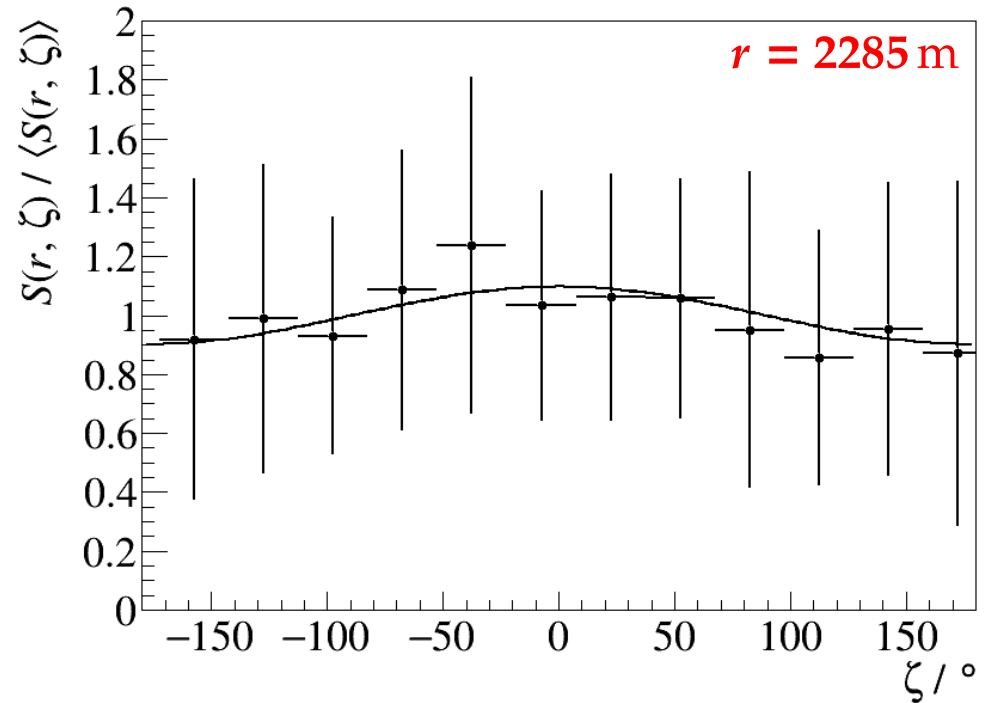
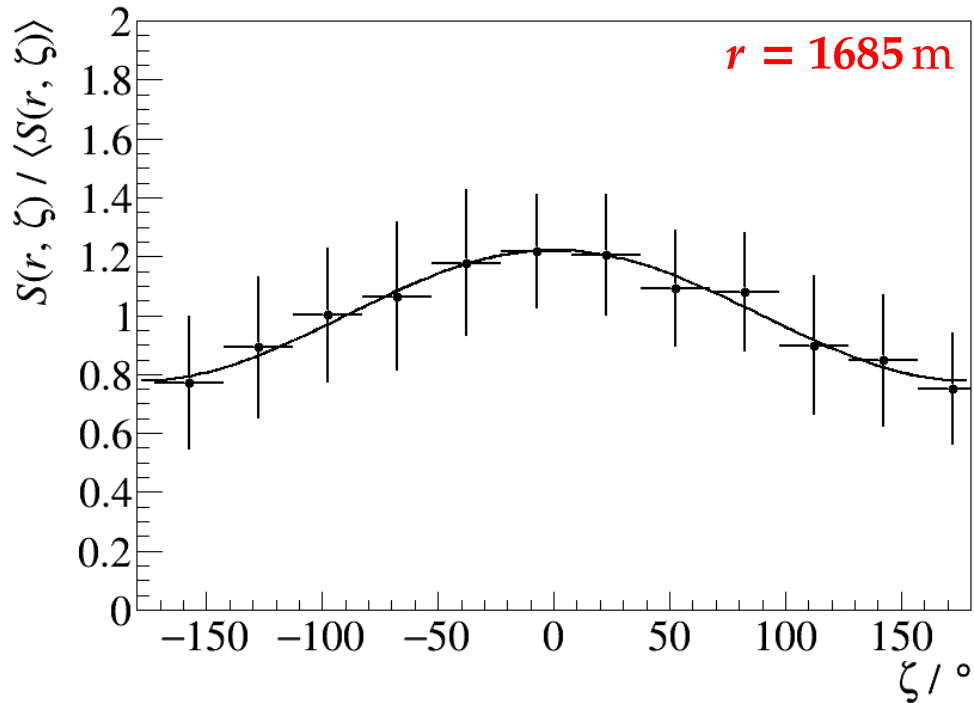
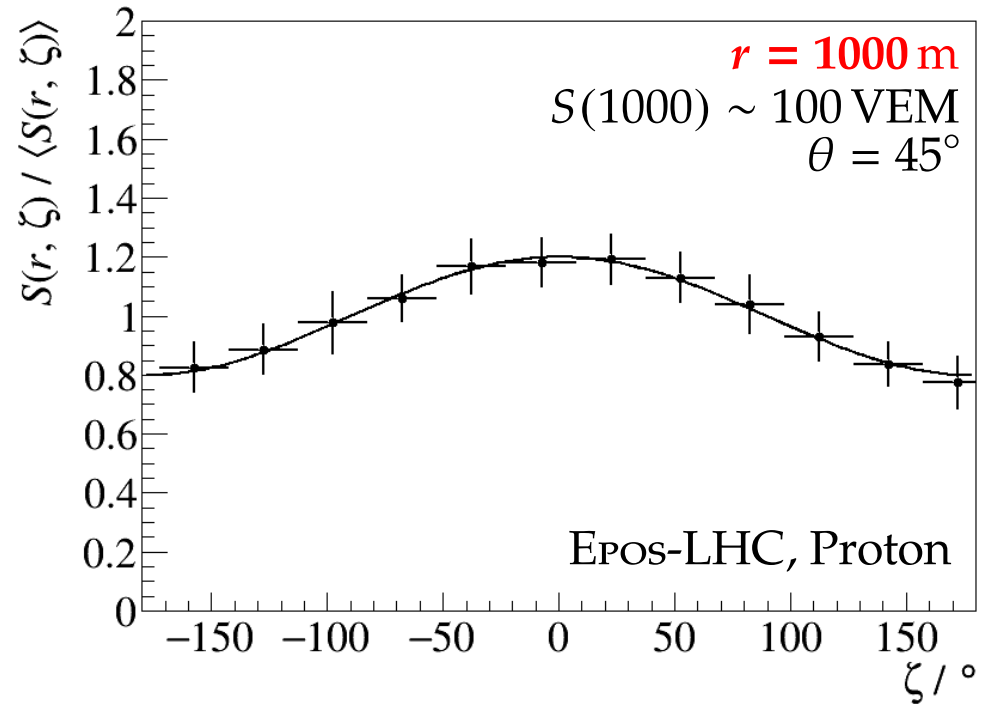
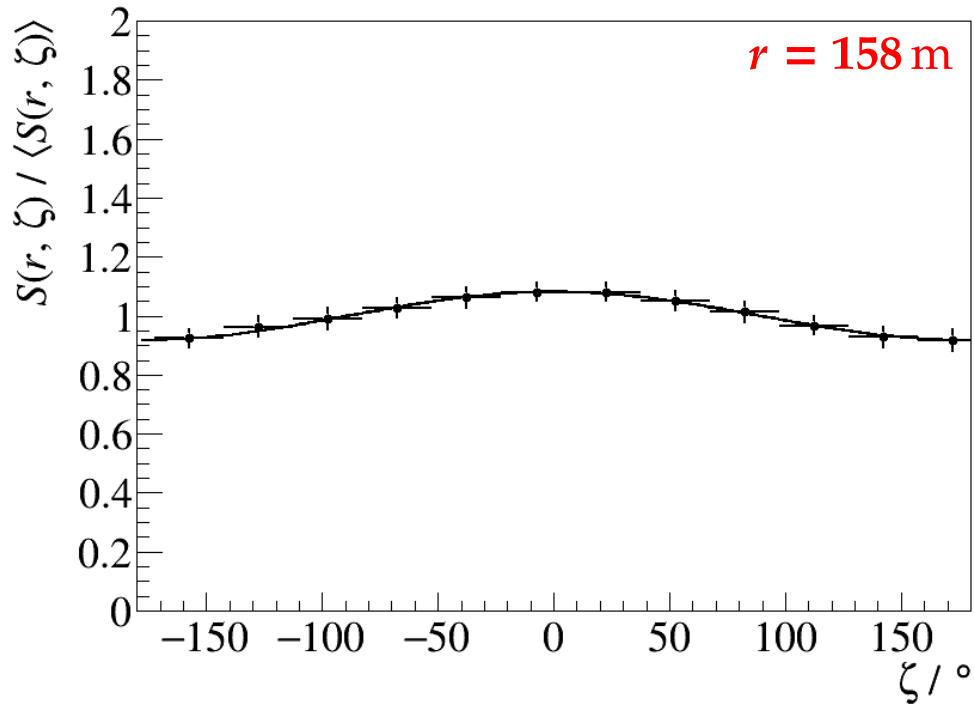
# Amplitude of the asymmetry



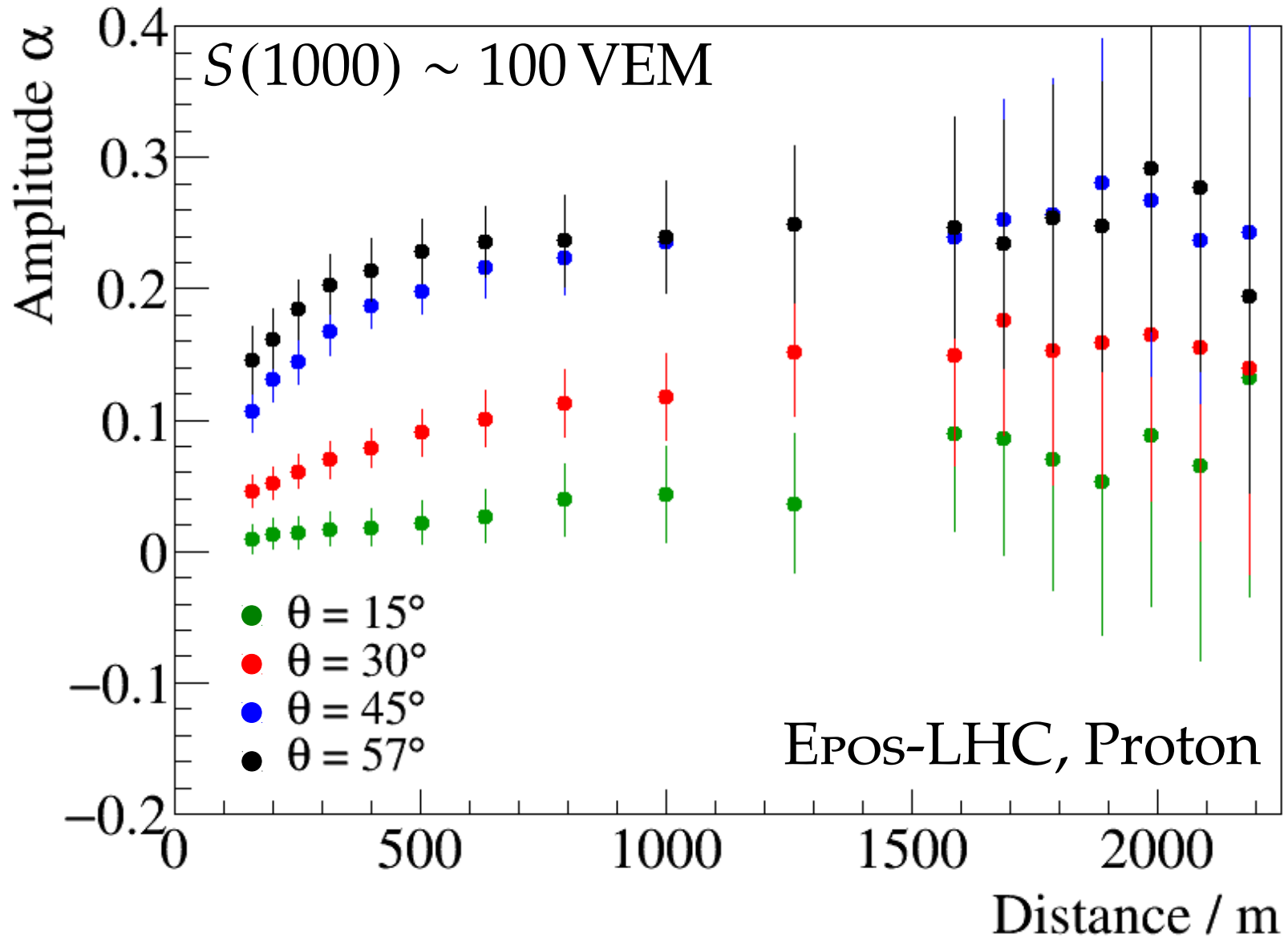
# Evolution with the inclination



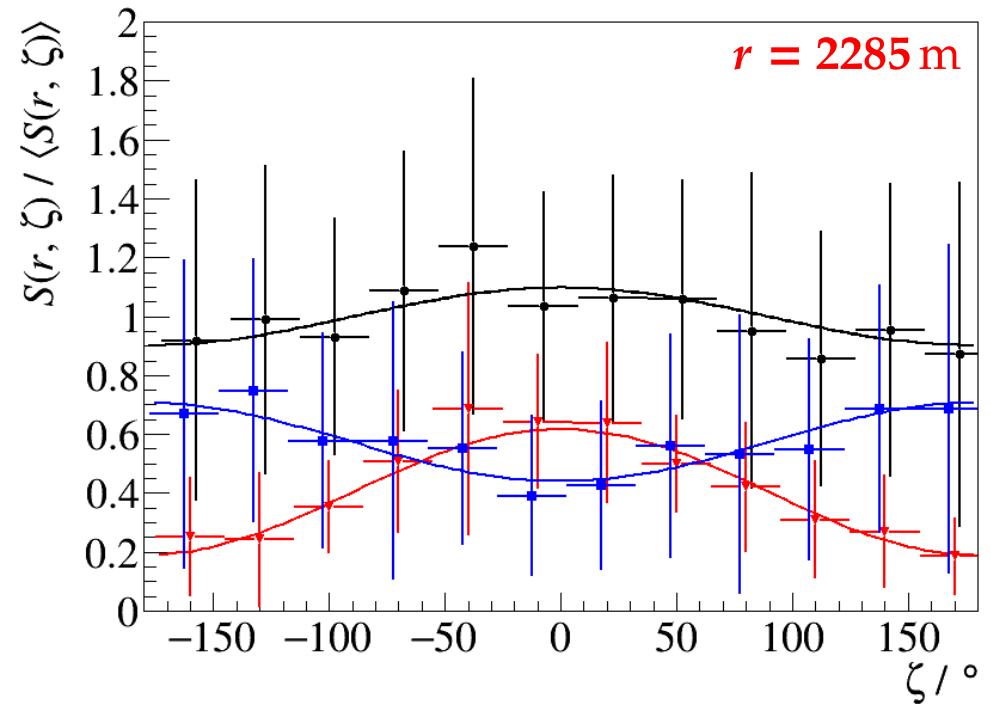
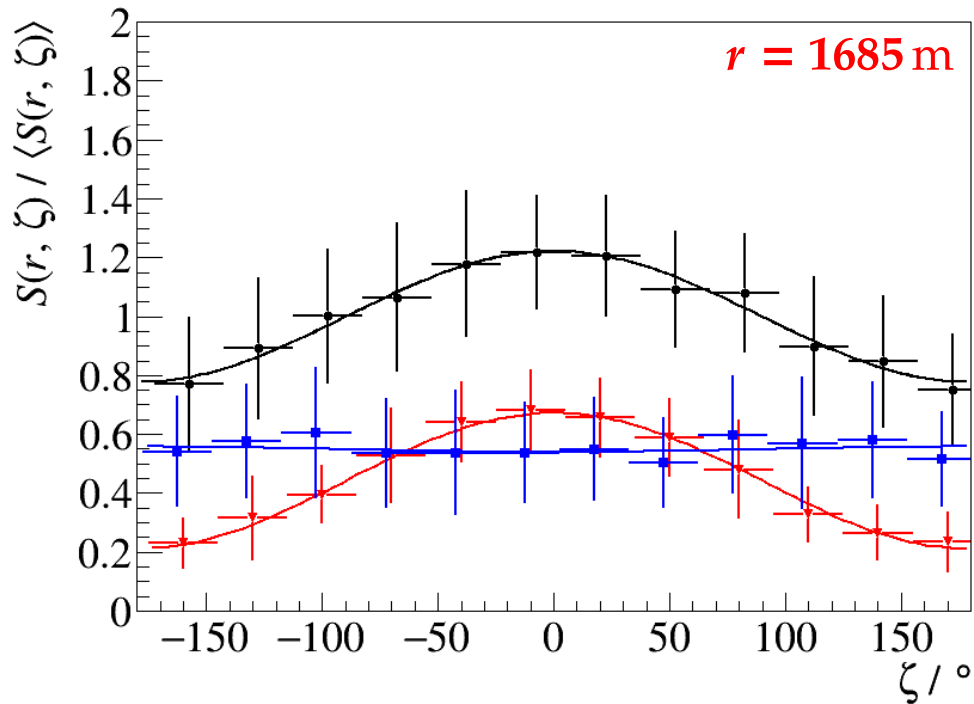
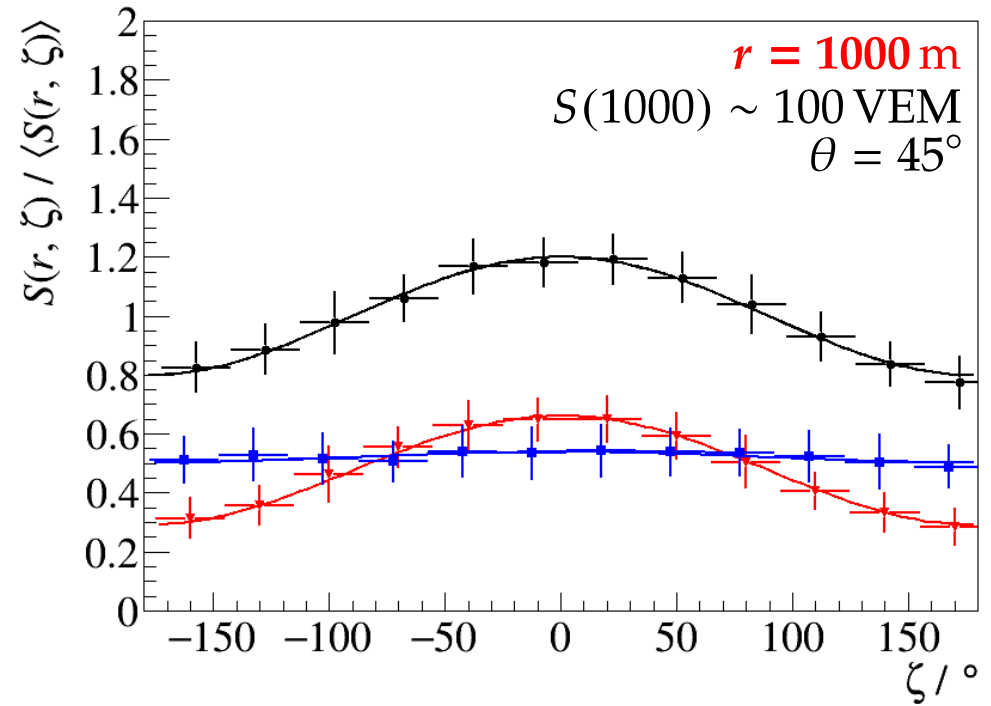
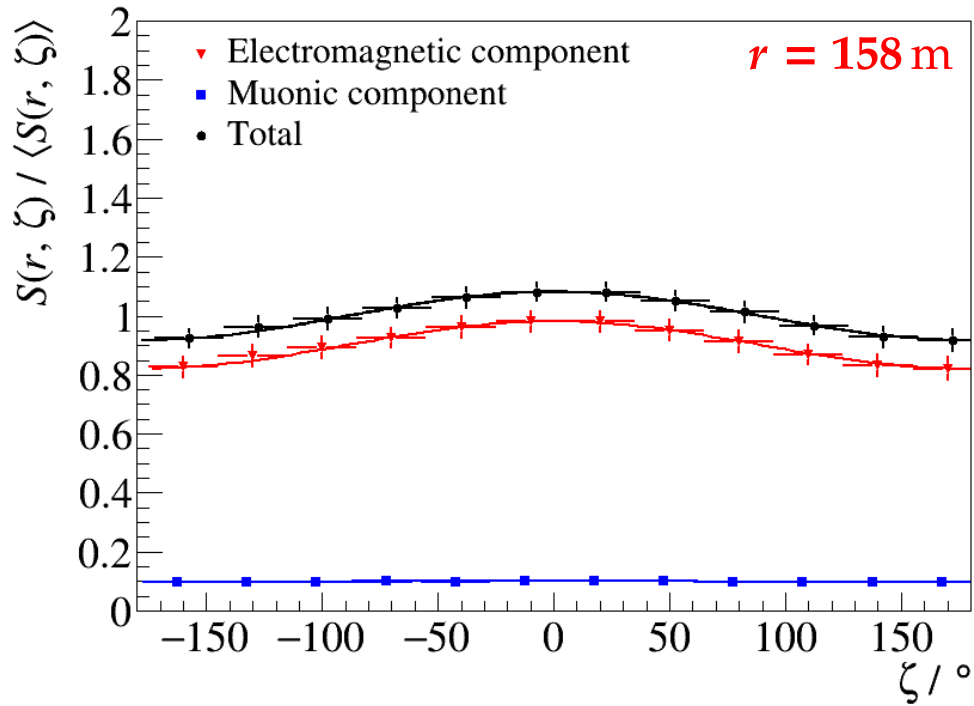
# Amplitude vs distance



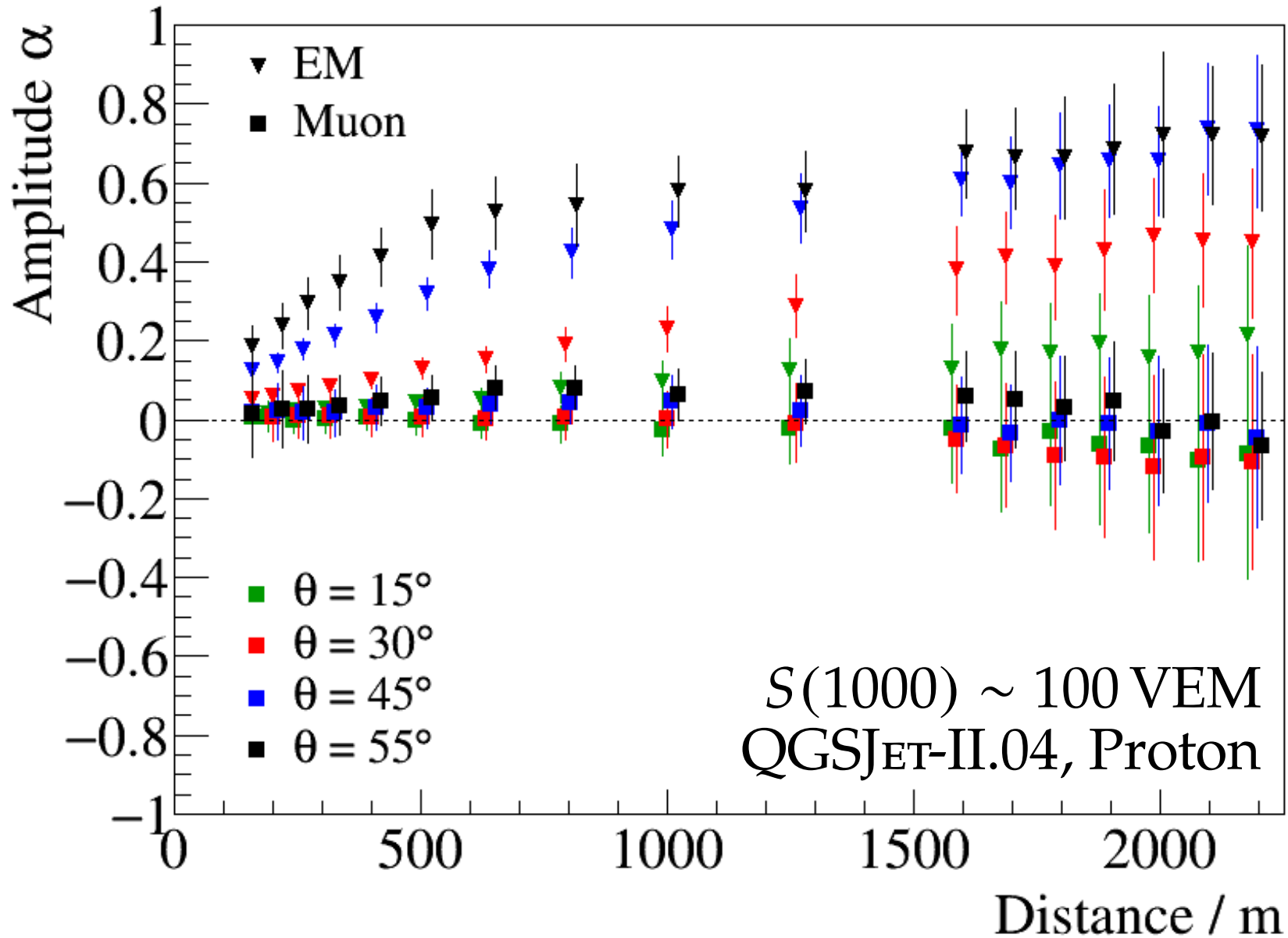
# Evolution with the distance



# Electromagnetic vs Muonic components



# Electromagnetic vs Muonic components





# Negative amplitude?

Amplitude of the asymmetry<sup>3</sup> = combination of:

- attenuation over the distance from the emission point to the ground

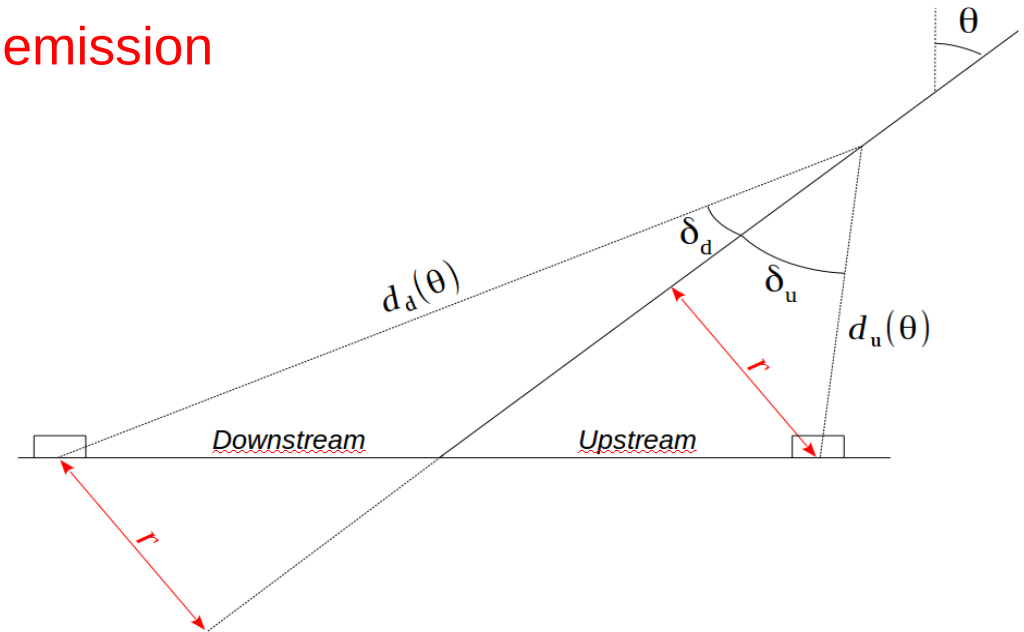
$$f_{\text{att}}(d(\theta)) = \exp(-d(\theta)/\lambda)$$

- dependence of particle density in a fixed solid angle

$$\Delta\Omega \propto 1/d^2$$

- angular distribution function of the emission of particles

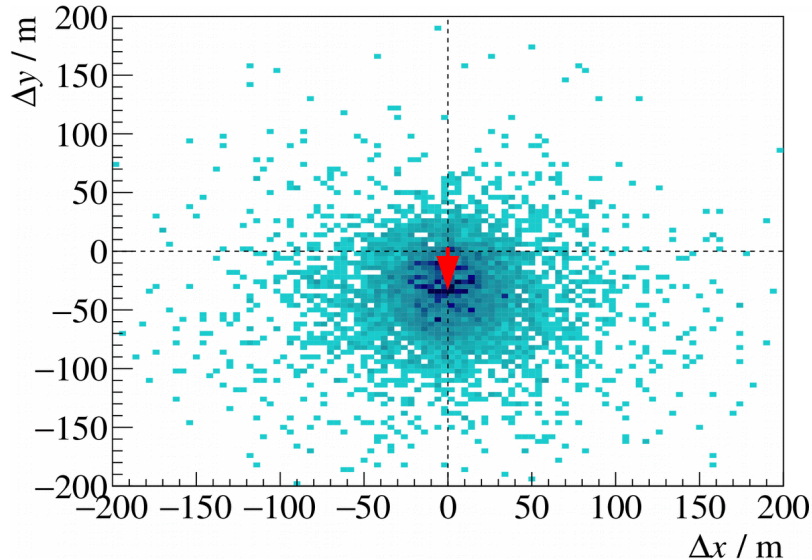
$$\text{ADF}(\delta) \propto (\delta/\delta_0)^{-\gamma}$$



$$\alpha \propto 2 + \frac{d(\theta)}{\lambda} - \gamma$$

# Impact on the core reconstruction

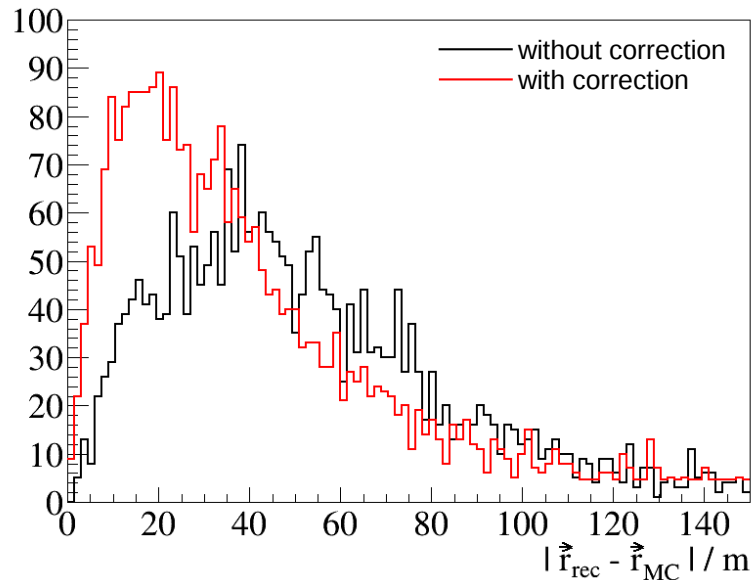
## Bias in the core position:



Core bias = mean value in the upstream-downstream direction

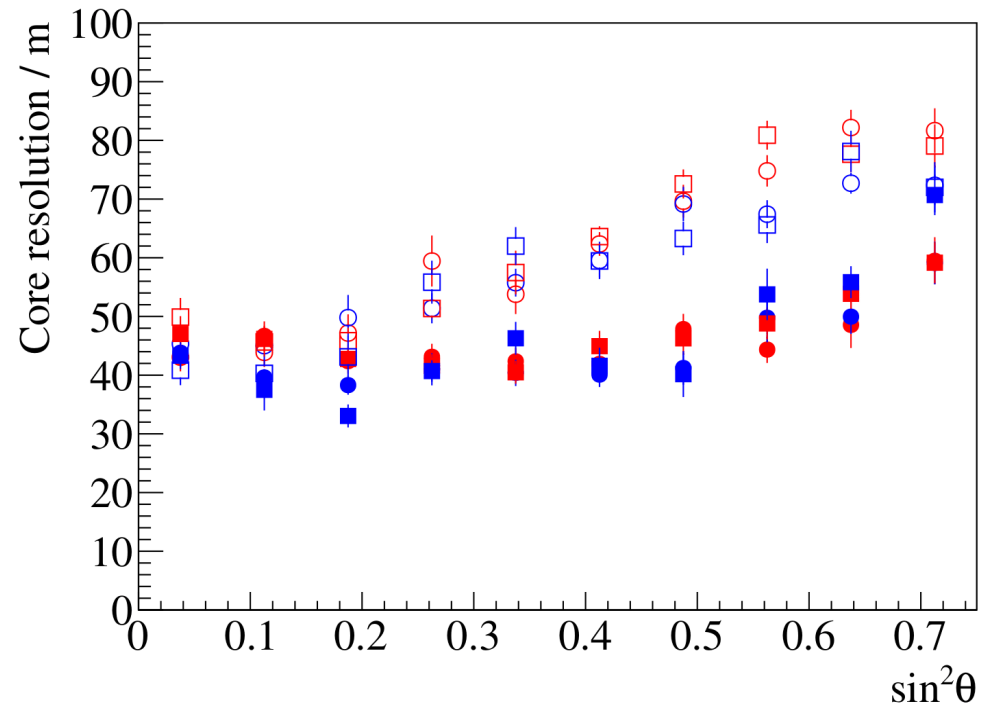
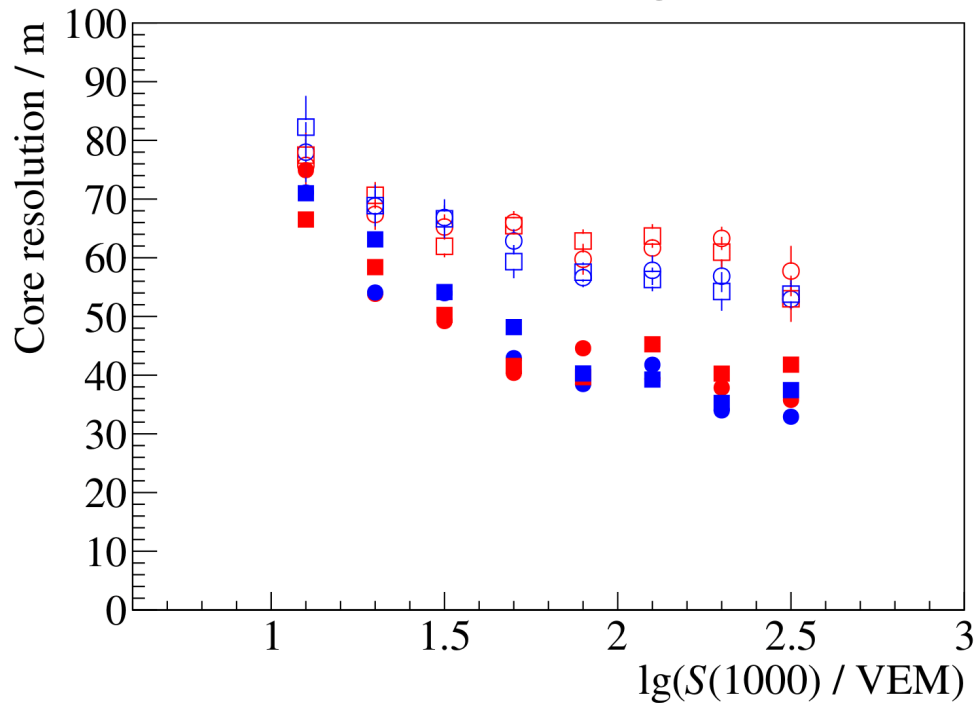
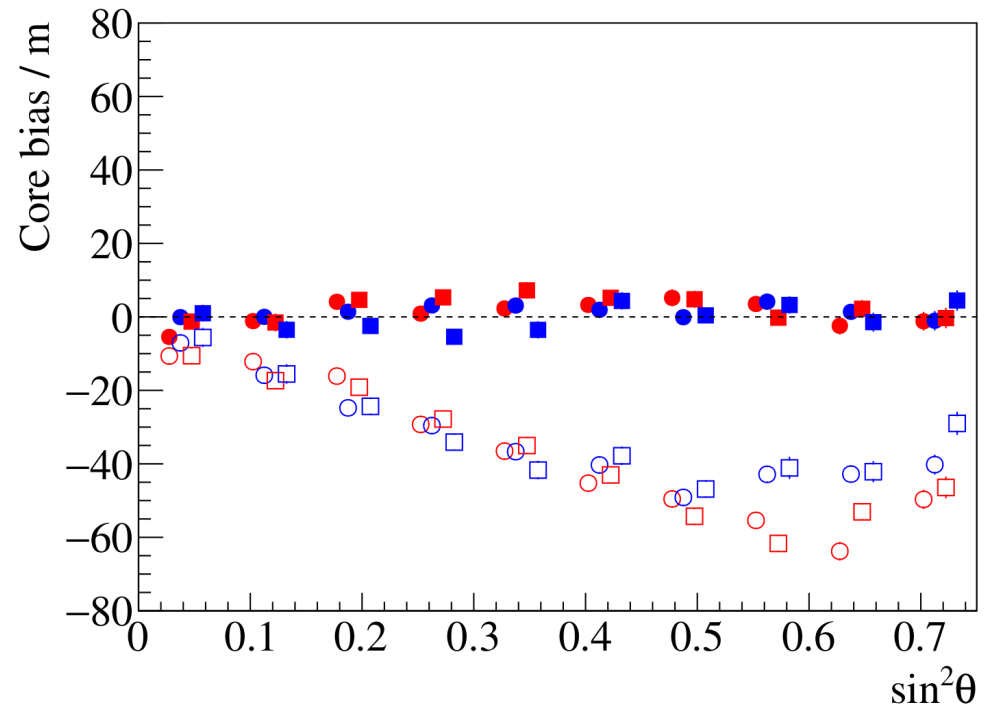
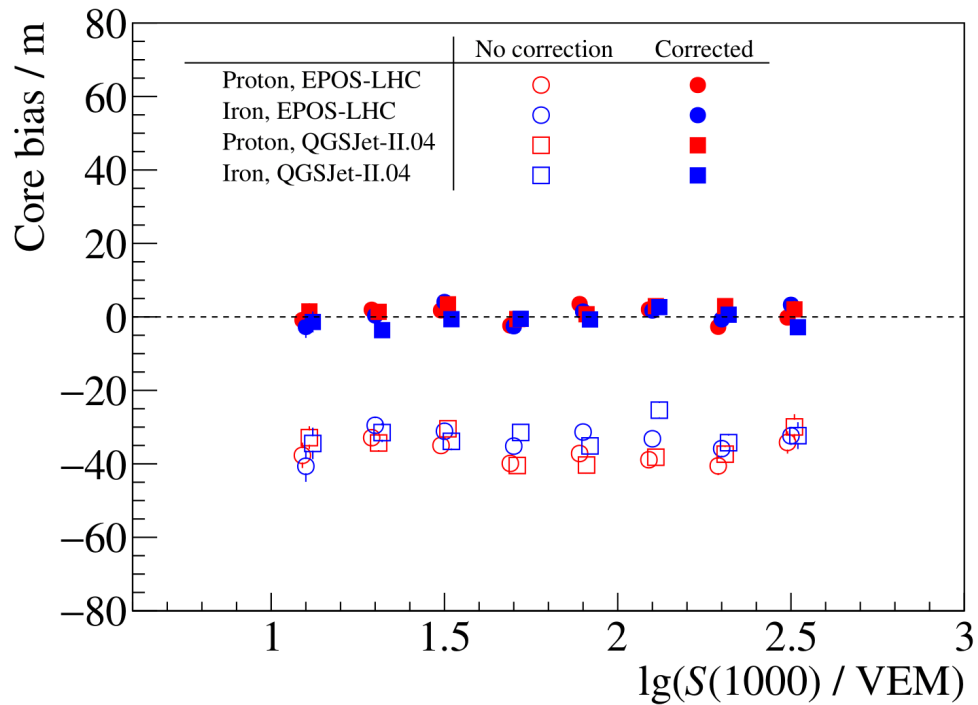
No bias in the perpendicular direction

## Resolution of the core position:



Core resolution = distance at which **68.3% of the cumulative distribution of the distance** between the simulated and reconstructed positions of the core, is reached

# Impact on the core reconstruction



# Impact on zenith and $S(1000)$

## Uncertainties in $S(1000)$ :

Comparison of:

$S(1000)_{MC}$   $S(1000)$  computed from the ring of 24 detectors at 1000 m

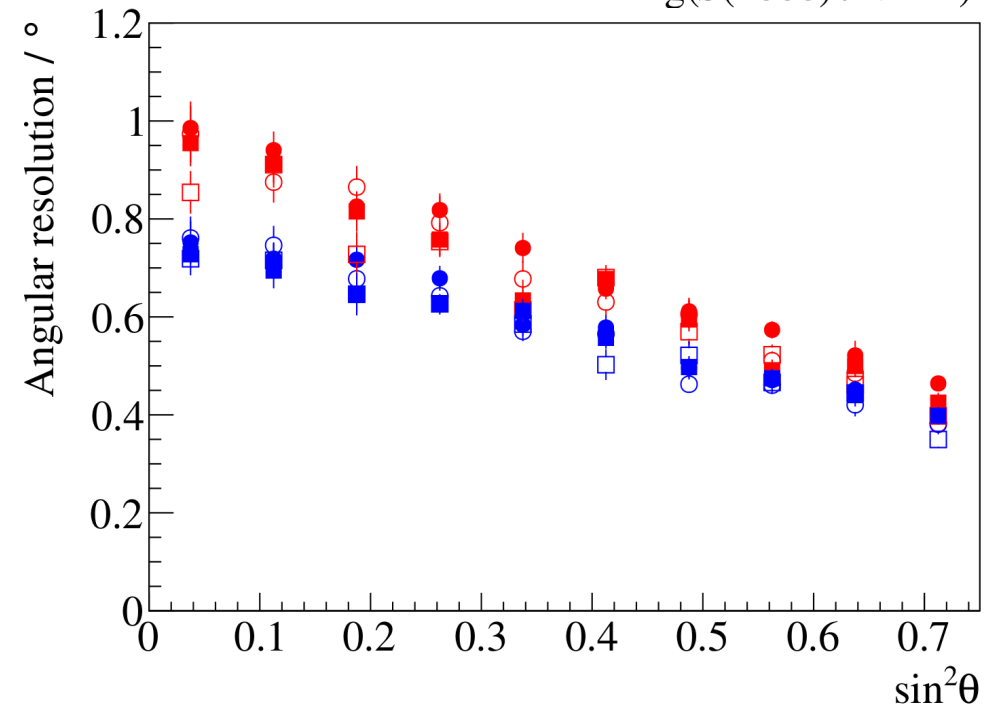
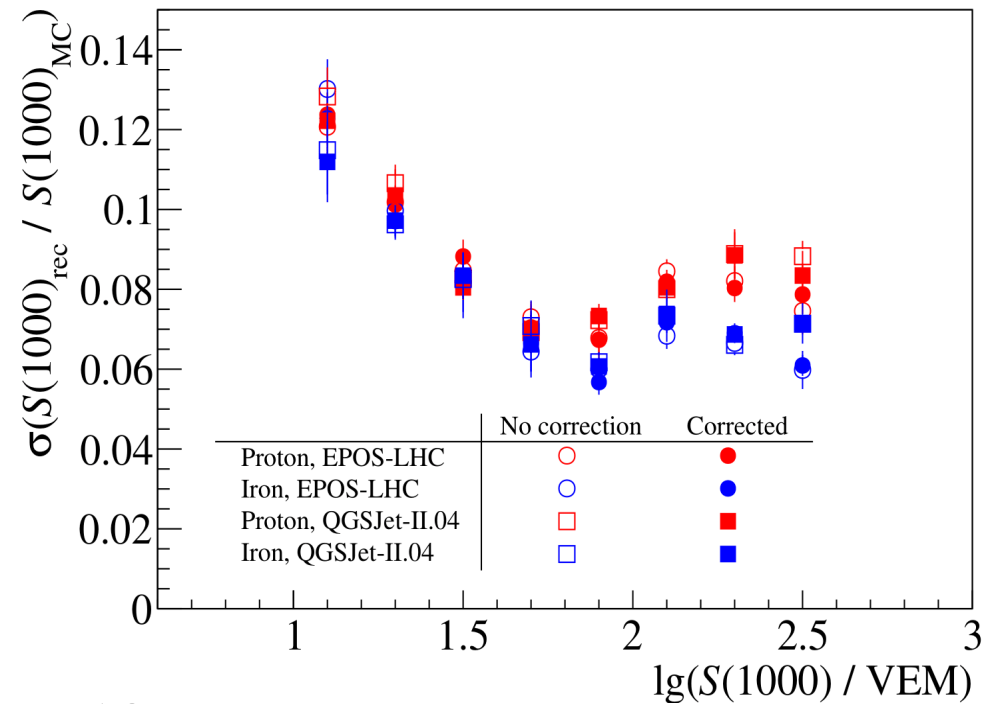
$S(1000)_{rec}$  reconstructed  $S(1000)$

## Angular resolution:

Opening-angle:

$$\sin \eta = |\hat{a}_{MC} \times \hat{a}_{rec}|$$

Angular resolution = angle at which **68.3%** of the cumulative distribution of  $\eta$  is reached



# To conclude...

---

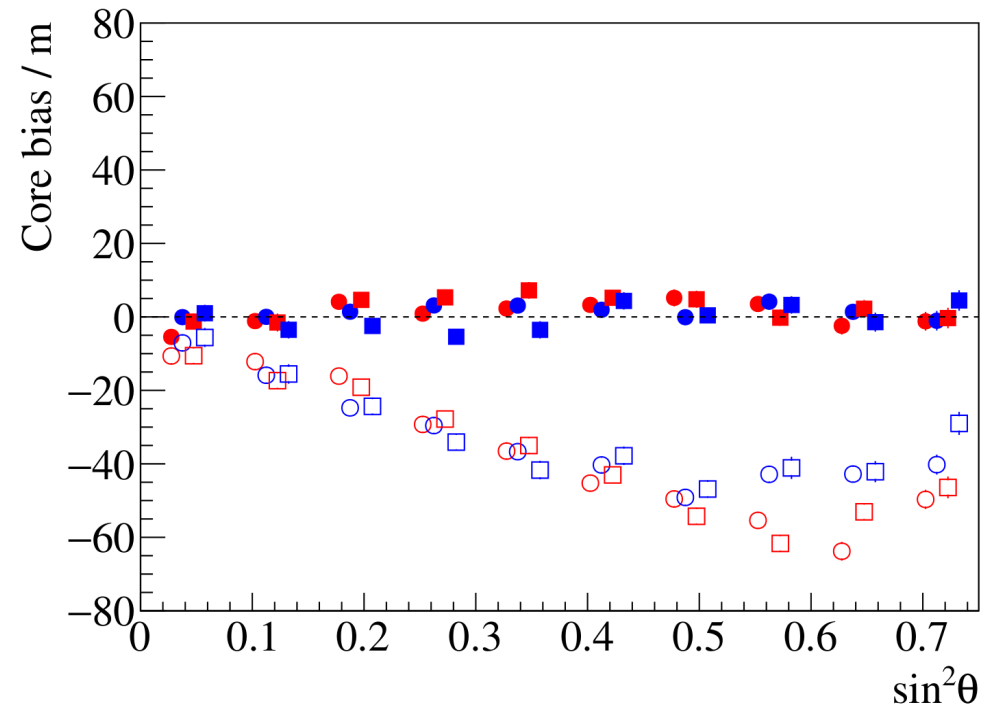
Azimuthal asymmetry in signals **observed in the water-Cherenkov detector** of the surface detector are combination of:

- **geometrical effects** from the inclination of the shower
- **attenuation of the charged particles** from their point of emission to the ground

From simulations, **development of a model describing the amplitude of the asymmetry:**

- suppression of the bias
- improvement of the resolution

No impact observed on the uncertainties in  $S(1000)$  or on the angular resolution



# To conclude...

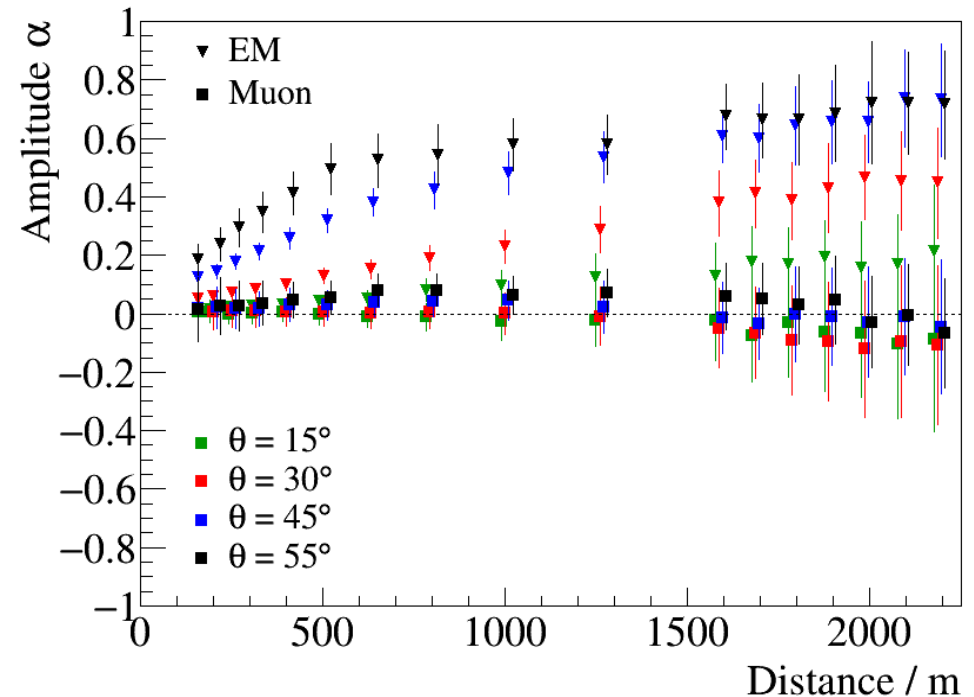
---

Azimuthal asymmetry in signals **observed in the water-Cherenkov detector** of the surface detector are combination of:

- **geometrical effects** from the inclination of the shower
- **attenuation of the charged particles** from their point of emission to the ground

Amplitude of the asymmetry is a **balance between electromagnetic and muonic components** of the shower:

- scintillator detectors?
- muon deficit in simulations?



---

***Trugarez !\****

\* Thank you!

---